

Nuclear Astrophysics Experiments in Cologne

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for the group of Prof. Dr. Andreas Zilges

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**Nuclear Astrophysics in Germany,
Darmstadt, November 16th, 2016**



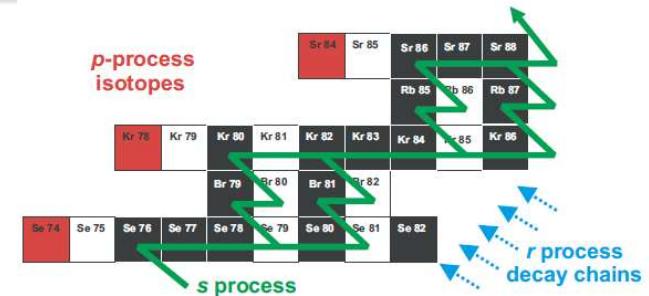
Supported by the ULDETIS project within the UoC Excellence Initiative institutional strategy and by DFG (ZI 510/8-1, INST 216/544-1).

Supported by the Bonn-Cologne Graduate School of Physics and Astronomy.

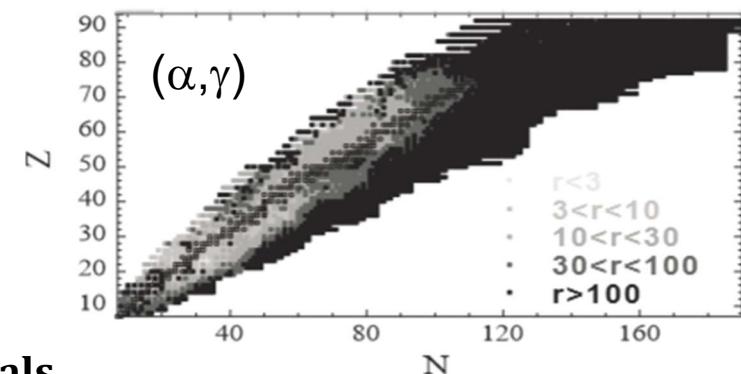
Nuclear Astrophysics Motivation

Nucleosynthesis of p-nuclei

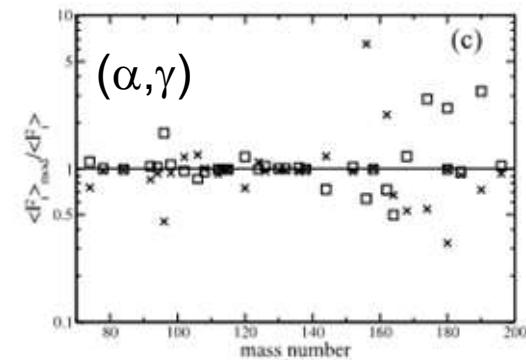
- mostly produced via photodisintegration reactions (γ process)
- measurements of inverse reactions, *e.g.*, $(p,\gamma), (\alpha,\gamma)$
- contributions of other processes?
 - neutrino-driven wind
 - rapid proton-capture process



T. Rauscher *et al.*, Phys. Rep. **76** (2013) 066201.



M. Arnould, S. Goriely, Phys. Rep. **384** (2003) 1.



W. Rapp *et al.*, ApJ **653** (2006) 474.

Sensitivity on reaction rates

- 25 solar mass supernova model
- varying reaction rates by a factor of 3
- improving nuclear physics input

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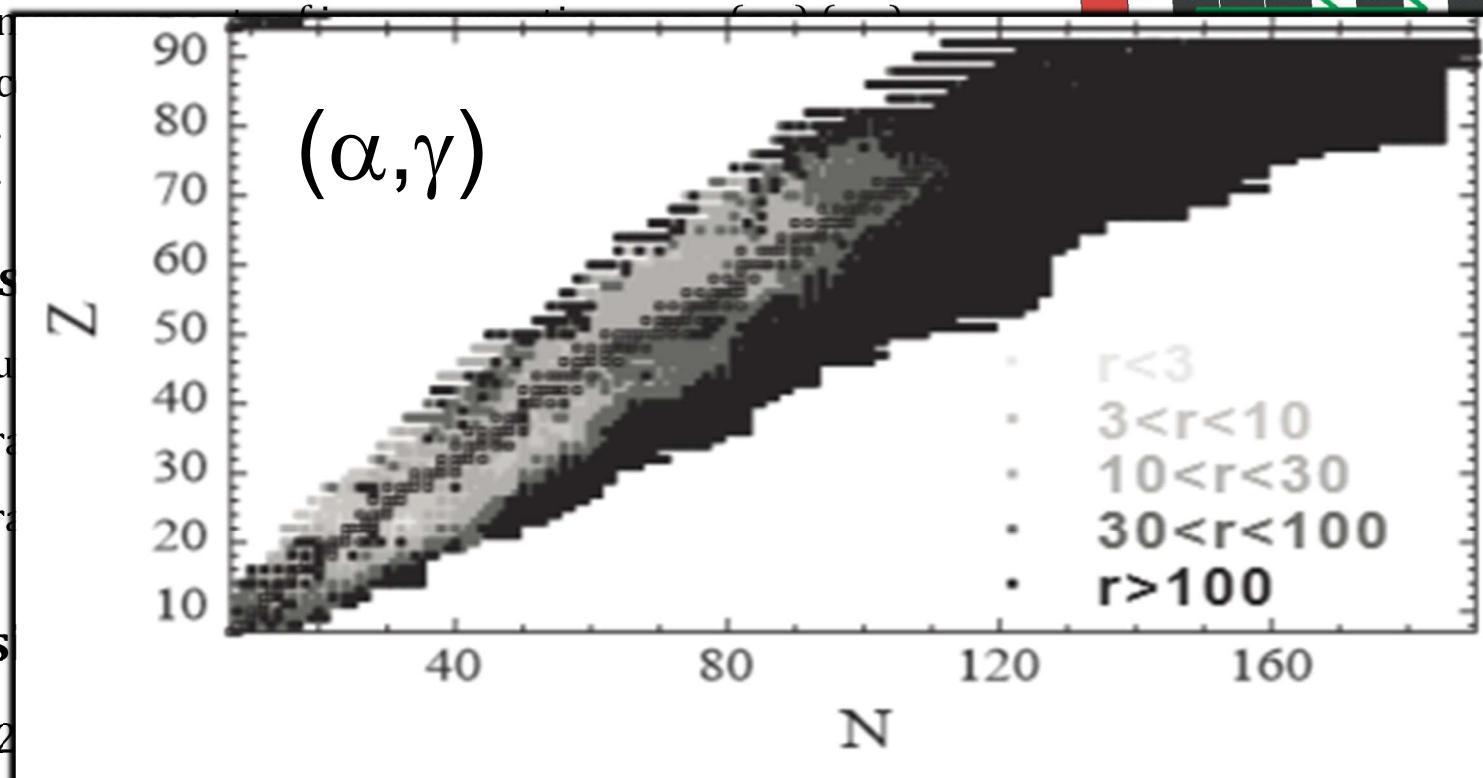
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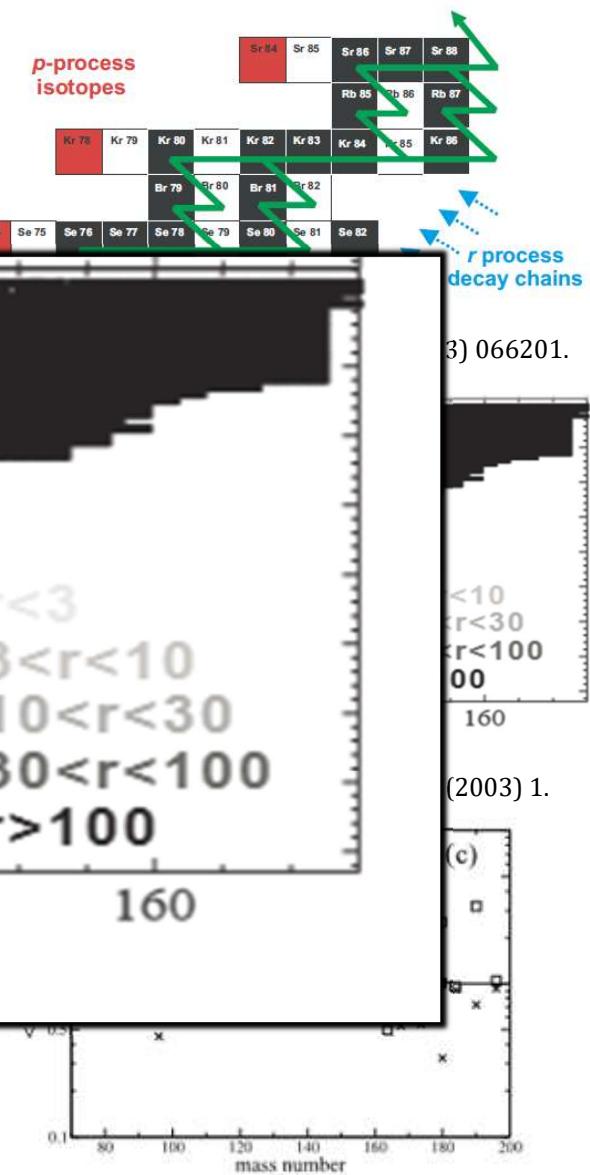
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- varying reaction rates by a factor of 3
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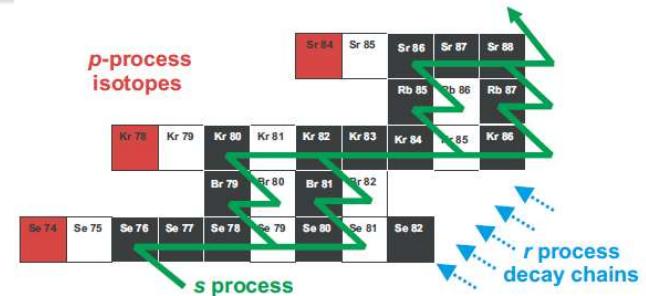


W. Rapp *et al.*, ApJ 653 (2006) 474.

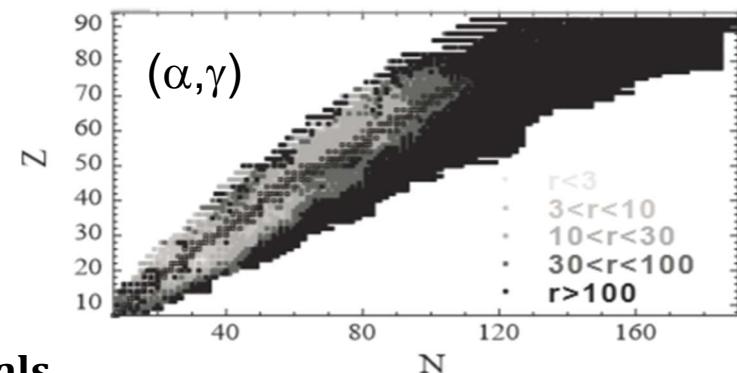
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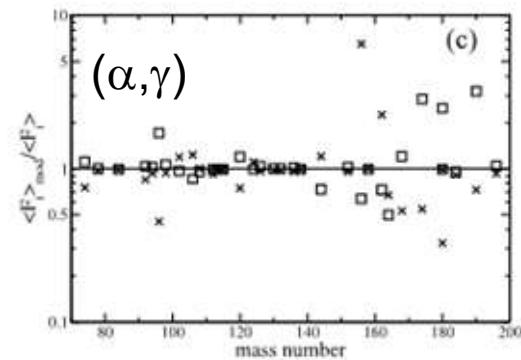
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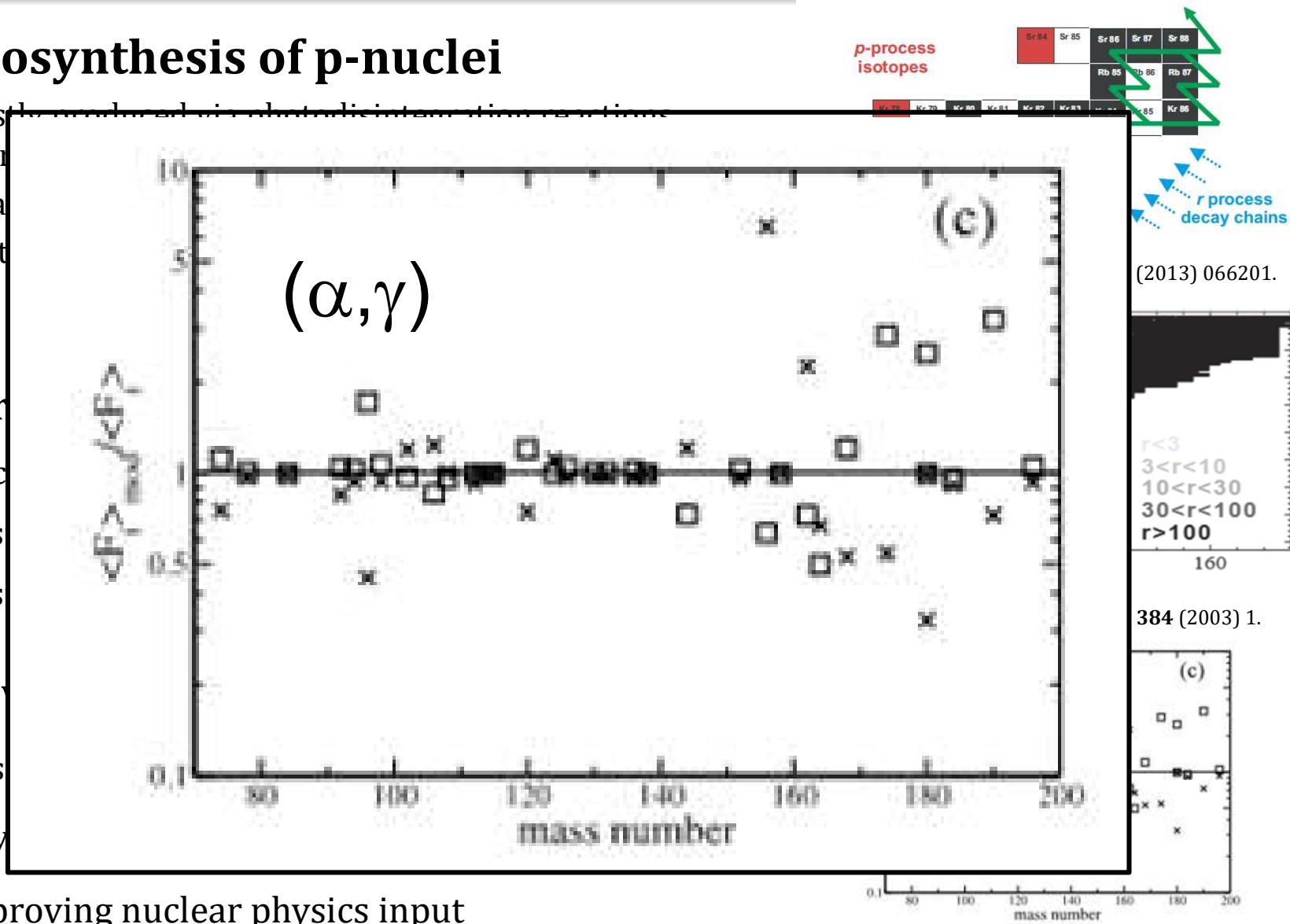
- most abundant elements in the universe
- (γ processes)
- measured
- containing
- ...
- ...

Hauser-Feshbach

- sufficiently
- transient
- transient

Sensitivity

- 25 s
 - vary
- improving nuclear physics input

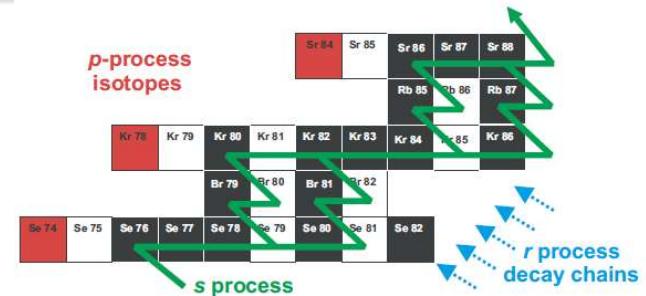


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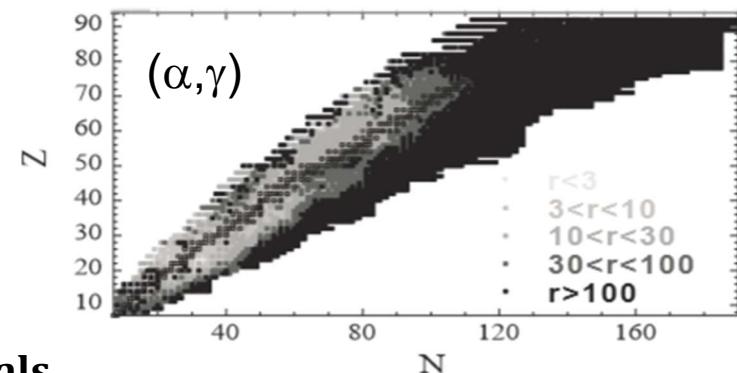
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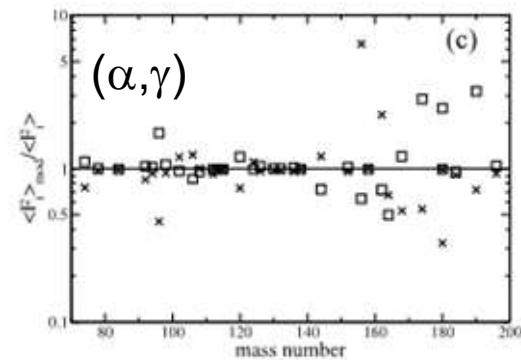
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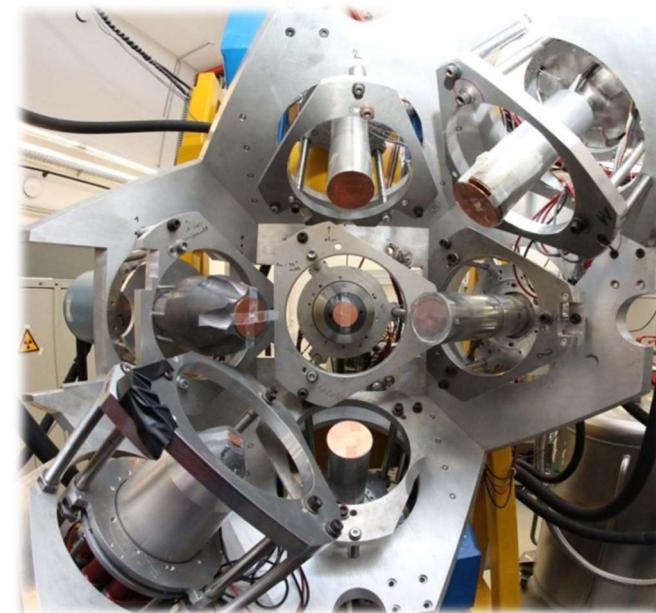
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γ -ray spectrometer HORUS

- 10 MV FN-Tandem ion accelerator

HORUS γ -ray spectrometer

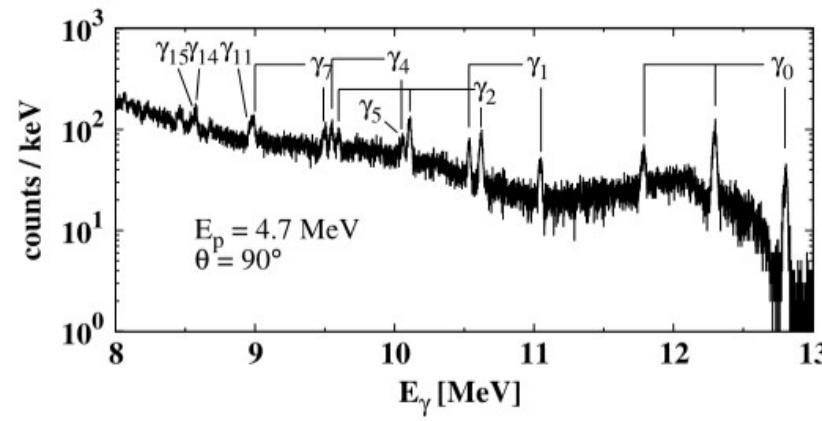
- 14 HPGe detectors
 - High resolution
 $\approx 2 \text{ keV} @ 1332 \text{ keV}$
 - High total efficiency
 $\approx 2\% @ 1332 \text{ keV}$
- 5 different angles with respect to beam axis
 - determination of angular distributions
- BGO shields



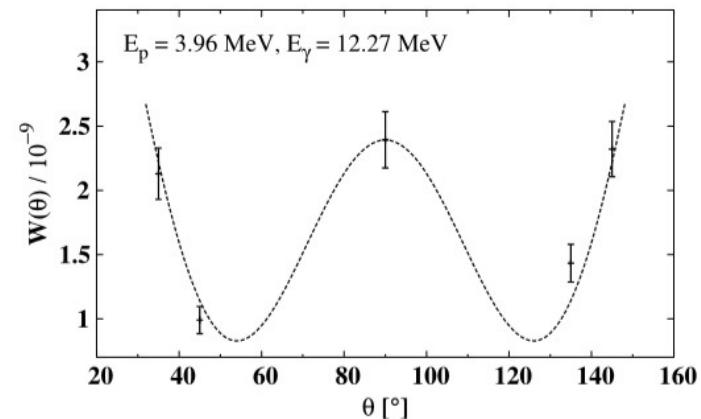
L. Netterdon *et al.*, NIM A 754 (2014) 94

In-beam measurement of cross sections

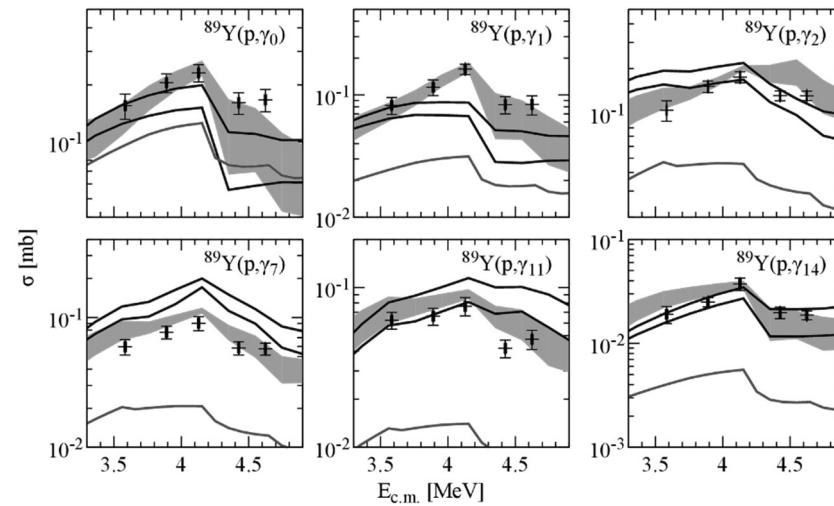
High-resolution γ -ray spectrum



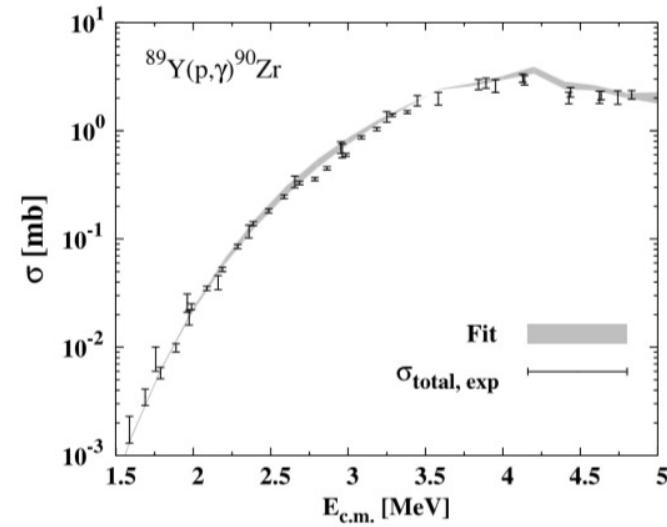
Angular distribution of γ -rays



Partial cross sections

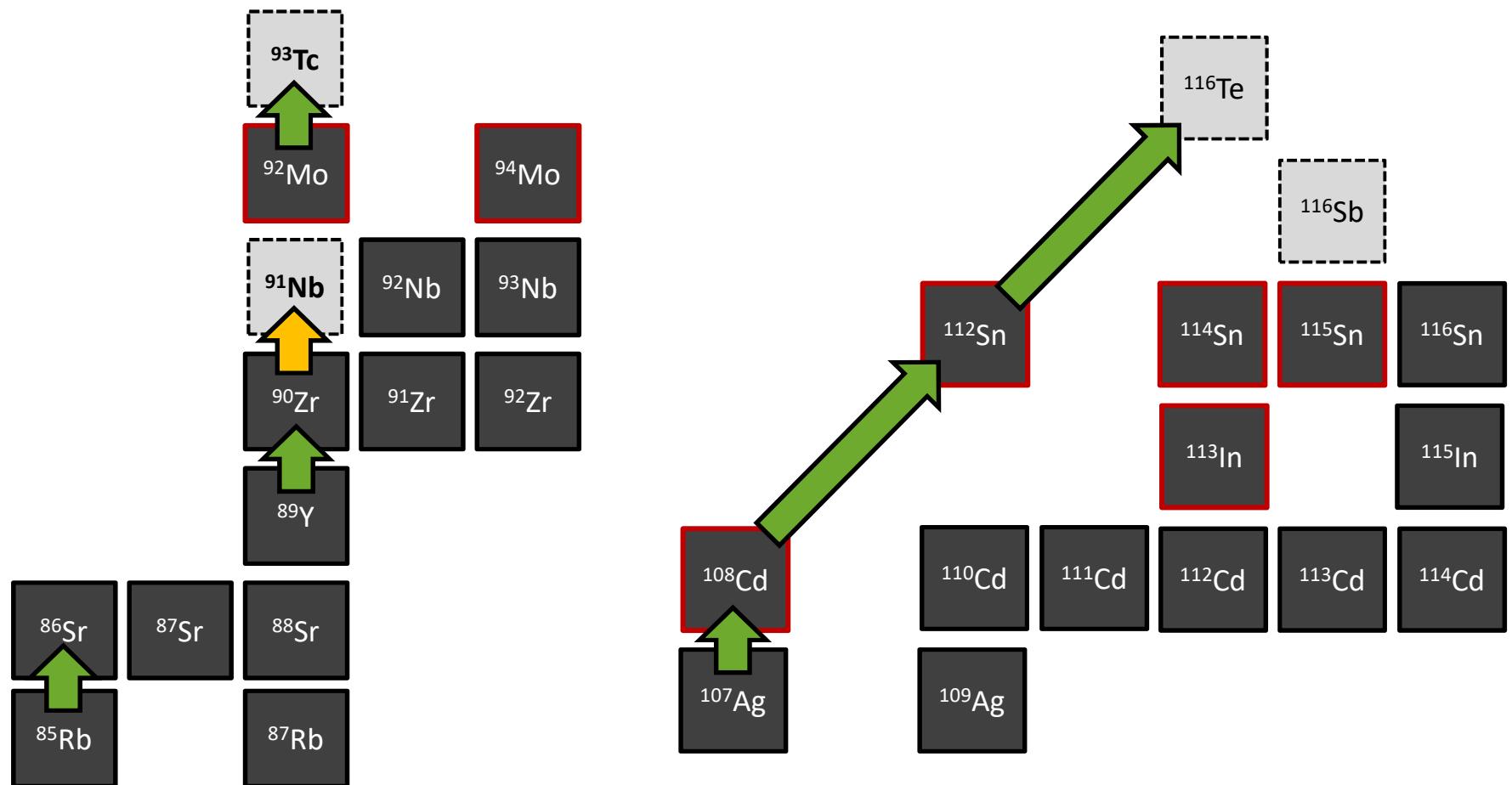


Total cross sections



L. Netterdon *et al.*, PLB 744 (2015) 358

In-beam measurements at HORUS



L. Netterdon *et al.*, PLB **744** (2015) 358

J. Mayer *et al.*, Phys. Rev. C **93** (2016) 045809

L. Netterdon *et al.*, Phys. Rev. C **91** (2015) 035801

P. Scholz *et al.*, PLB **761** (2016) 247

In-beam measurements at HORUS

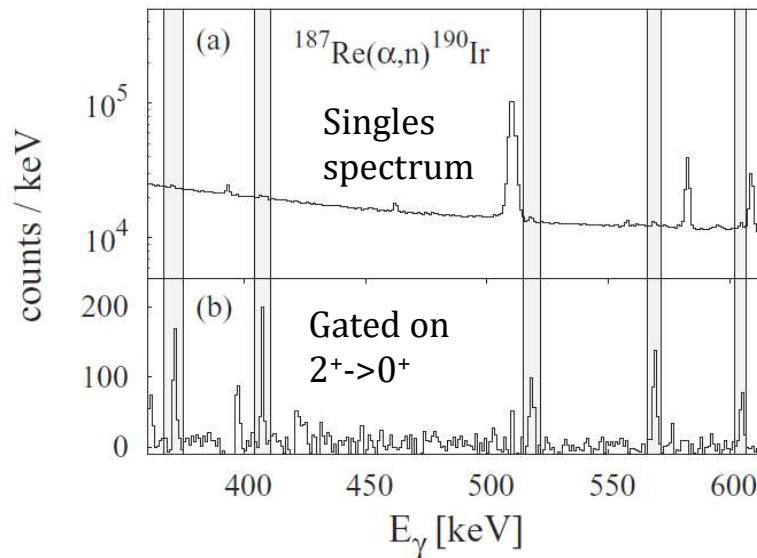
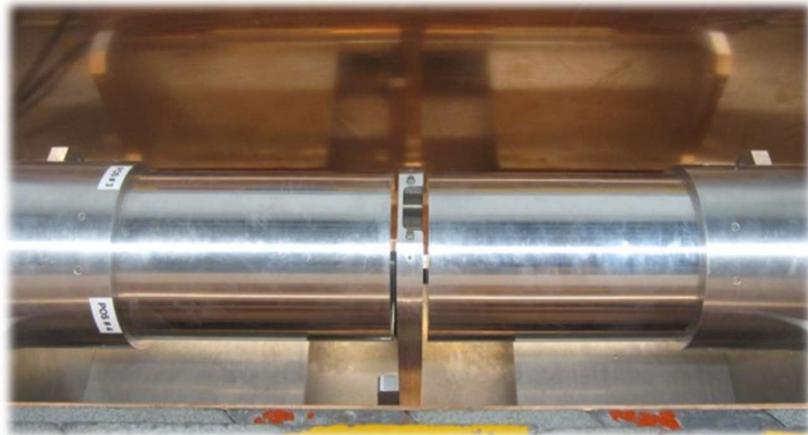
Advantages

- highly resolved γ -decay information
- partial cross sections accessible
- wide choice of reaction products
- selection of reaction channel via angular distributions

Disadvantages

- relatively time consuming measurements
- limitations due to beam current/contaminants
- new α -particle source in 2017/18

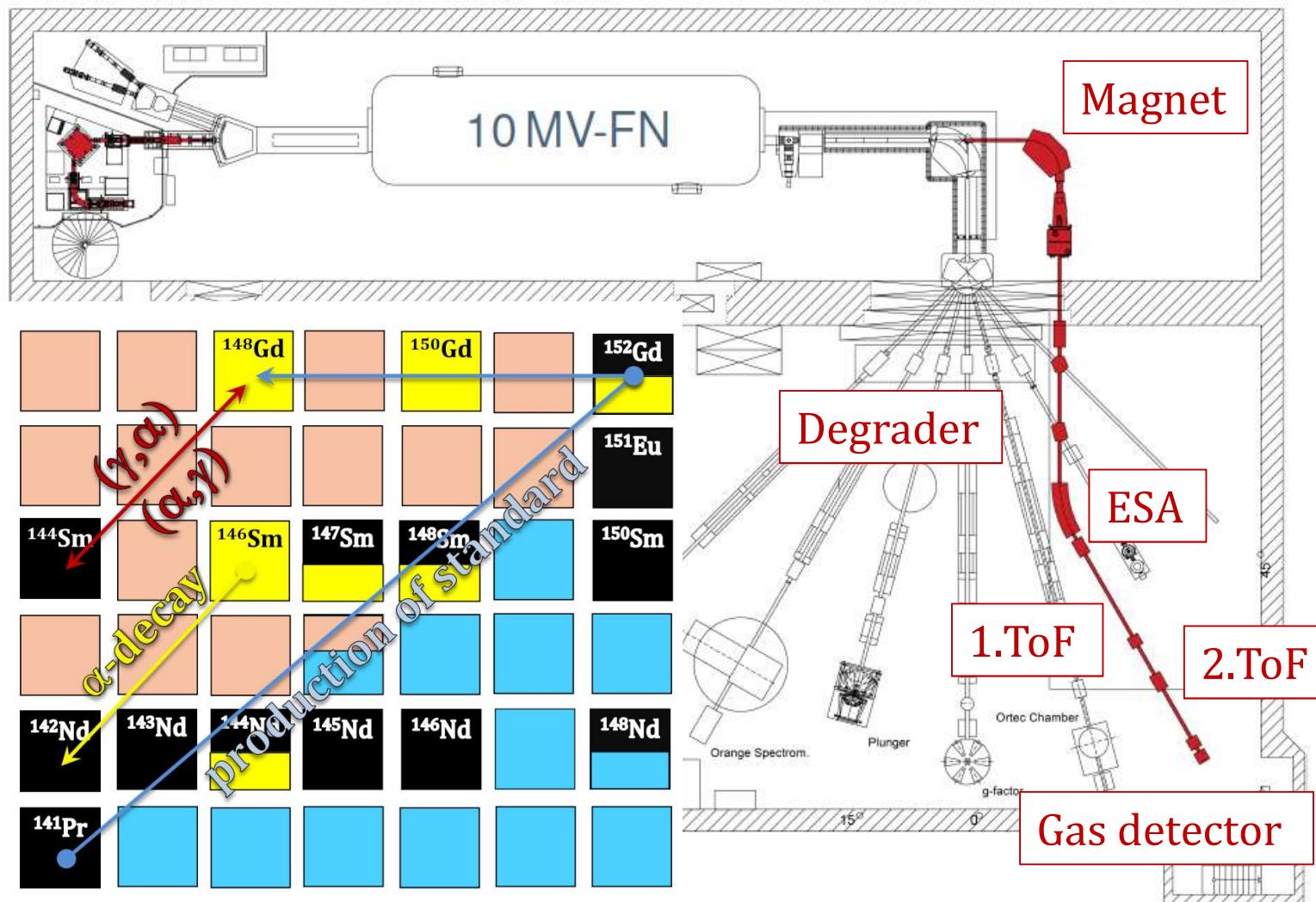
Cologne Clover Counting Setup



P. Scholz *et al.*, Phys. Rev. C **90** (2014) 065807

- complementary measurement of (α ,n) cross sections
- sensitive to the α -OMP
- half-lives down to several minutes possible
- even higher mass α -induced reactions possible
- using $\gamma\gamma$ coincidence method

AMS cross-section measurements



Summary

Nuclear Astrophysics experiments in Cologne

- measuring radiative charged-particle induced cross-sections
- sensitive to particle+nucleus optical-model potentials and γ -strength function
- experimental methods:
 - In-beam technique with HPGe detectors
 - Activation and Accelerator Mass Spectrometry
 - Two Step γ -cascade technique

Systematic studies of nuclear physics input-parameters

- constraining global models means systematic investigation
- need for a larger experimental database
- using complementary methods, *e.g.*, in-beam, activation, 4π summing, etc.
- see own results in the context of other measurements