

Perturbative treatment of the many-body problem in nuclear matter.

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1 Motivations

- The nuclear many-body problem with Skyrme phenomenological forces.
- Which nuclei can be treated (EDF framework)?

2 Green's function formalism

- First-order energy diagrams.
- The mean-field approximation: Good for bulk properties of nuclei!
- Second-order energy diagrams.

3 Applications in nuclear matter

- A simple Dirac-delta force.
- Inclusion of 2-body density-dependent force (2BDDF).

4 Full Skyrme interaction

- Dimensional regularization with MS.
- Cutoff dependence.

5 Conclusions and Perspectives

- The essential goal of quantum many-body physics is to study the nature and the effects of interactions between particles as well as the observable properties of many-particle systems.
- The Schrödinger equation that describes the dynamics of a many-body system composed by A nucleons is given by:

$$H\Psi = (T + V)\Psi = E\Psi,$$

where the Hamiltonian is written as the sum of a kinetic term T and an interaction term V , that represents in principle a 2-body, 3-body, \dots , up to a A -body force,

$$H = T + V.$$

- In other words, the Schrödinger equation may be written as:

$$H\Psi = \left\{ \sum_{i=1}^A -\frac{\hbar^2}{2m} \nabla_i^2 + 2\text{-body} + 3\text{-body} + \dots + \sum_{i_1 < \dots < i_A} v(i_1, \dots, i_A) \right\} = E\Psi,$$

where i represents all coordinates of the i^{th} nucleon.

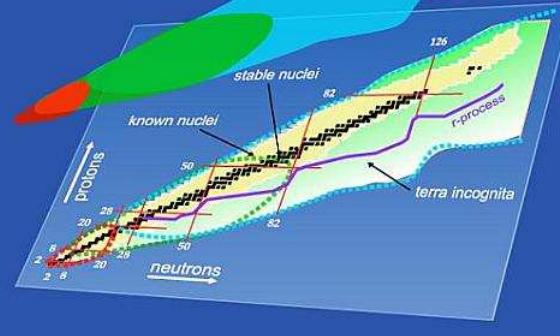
- From the phenomenological point of view, it turns out that, in most cases, the interaction is well enough described by the 2-body (and possibly the 3-body) terms, and therefore the Hamiltonian reduces to:

$$H \sim \sum_{i=1}^A -\frac{\hbar^2}{2m} \nabla_i^2 + \frac{1}{2} \sum_{i \neq j=1} V(i,j).$$

- A simple example of a two-body phenomenological interaction in nuclear physics is the Skyrme effective force (zero-range force) which is given by:

$$\begin{aligned} V_{12}(\vec{r}, \vec{R}) = & t_0(1 + x_0 P^\sigma) \delta(\vec{r}) + \frac{1}{2} t_1(1 + x_1 P^\sigma) \left[\delta(\vec{r}) \vec{k}^2 + \vec{k}'^2 \delta(\vec{r}) \right] \\ & + t_2(1 + x_2 P^\sigma) \vec{k}' \cdot \delta(\vec{r}) \vec{k} + \frac{1}{6} t_3(1 + x_3 P^\sigma) \delta(\vec{r}) \rho^\alpha(\vec{R}) \\ & + iW_0 \vec{\sigma} \cdot \vec{k}' \times \delta(\vec{r}) \vec{k}. \end{aligned}$$

Nuclear Landscape



- Which nuclei can be treated? From medium-mass to heavy nuclei.
- Blue region: Energy Density Functional (EDF) framework (for example Skyrme forces): Mean-field framework.
 - Mean field for ground-state nuclear structure (HF, HFB,...)
 - RPA and QRPA for small-amplitude oscillations
 - Beyond small amplitude oscillations: time-dependent mean field for dynamics (TDHF, TDHFB,...).

- The Green's function, also known as the Feynman propagator is defined as:

$$G(r, t; r', t') = -i \langle \Phi_0 | T [\Psi(r, t) \Psi^\dagger(r', t')] | \Phi_0 \rangle, \quad \text{by assuming } \langle \Phi_0 | \Phi_0 \rangle = 1.$$

- When $t' > t$, the function $G(r, t; r', t')$ creates a particle at time t and position r , then destroys it again at time t' and position r' , in other words, it measures the probability of a particle propagating from (r, t) to (r', t') .
- The formula due to Gell-Mann and Low expresses the shift energy of the ground state with respect to the unperturbed system as:

$$E - \epsilon_k^{(0)} = \sum_{m=0}^{\infty} (-i)^m \frac{1}{m!} \int_{-\infty}^0 dt_1 \cdots dt_m \langle \Phi_0 | T [\hat{H}_1 \hat{H}_1(t_1) \cdots \hat{H}_1(t_m)] | \Phi_0 \rangle_{\text{connected}},$$

where the interacting Hamiltonian \hat{H}_1 can be written as:

$$H_1 = \frac{1}{2!} \sum_{\substack{\alpha\beta\alpha' \beta' \\ \gamma\mu\gamma' \mu'}} \int d^3\mathbf{r} d^3\mathbf{r}' \hat{\psi}_{\alpha;\gamma}^\dagger(\mathbf{r}) \hat{\psi}_{\beta;\mu}^\dagger(\mathbf{r}') V(\mathbf{r}, \mathbf{r}')_{\alpha\alpha', \beta\beta'} \hat{\psi}_{\beta';\mu'}(\mathbf{r}') \hat{\psi}_{\alpha';\gamma'}(\mathbf{r}).$$

- The first-order contributions to the total energy is given by:

$$E^{(1)} = \langle \Phi_0 | T [\hat{H}_1] | \Phi_0 \rangle_{connected}.$$

- The time-ordering can be expressed as:

$$T [\hat{\psi}_{\alpha;\gamma}^\dagger(\mathbf{r}) \hat{\psi}_{\beta;\mu}^\dagger(\mathbf{r}') \hat{\psi}_{\beta';\mu'}(\mathbf{r}') \hat{\psi}_{\alpha';\gamma'}(\mathbf{r})] = \hat{\psi}_{\alpha;\gamma}^\dagger(\mathbf{r}) :: \hat{\psi}_{\alpha';\gamma'}(\mathbf{r}) \hat{\psi}_{\beta;\mu}^\dagger(\mathbf{r}') :: \hat{\psi}_{\beta';\mu'}(\mathbf{r}') \\ - \hat{\psi}_{\alpha;\gamma}^\dagger(\mathbf{r}) :: \hat{\psi}_{\beta';\mu'}(\mathbf{r}') \hat{\psi}_{\beta;\mu}^\dagger(\mathbf{r}') :: \hat{\psi}_{\alpha';\gamma'}(\mathbf{r}).$$

- By noting that: $\hat{\psi}_{a;b}^\dagger(\mathbf{r}) \hat{\psi}_{a';b'}^\dagger(\mathbf{r}') = iG_{ab;a'b'}^0(\mathbf{r}, \mathbf{r}')$, we get:

$$T [\hat{\psi}_{\alpha;\gamma}^\dagger(\mathbf{r}) \hat{\psi}_{\beta;\mu}^\dagger(\mathbf{r}') \hat{\psi}_{\beta';\mu'}(\mathbf{r}') \hat{\psi}_{\alpha';\gamma'}(\mathbf{r})] = iG_{\alpha'\alpha;\gamma'\gamma}^0(\mathbf{r}, \mathbf{r}') iG_{\beta'\beta;\mu'\mu}^0(\mathbf{r}', \mathbf{r}') \\ - iG_{\beta'\alpha;\mu'\gamma}^0(\mathbf{r}', \mathbf{r}) iG_{\alpha'\beta;\gamma'\mu}^0(\mathbf{r}, \mathbf{r}').$$

- Therefore, we are able to write the first-order energy as a sum of a direct and exchange part:

$$E^{(1)} = E_{direct}^{(1)} + E_{exchange}^{(1)}.$$

- The expression for the first-order correction (direct term) is given by:

$$E_{direct}^{(1)} = \frac{1}{2!} \sum_{\substack{\alpha\beta\alpha'\beta' \\ \gamma\mu\gamma'\mu'}} \int d^4r d^4r' U(r, r')_{\alpha\alpha', \beta\beta'} \left[iG_{\alpha'\alpha; \gamma'\gamma}^0(\mathbf{r}, \mathbf{r}) iG_{\beta'\beta; \mu'\mu}^0(\mathbf{r}', \mathbf{r}') \right].$$

- The expression for the first-order correction (exchange term) is given by:

$$E_{exchange}^{(1)} = \frac{1}{2!} \sum_{\substack{\alpha\beta\alpha'\beta' \\ \gamma\mu\gamma'\mu'}} \int d^4r d^4r' U(r, r')_{\alpha\alpha', \beta\beta'} \left[-iG_{\beta'\alpha; \mu'\gamma}^0(\mathbf{r}', \mathbf{r}) iG_{\alpha'\beta; \gamma'\mu}^0(\mathbf{r}, \mathbf{r}') \right].$$

- Mean-field models are described by first-order contributions.



Figure : Left (Right): First-order energy correction direct (exchange) term.

Advantages of mean-field approaches.

In nuclear physics, mean-field approaches lead to satisfactory results when applied to bulk properties of atomic nuclei.

- Masses, radii or ground state deformations.

Mean-field approaches are not always very accurate.

- For example, in nuclear physics, mean-field approaches do not predict accurately the single-particle spectra.
- Spectroscopic factors and the fragmentation of the single-particle energies cannot be reproduced in a precise way.
- This why we want to formulate the nuclear many-body problem in a beyond mean-field framework.

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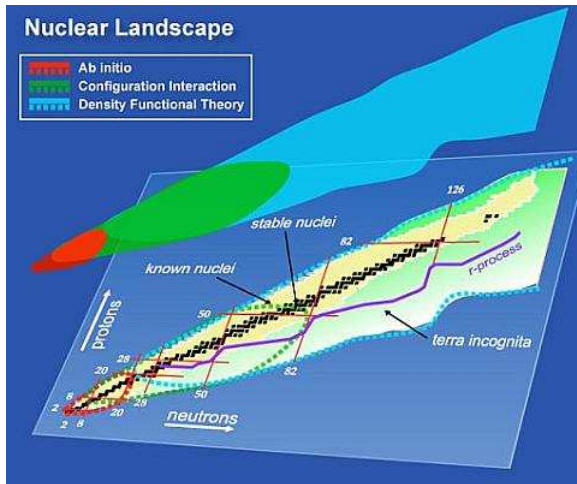
Self-consistent mean-field models for nuclear structure

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- Which nuclei can be treated? From medium-mass to heavy nuclei.
- Blue region: Energy Density Functional (EDF) framework (for example Skyrme forces): Mean-field framework.
- Beyond-mean-field models.
 - Particle-Vibration Coupling (PVC)
 - Multiparticle-Multihole Configuration Mixing.
 - 2nd order energy correction.
 - Second RPA,.....etc
- What are the problems?

- The double counting problem:
 - 1 The first question that one should address before adding correlations is: What are the many-body correlations that are effectively included implicitly in the mean-field approach?
 - 2 When one uses phenomenological interactions that are adjusted at the mean-field level, the problem of the double counting of correlations (that are implicitly contained to some extent in the parameters) should be addressed.
- The problem of ultraviolet divergences:
 - 1 UV divergences may appear for instance when one uses models beyond the mean-field level with zero-range interactions.
 - 2 For example, UV divergences appears in the so-called Bogoliubov-de Gennes or Hartree-Fock-Bogoliubov (HFB) theories if a zero-range interaction is employed in the pairing channel to treat a superfluid many-fermion system.

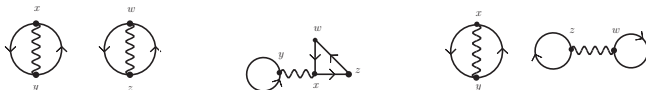
- The second-order energy correction is given by:

$$E^{(2)} = -i \int_{-\infty}^0 dt_1 \langle \Phi_0 | T [\hat{H}_1 \hat{H}_1(t_1)] | \Phi_0 \rangle_{connected}.$$

- The expectation value of all the terms containing normal-ordered products of operators vanishes in the non-interacting ground state $|\Phi_0\rangle$, leaving only the fully contracted products of field operators. We have:

$$T [\hat{\psi}_a^\dagger(x) \hat{\psi}_b^\dagger(y) \hat{\psi}_d(y) \hat{\psi}_c(x) \hat{\psi}_\alpha^\dagger(z) \hat{\psi}_\beta^\dagger(w) \hat{\psi}_\mu(w) \hat{\psi}_\gamma(z)] = 24 \text{ terms.}$$

- Examples of second-order disconnected and anomalous diagrams are:



- Finally, according to Goldstone's theorem, the only contributions coming from the second-order terms are:

- The expression of the second-order correction due to the direct contribution is given by:

$$E_d^{(2)} = \left(\frac{g^2}{2}\right) \frac{1}{(2\pi)^9} \int_{C_I} d^3 p_1 d^3 p_2 d^3 q \frac{V^2(q)}{\epsilon_{p_1}^{(0)} + \epsilon_{p_2}^{(0)} - \epsilon_{p_1+q}^{(0)} - \epsilon_{p_2-q}^{(0)}}.$$

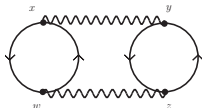


Figure : Direct diagram.

- Similarly, the expression of the second-order correction due to the exchange contribution is given by:

$$E_{exch}^{(2)} = - \left(\frac{g}{2}\right) \frac{1}{(2\pi)^9} \int_{C_I} d^3 p_1 d^3 p_2 d^3 q \frac{V(q)V(p_1 - p_2 - q)}{\epsilon_{p_1}^{(0)} + \epsilon_{p_2}^{(0)} - \epsilon_{p_1+q}^{(0)} - \epsilon_{p_2-q}^{(0)}}.$$

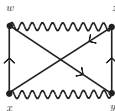


Figure : Exchange contribution.

- As a first attempt, we use a simple delta force which is spin-independent and we deal with nuclear matter, where all the calculations may be done analytically:

$$V(r_1, r_2) = t_0 \delta(r_1 - r_2).$$

- We calculate the equation of state for symmetric nuclear matter at the second-order in perturbation theory.

$$\frac{E}{A}(\rho) = \frac{E^{(0)}}{A}(\rho) + \frac{E^{(1)}}{A}(\rho) + \frac{E^{(2)}}{A}(\rho, \infty).$$

- Due to the short range character of the interaction, this equation of state diverges. We use the momentum cutoff Λ procedure to regularize the divergent integrals:

$$\frac{E}{A}(\rho, \Lambda) = \frac{3}{10m} k_F^2 + \frac{t_0}{4\pi^2} k_F^3 - \left(\frac{9m}{8\pi^4} \right) t_0^2 k_F^4 I(k_F, \Lambda).$$

- $I(k_F, \Lambda)$ is an analytic function of ρ and Λ and it diverges linearly for large values of the cutoff Λ :

$$I(k_F, \Lambda \gg M) = \frac{\Lambda}{9k_F} + \frac{-11 + 2 \log 2}{105} + \mathcal{O}\left(\frac{k_F}{\Lambda}\right).$$

- The asymptotic behaviour of the equation of state for large values of Λ is given by:

$$\frac{E}{A}(\rho, \Lambda \gg M) = \frac{3k_F^2}{10m} + \frac{t_0}{4\pi^2} k_F^3 - \left(\frac{9m}{8\pi^4}\right) t_0^2 k_F^4 \left[\frac{-11 + 2 \log 2}{105} + \frac{\Lambda}{9k_F} + \mathcal{O}\left(\frac{k_F}{\Lambda}\right) \right].$$

- The terms that depend on the cutoff can be regrouped with the terms coming from the mean-field contribution:

$$\frac{E}{A}(\rho, \Lambda \gg M) = \frac{3k_F^2}{10m} + \frac{1}{4\pi^2} k_F^3 \left[t_0 - \frac{mt_0^2}{2\pi^2} \Lambda \right] + \left(\frac{9m}{8\pi^4}\right) \left(\frac{11 - 2 \log 2}{105}\right) t_0^2 k_F^4.$$

- The dependence on the cut-off will be eliminated by defining renormalized parameters t_0^R from the bare parameter $t_0(\Lambda)$:

$$t_0^R = t_0(\Lambda) + C_2 \Lambda t_0^2(\Lambda) = t_0(\Lambda) [1 + C_2 \Lambda t_0(\Lambda)] \text{ such that: } \frac{d}{d\Lambda} t_0^R = 0.$$

- It is to be noted that:

$$t_0(\Lambda) = t_0^R - C_2 \Lambda t_0^2(\Lambda) = t_0^R - C_2 \Lambda \left(t_0^R \right)^2 + \mathcal{O} \left(t_0^R \right)^3 ;$$

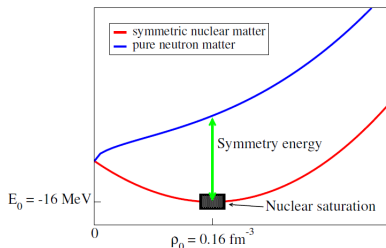
$$t_0^2(\Lambda) = \left[t_0^R - C_2 \Lambda \left(t_0^R \right)^2 + \mathcal{O} \left(t_0^R \right)^3 \right]^2 = \left(t_0^R \right)^2 + \mathcal{O} \left(t_0^R \right)^3 .$$

- Therefore, the equation of state for symmetric matter evaluated at the second-order can be written as:

$$\frac{E}{A}(\rho, \Lambda \gg M) = \frac{3k_F^2}{10m} + \frac{t_0^R}{4\pi^2} k_F^3 + \left(\frac{9m}{8\pi^4} \right) \left(\frac{11 - 2 \log 2}{105} \right) \left(t_0^R \right)^2 k_F^4 .$$

- We conclude that the problem is renormalized in this case by redefining the parameter t_0 .
- However, this simple model does not provide any saturation point for symmetric matter, even at the mean-field level.

- Calculations of the properties in nuclear matter show that the binding energy is $E_0 \approx -16$ MeV, and $\rho_0 \approx 0.16 \text{ fm}^{-3}$.



- Three-body forces are considered as an indispensable ingredient in accurate calculations not only of few-nucleon systems and the structure of light nuclei but also for many-body systems in many cases.

- We have already concluded that in non-relativistic approaches the model of nucleons interacting only via a two-body force fails to reproduce the empirical saturation observables. Thus phenomenological three-body forces have been introduced with few adjustable parameters.
- For example, Skyrme introduced a zero-range force in order to achieve saturation in nuclear matter:

$$V_{123}(r_1, r_2, r_3) = t_3 \delta(r_1 - r_2) \delta(r_2 - r_3).$$

- Later, Vautherin and Brink replaced the contact three-body force by a contact density-dependent two-body force.

$$t_3 \delta(r_1 - r_2) \delta(r_2 - r_3) \longrightarrow \frac{t_3}{6} \rho^\alpha \delta(r_1 - r_2), \quad \text{where } \alpha = 1.$$

- The next step is to add a zero-range density-dependent force $\frac{1}{6}t_3\rho^\alpha\delta(r)$ to the t_0 interaction.

$$V(r_1, r_2) = \left(t_0 + \frac{t_3}{6}\rho^\alpha \right) \delta(r_1 - r_2).$$

- It has to be noted that it corresponds to a three-body force at the mean-field level when $\alpha = 1$. The mean-field contribution to the density-dependent term is equal to:

$$e/A = \left(\frac{t_3}{36\pi^4} \right) k_F^6$$



Figure : A contact three-body force at first-order.

- The density-dependent effective two-body interaction at second order is given by:

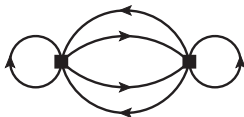


Figure : Density-dependent effective two-body interaction to second order.

- The equation of state for symmetric nuclear matter evaluated at the second-order is given by:

$$\frac{E}{A}(\rho, \Lambda) = \frac{3k_F^2}{10m} + \frac{t_0}{4\pi^2} k_F^3 + \left(\frac{t_3}{36\pi^4}\right) k_F^6 - \left(\frac{9m}{8\pi^4}\right) \left(t_0 + \frac{t_3}{6}\rho^\alpha\right)^2 k_F^4 I(k_F, \Lambda).$$

Where the expression $I(k_F, \Lambda)$ is an analytic function of ρ and Λ and it diverges linearly for large cutoff Λ :

$$I(k_F, \Lambda \gg M) = \frac{-11 + 2 \log 2}{105} + \frac{\Lambda}{9k_F} + \mathcal{O}\left(\frac{k_F}{\Lambda}\right).$$

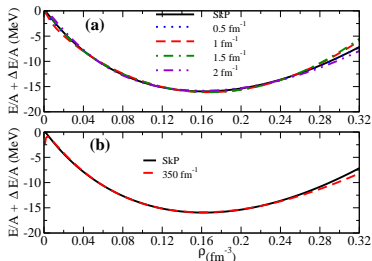
- Then, the EoS for symmetric nuclear matter becomes (when $\alpha = 1$):

$$\begin{aligned} \frac{E}{A}(\rho, \Lambda \gg M) &= \frac{3k_F^2}{10m} + \frac{1}{4\pi^2} \left[t_0 - \frac{m\Lambda}{2\pi^2} t_0^2 \right] k_F^3 + \frac{1}{36\pi^4} \left[t_3 - \frac{m\Lambda}{\pi^2} t_0 t_3 \right] k_F^6 \\ &+ \left(\frac{9m}{8\pi^4}\right) \left(t_0 + \frac{t_3}{6}\rho\right)^2 \left[\frac{11 - 2 \ln 2}{105}\right] k_F^4 - \left(\frac{m\Lambda}{648\pi^8} t_3^2\right) k_F^9. \end{aligned}$$

- The last term can not be regrouped into the existing parameters unless we add a four-body term treated perturbatively at the mean-field level.

- Since the technique of adding counter-terms will complicate our problem, we will follow a phenomenological approach by which the divergence is absorbed by adjusting the parameters of the Skyrme interaction.
- In the case of effective interactions between point-nucleons, the cutoff Λ must certainly be smaller than the momentum associated with the nucleon size, i.e., $\Lambda \leq 2 \text{ fm}^{-1}$.
- Moreover, the energy scale of low-energy nuclear phenomena in finite nuclei is much lower. See, for instance:
[J. Dobaczewski, K. Bennaceur, F. Raimondi, J. Phys. G39 \(2012\) 125103.](#)
- In fact, to describe giant resonances or rotational bands of nuclei, the scale should be even smaller, perhaps around 0.5 fm^{-1} .
- For each value of Λ , we can perform a least-square fit to determine a new parameter set SkP_Λ , such that the EoS including the second-order correction matches the one obtained with the original force SkP at the mean-field level.

- The SkP-equation of state is chosen as our reference on which we perform the fit.



K. Moghrabi, et al., Phys. Rev. Lett. 105, 262501 (2010).

- The quality of the fits demonstrates that the refitted interactions can describe satisfactorily the empirical equation of state for the case of symmetric nuclear matter treated with a simplified contact force.

Table : From the second line, columns 2, 3 and 4: parameter sets obtained in the fits associated with different values of the cutoff Λ compared with the original set SkP (first line). In the fifth column the χ^2/N -value (χ^2 divided by the number of fitted points) associated to each fit is shown. In columns 6 and 7 the saturation point is shown.

	t_0 (MeV fm ³)	t_3 (MeV fm ^{3+3α)}	α	χ^2/N	ρ_0 (fm ⁻³)	$E/A(\rho_0)$ (MeV)
SkP	-2931.70	18708.97	1/6		0.16	-15.95
$\Lambda = 0.5$ fm ⁻¹	-2352.900	15379.861	0.217	0.00004	0.16	-15.96
$\Lambda = 1$ fm ⁻¹	-1155.580	9435.246	0.572	0.00142	0.17	-16.11
$\Lambda = 1.5$ fm ⁻¹	-754.131	8278.251	1.011	0.00106	0.17	-16.09
$\Lambda = 2$ fm ⁻¹	-632.653	5324.848	0.886	0.00192	0.16	-15.82
$\Lambda = 350$ fm ⁻¹	-64.904	360.039	0.425	0.00042	0.16	-15.88

- For any value of the cutoff Λ , it is possible to find a new refitted interaction that can be used to take into account the mean-field contribution plus the second-order corrections.

- Let us turn our attention to the full Skyrme interaction by including the velocity-dependent terms:

$$\begin{aligned}
 V_{12}(\vec{r}, \vec{R}) = & t_0(1 + x_0 P^\sigma) \delta(\vec{r}) + \frac{1}{6} t_3 (1 + x_3 P^\sigma) \delta(\vec{r}) \rho^\alpha(\vec{R}) \\
 & + \frac{1}{2} t_1 (1 + x_1 P^\sigma) \left[\delta(\vec{r}) \vec{k}^2 + \vec{k}'^2 \delta(\vec{r}) \right] + t_2 (1 + x_2 P^\sigma) \vec{k}' \cdot \delta(\vec{r}) \vec{k} \\
 & + i W_0 \vec{\sigma} \cdot \vec{k}' \times \delta(\vec{r}) \vec{k}.
 \end{aligned}$$

- We calculate analytically the equation of state at second order in both symmetric and asymmetric nuclear matter with different neutron-to-proton ratios δ .
- The equation of state at the mean-field level for asymmetric nuclear matter is given by:

$$\begin{aligned}
 \frac{E}{A}(\delta, \rho) = & \frac{3}{10m} G_{5/3} k_F^2 + \frac{t_0}{12\pi^2} [2(2 + x_0) - (1 + 2x_0) G_2] k_F^3 \\
 & + \frac{t_3}{48} \left(\frac{2}{3\pi^2} \right)^{1+\alpha} [2(2 + x_3) - (1 + 2x_3) G_2] k_F^{3+3\alpha} \\
 & + \frac{1}{40\pi^2} \left[\Theta_v G_{5/3} + \frac{1}{2} (\Theta_s - 2\Theta_v) G_{8/3} \right] k_F^5,
 \end{aligned}$$

- where $G_\beta = \frac{1}{2} [(1 + \delta)^\beta + (1 - \delta)^\beta]$, $\Theta_s = 3t_1 + t_2(5 + 4x_2)$, and $\Theta_v = t_1(2 + x_1) + t_2(2 + x_2)$.

- The asymptotic behavior of the second-order energy contribution ($\Lambda \gg M$) is given by:

$$\begin{aligned} \frac{\Delta E^{(2)}}{A}(\delta, \rho) = & k_F^3 [a_0(\delta)\Lambda + a_1(\delta)\Lambda^3 + a_2(\delta)\Lambda^5] + k_F^5 [b_0(\delta)\Lambda + b_1(\delta)\Lambda^3] \\ & + k_F^{3+3\alpha} [c_1(\delta)\Lambda + c_2(\delta)\Lambda^3] \\ & + k_F^7 [c_0(\delta)] + k_F^{3+6\alpha} [c_3(\delta)\Lambda] + k_F^{5+3\alpha} [c_4(\delta)]. \end{aligned}$$

- What counter terms should be added?

- The first three terms can be regrouped with the mean-field contributions by redefining some existing parameters.
- The term $k_F^7 [c_0(\delta)]$ can be absorbed by adding to the original Skyrme interaction the term: $\nabla^4 \delta(r_1 - r_2)$.
- The term: $k_F^{3+6\alpha} [c_3(\delta)\Lambda] = k_F^9 [c_3(\delta)\Lambda]$ ($\alpha = 1$), corresponds to a four-body force: $\delta(r_1 - r_2)\delta(r_2 - r_3)\delta(r_3 - r_4)$;
- Term $k_F^{5+3\alpha} [c_4(\delta)] = k_F^8 [c_4(\delta)]$ corresponds to $\nabla^2 \delta(r_1 - r_2)\delta(r_2 - r_3)$.

- Possible directions:

- Dimensional regularization (DR) with minimal subtraction (MS) scheme.
- Absorbing the divergence through a fitting procedure.

- The first + second-order EoS for symmetric matter is given by

$$\frac{E}{A}(k_F) = \frac{3}{10m} k_F^2 + \frac{t_0}{4\pi^2} k_F^3 + \frac{t_3}{16} \left(\frac{2}{3\pi^2} \right)^{1+\alpha} k_F^{3+3\alpha} + \frac{1}{40\pi^2} \Theta_s k_F^5 + \frac{\Delta E^S}{A}(k_F)$$

- The second-order correction is given by:

$$\begin{aligned} \frac{\Delta E^S}{A}(k_F) = & k_F^4 \left(\frac{-11 + 2 \ln 2}{105} \right) \chi_1^S(k_F) + k_F^6 \left(\frac{-167 + 24 \ln 2}{945} \right) \chi_2^S(k_F) \\ & + k_F^8 \left(\frac{-2066 + 312 \ln 2}{31185} \right) \chi_3^S(k_F) + k_F^8 \left(\frac{-9997 + 1236 \ln 2}{62370} \right) \chi_4^S(k_F). \end{aligned}$$

where all the χ terms depend on the parameters of Skyrme interaction and on the Fermi momentum k_F .

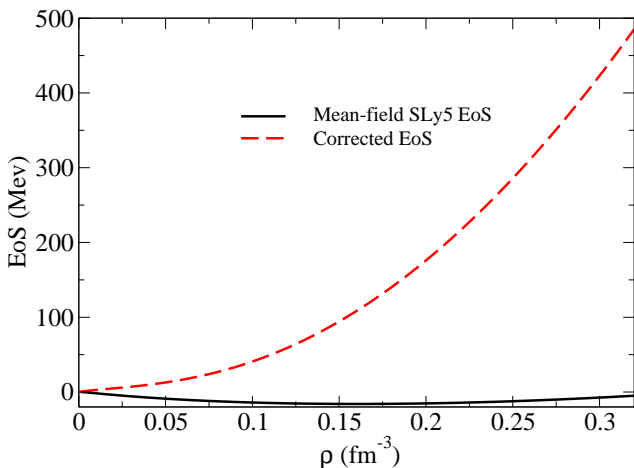


Figure : SLy5 mean-field EoS (full line) and mean-field + second-order EoS calculated with the SLy5 parameters (dashed line) for symmetric matter.

- The beyond mean-field EoS evaluated at second order for pure neutron matter is given by:

$$\begin{aligned} \frac{E}{A}(k_n) &= \frac{3}{10m} k_n^2 + \frac{1}{12\pi^2} t_0 (1 - x_0) k_n^3 + \frac{1}{24} \left(\frac{1}{3\pi^2} \right)^{1+\alpha} (1 - x_3) t_3 k_n^{3+3\alpha} \\ &+ \frac{1}{40\pi^2} (\Theta_s - \Theta_v) k_n^5 + \frac{\Delta E^N}{A}(k_n), \end{aligned}$$

- This time the second-order correction is given by:

$$\begin{aligned} \frac{\Delta E^N}{A}(k_n) &= k_n^4 \left(\frac{-11 + 2 \ln 2}{105} \right) \chi_1^N(k_n) + k_n^6 \left(\frac{-167 + 24 \ln 2}{2835} \right) \chi_2^N(k_n) \\ &+ k_n^6 \left(\frac{167 - 24 \ln 2}{5670} \right) \chi_3^N(k_n) + k_n^8 \left(\frac{461 - 24 \ln 2}{31185} \right) \chi_4^N(k_n) \\ &+ k_n^8 \left(\frac{-4021 + 516 \ln 2}{124740} \right) \chi_5^N(k_n) \end{aligned}$$

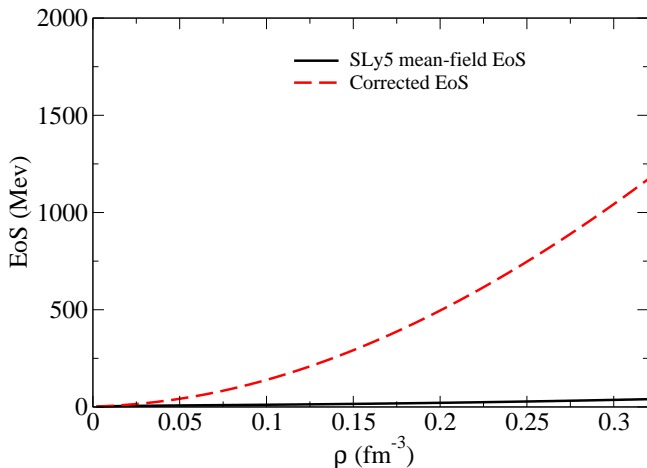
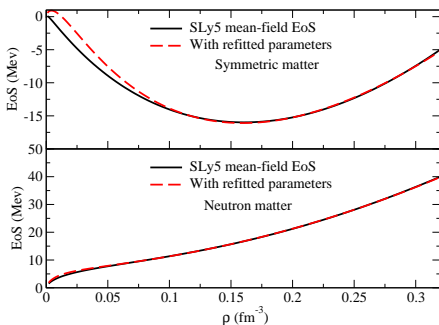


Figure : SLy5 mean-field EoS (full line) and mean-field + second-order EoS calculated with the SLy5 parameters (dashed line) for neutron matter.

- A problem of double counting! Need to refit the parameters.
- The following χ^2 is minimized:

$$\chi^2 = \frac{1}{N-1} \sum_{i=1}^N \frac{(E_i - E_{i,ref})^2}{\Delta E_i^2}.$$

- 1 The number N of fitted points is 15 and the points are in the range of densities between 0.1 to 0.30 fm⁻³.
 - 2 The errors ΔE_i are chosen equal to 1% of the reference SLy5 mean-field energies $E_{i,ref}$.
- We perform a global fit by considering together symmetric and pure neutron matter. In this case: $\chi^2 = \frac{1}{2} (\chi_{\delta=0}^2 + \chi_{\delta=1}^2)$.



K.M et al., Phys. Rev. C 86, 044319 (2012).

Table : Parameter sets obtained in the fit of the EoS of symmetric and pure neutron matter compared with the original set SLy5. In the last column the χ^2 value is shown.

	t_0 (MeV fm ³)	t_1 (MeV fm ⁵)	t_2 (MeV fm ⁵)	t_3 (MeV fm ^{3+3α})	x_0	x_1	x_2	x_3	α	χ^2
SLy5	-2484.88	483.13	-549.40	13763.0	0.778	-0.328	-1.0	1.267	0.16667	-
New	-460.73	10403.66	-8485.73	-141558.6	1.460	-0.681	-0.641	-0.779	0.650	0.202

- It has to be noted that this set of parameters is cutoff-independent which is a good result.
- However, the values of the new parameter are far from their original values (SLy5) which may lead to problems of non-convergence at the Hartree-Fock level, when applied to finite nuclei.
- Let us go back to the asymptotic behavior of the second-order energy correction::

$$\frac{\Delta E^{(2)}}{A}(\delta, \rho) = k_F^3 [a_0(\delta)\Lambda + a_1(\delta)\Lambda^3 + a_2(\delta)\Lambda^5] + k_F^5 [b_0(\delta)\Lambda + b_1(\delta)\Lambda^3] \\ + k_F^{3+3\alpha} [c_1(\delta)\Lambda + c_2(\delta)\Lambda^3] \\ + k_F^7 [c_0(\delta)] + k_F^{3+6\alpha} [c_3(\delta)\Lambda] + k_F^{5+3\alpha} [c_4(\delta)].$$

- Adding counter terms to the original Skyrme interaction to suppress the terms in red would strongly complicate the calculations, especially in the perspective of doing applications to finite nuclei.
- Therefore, we employ a simple procedure and absorb the UV divergences by adjusting the parameters of Skyrme in order to have a reasonable second-order EoS (as we have already done for the $t_0 - t_3$ model).

- We are dealing with phenomenological interactions where their corresponding parameters can be adjusted according to which diagrams are explicitly introduced.
- To have a reasonable second-order EoS, we have adjusted the nine parameters of the Skyrme interaction entering in the expression of the EoS to reproduce the reference SLy5 mean-field EoS.
- We have chosen 15 equidistant reference points (N) for densities ranging from 0.02 fm^{-3} to 0.30 fm^{-3} . All the parameters are kept free in the adjustment procedure.
- The minimization has been performed using the following definition for the χ^2 ,

$$\chi^2 = \frac{1}{N-1} \sum_{i=1}^N \frac{(E_i - E_{i,ref})^2}{\Delta E_i^2}.$$

- The errors or *adopted* standard deviations, ΔE_i are chosen equal to 1% of the reference SLy5 mean-field energies $E_{i,ref}$.

- K.M et al., Phys. Rev. C 85, 044323 (2012).

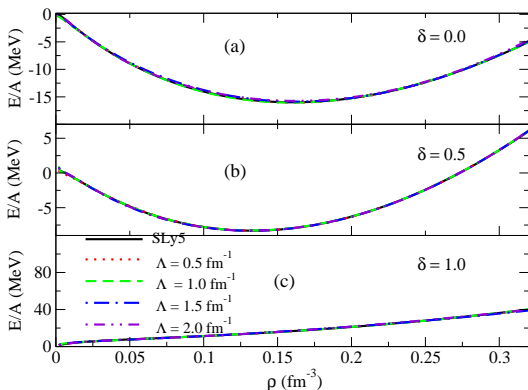


Figure : Refitted EoS (global fit) for symmetric (a), asymmetric (b) and pure neutron (c) matter. The reference SLy5 mean-field curves are also plotted in the 3 panels (solid lines).

Table : Parameter sets obtained in the fit of the EoS of symmetric, asymmetric and pure neutron matter for different values of the cutoff Λ compared with the original set SLy5. In the last column the χ^2 values are shown.

	t_0	t_1	t_2	t_3	x_0	x_1	x_2	x_3	α	
	(MeV fm ³)	(MeV fm ⁵)	(MeV fm ⁵)	(MeV fm ^{3+3α)}						
SLy5	-2484.88	483.13	-549.40	13736.0	0.778	-0.328	-1.0	1.267	0.16667	
$\Lambda(\text{fm}^{-1})$										χ^2
0.5	-2022.142	290.312	1499.483	12334.459	0.481	-5.390	-1.304	0.880	0.259	0.411
1.0	-627.078	83.786	-971.384	186.775	3.428	-1.252	-1.620	200.360	0.338	0.540
1.5	-743.227	112.246	-42.816	5269.849	1.013	3.478	-2.114	0.189	0.814	1.733
2.0	-718.397	573.884	-497.766	6179.243	0.391	-0.393	-0.574	0.785	1.051	1.313

Conclusions/Nuclear Matter

- Analytical study of the divergent terms at the second-order in the equation of state of nuclear matter with Skyrme interaction.
- We were not able to remove the divergences by their corresponding adding counter terms because the problem will become complicated.
 - ① In the simple case of $t_0 - t_3$ model, the divergence was absorbed through a fitting procedure.
 - ② In the case of the full interaction: the techniques of DR/MS were introduced. In this case, a unique set of parameters is obtained for the adjusted effective interaction.
 - ③ The UV divergences were **absorbed** so that the reference EOS (SLy5) of both symmetric and nuclear matter with different neutron-to-proton ratio are reproduced in the case of MC.

Perspectives/ Finite Nuclei

- These encouraging results open new perspectives for future applications of this kind of interactions to treat finite nuclei in beyond mean-field models.
- Introduce the coupling between nucleon individual degrees of freedom and collective degrees of freedom by means of the particle-vibration coupling approach (pvc).

- Motivation: Let us see what would happen if we replace the density-dependent term by the three-body term.
- At the first order, they give the same energy contribution. What about the second-order? Can we renormalize our problem?
- The 1st order diagrams generated by a 3–body contact interaction is given by:



Figure : 1st order diagrams generated by a 3–body contact interaction.

- The energy contribution is given by:

$$e/A = \left(\frac{t_3}{36\pi^4} \right) k_F^6.$$

- The 2nd order diagrams generated by a 3-body contact interaction is given by:

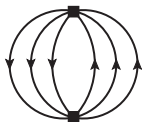


Figure : 2nd order diagrams generated by a 3-body contact interaction.

- The second-order energy correction due to the three-body force is given by:

$$E = \frac{k_F^{13}}{(2\pi)^{15}} (t_3^2 M) \int_{C_I} d^3 p_1 d^3 p_2 d^3 p_3 d^3 x d^3 y \frac{1}{\bar{x}^2 + 3\bar{y}^2/4 - H/6 - i\epsilon},$$

where the expression for C_I is given by:

$$C_I = \left[1 - \theta(1 - |\vec{p} \pm \vec{x} - \vec{y}/2|) \right] \left[1 - \theta(1 - |\vec{p} + \vec{y}|) \right] \theta(1 - |\vec{p}_1|) \theta(1 - |\vec{p}_2|) \theta(1 - |\vec{p}_3|).$$

- The chosen intermediate variables are: $\vec{p} + \vec{x} - \vec{y}/2$, $\vec{p} + \vec{y}$, $\vec{p} - \vec{x} - \vec{y}/2$, with $\vec{p} = (\vec{p}_1 + \vec{p}_2 + \vec{p}_3)/3$.
- The other kinematical quantity H appearing in the energy denominator is the Galilean invariant:

$$H = \frac{1}{k_F^2} [(\vec{p}_1 - \vec{p}_2)^2 + (\vec{p}_1 - \vec{p}_3)^2 + (\vec{p}_2 - \vec{p}_3)^2] < 9.$$

- We use the techniques of the dimensional regularization (DR) with a power-divergent subtraction (PDS) scheme because the second-order correction diverges at high momenta.
- After analyzing the asymptotic behavior of the scattering amplitude at high momenta, it becomes straightforward to analyze the divergent parts of the binding energy calculated at the second-order:

$$E^{div} \propto k_F^{13} \delta t_3^a \left[\frac{1}{3-d} + \frac{3}{2} - \gamma \right] - \mu^2 k_F^{11} \delta t_3^b \left[\frac{1}{2-d} - \gamma \right] + \mu k_F^{12} \delta t_3^c \left[\frac{1}{2-d} - \gamma \right] \\ + \mu^4 k_F^9 \delta t_3^d \left[\frac{1}{1-d} - \gamma \right]$$

- The terms $\delta t_3^{a,b,c,d}$ are proportional to t_3^2 ; μ is an auxiliary momentum parameter; γ is the Euler's constant.
- We see that E has poles:
 - 1 at $d = 3$ which corresponds to a logarithmic divergence ($\propto k_F^{13} \ln \Lambda$);
 - 2 at $d = 2$ which corresponds to a power-divergence ($\propto k_F^{12} \Lambda$ and $k_F^{11} \Lambda$);
 - 3 at $d = 1$ corresponds to a power-divergence ($\propto k_F^9 \Lambda^4$).

- The asymptotic behavior of the second-order energy correction can be written as:

$$E^{div} \propto \underbrace{k_F^{13} \delta t_3^a \left[\frac{1}{3-d} + \frac{3}{2} - \gamma \right]}_{l_1} - \underbrace{\mu^2 k_F^{11} \delta t_3^b \left[\frac{1}{2-d} - \gamma \right]}_{l_2} + \underbrace{\mu k_F^{12} \delta t_3^d \left[\frac{1}{2-d} - \gamma \right]}_{l_3} \\ + \underbrace{\mu^4 k_F^9 \delta t_3^d \left[\frac{1}{1-d} - \gamma \right]}_{l_4}$$

- The last term l_4 is proportional to k_F^9 . It can be regrouped with the mean-field three-body contribution.
- Counter terms:
 - l_1 is suppressed by adding a counter term: $\delta t_3^a \nabla_{12}^2 \delta(r_1 - r_2) \nabla_{23}^2 \delta(r_2 - r_3)$;
 - l_1 is suppressed by adding a counter term: $\delta t_3^b \nabla_{12}^2 \delta(r_1 - r_2) \delta(r_2 - r_3)$;
 - l_1 is suppressed by adding a counter term: $\delta t_3^c \delta(r_1 - r_2) \delta(r_2 - r_3) \delta(r_3 - r_4)$.