

Super-Kamiokande

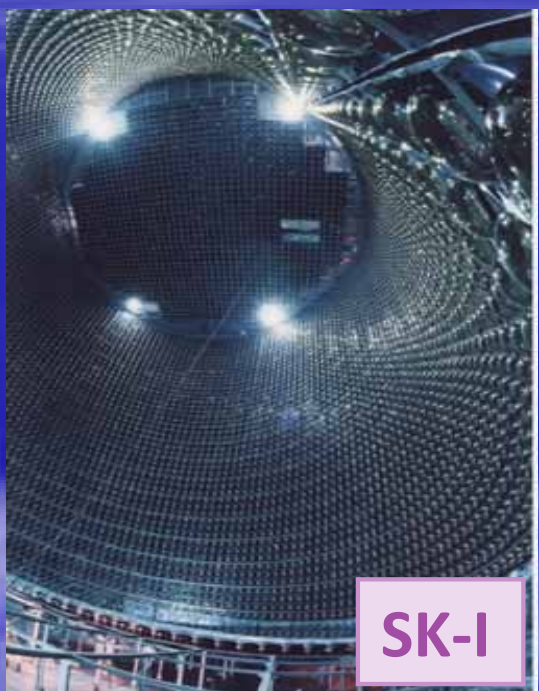
***Kamioka Observatory, Institute for Cosmic Ray Research, U of Tokyo, and
Kamioka Satellite, Institute for the Mathematics and Physics of the Universe, U of Tokyo***

***Masato Shiozawa**
for the Super-Kamiokande collaboration*

Super-Kamiokande History

inner detector mass: 32kton fiducial mass: 22.5kton

1996	1997	1998	1999	2000	2001	2002	2003	2004	2005	2006	2007	2008	2009
------	------	------	------	------	------	------	------	------	------	------	------	------	------



SK-I

11146 ID PMTs
(40% coverage)

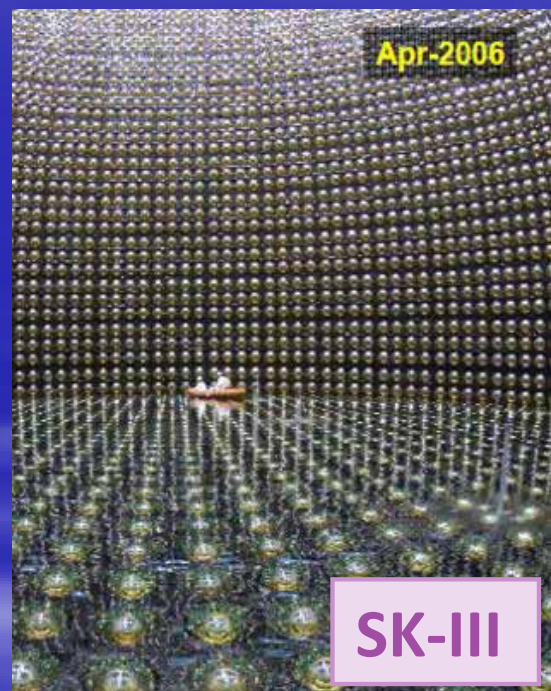
Energy Threshold **5.0 MeV**
(total electron energy)



SK-II

5182 ID PMTs
(19% coverage)

7.0 MeV



SK-III

11129 ID PMTs
(40% coverage)

4.5 MeV
work in progress



SK-IV

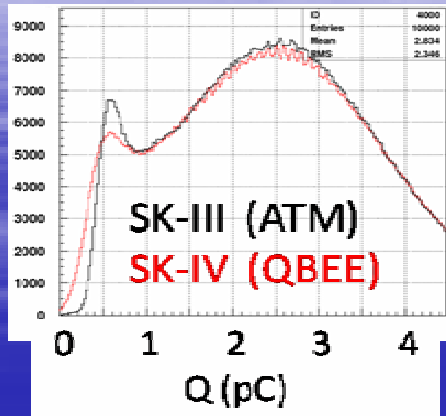
Electronics Upgrade

< 4.0 MeV
target

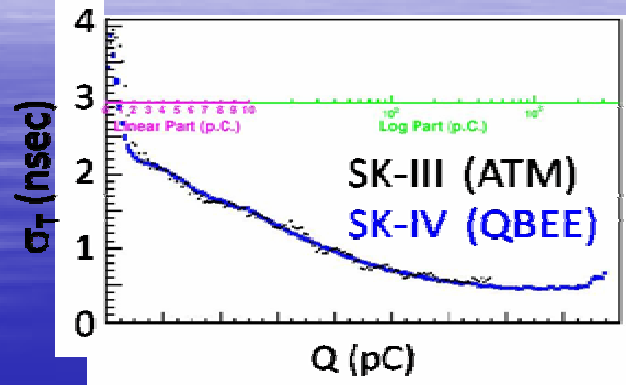
SK-IV commissioning since Sep.-2008

detector calibrations and detector simulation tuning

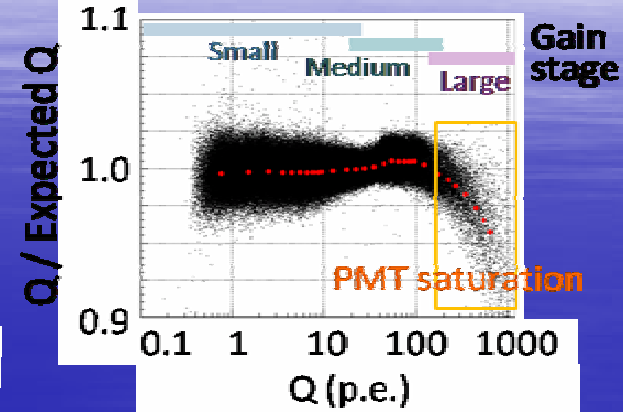
Single p.e. distribution (by Ni)



Timing resolution (by Laser)

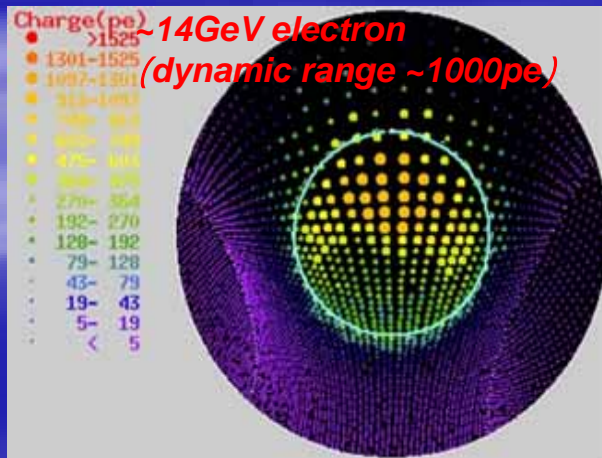


Charge linearity (by Laser)

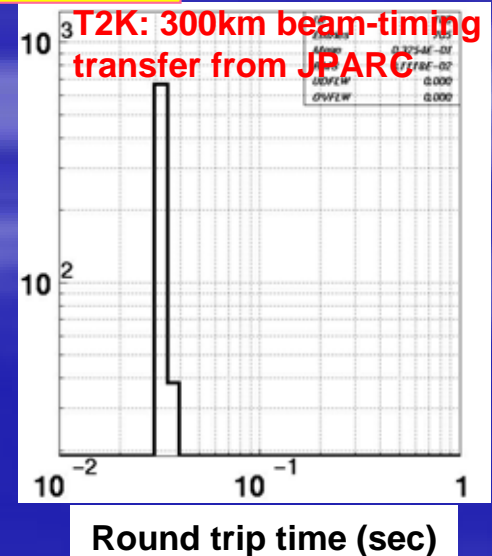
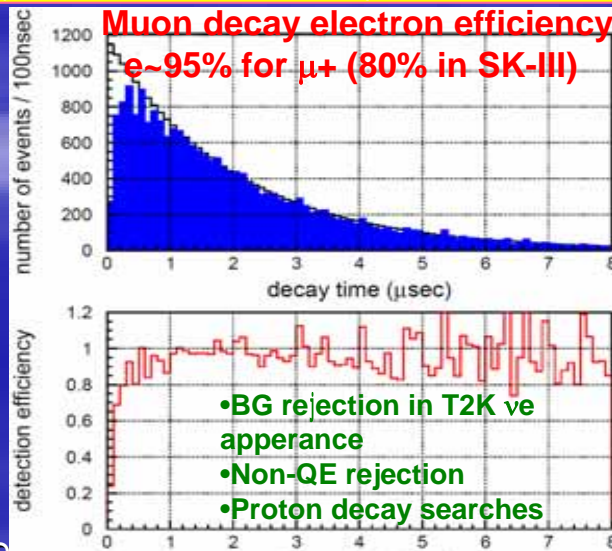


Aiming to expand observation windows

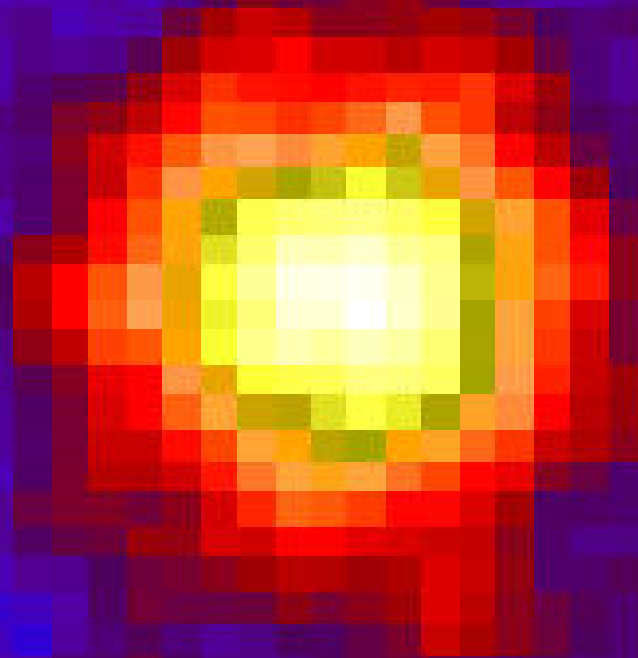
- Lower energy (<5MeV) solar ν and high energy (>10GeV) atmospheric ν
- Better efficiency of muon decay electrons
- Nearby supernova (~0.4kpc, 10,000,000events)
- Neutron tagging from anti electron neutrino interactions
- Observe accelerator ν (T2K from this winter)



ν_e appearance (θ_{13}) in atm ν

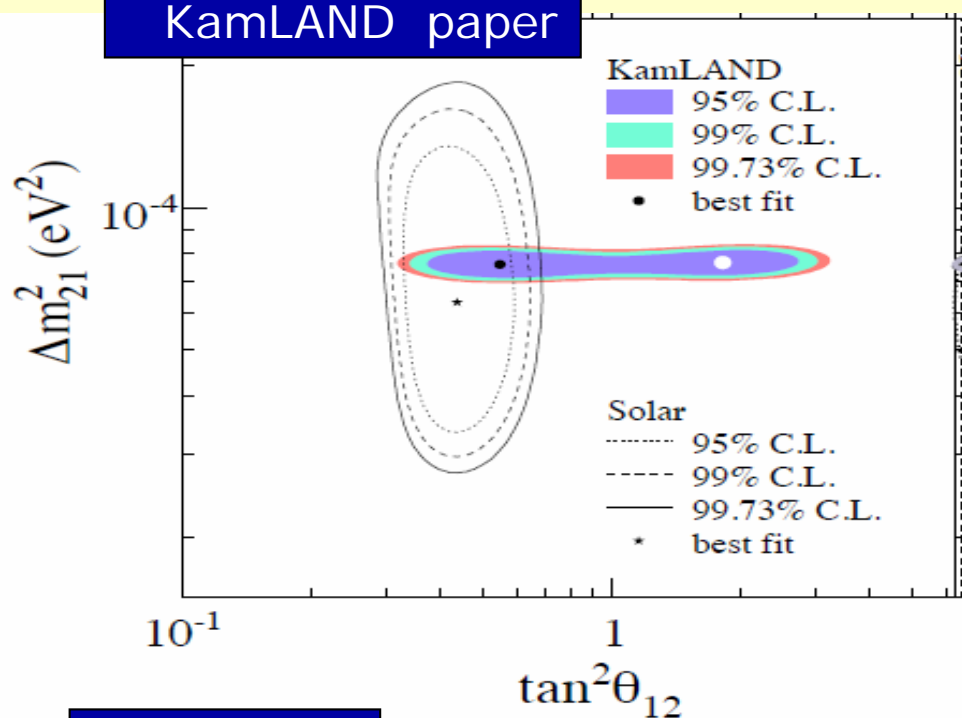


Solar Neutrinos

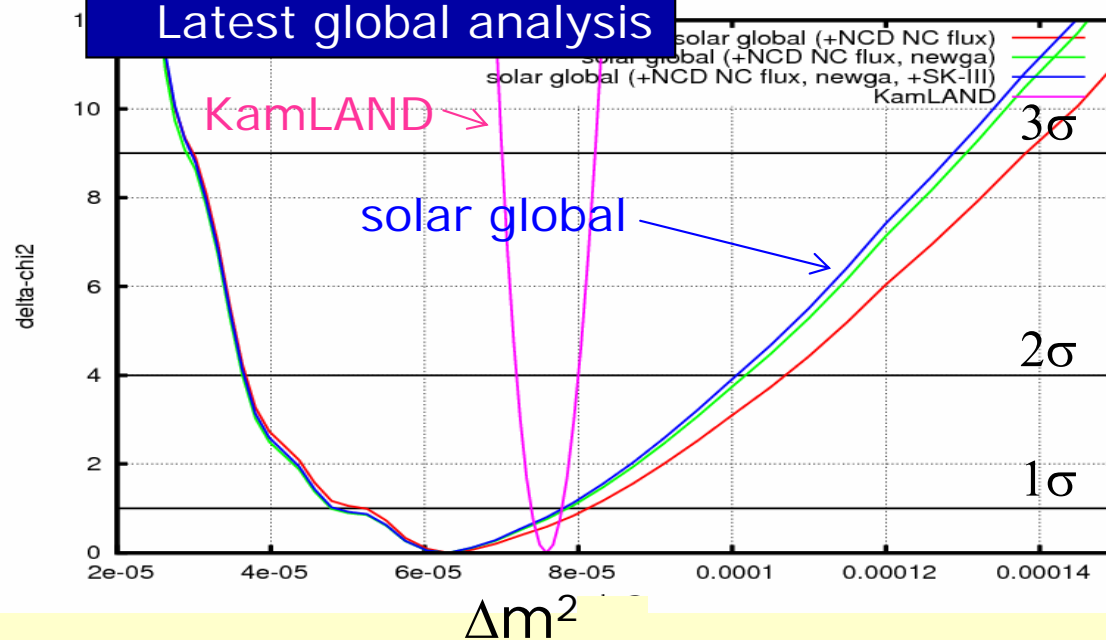


Solar global and KamLAND

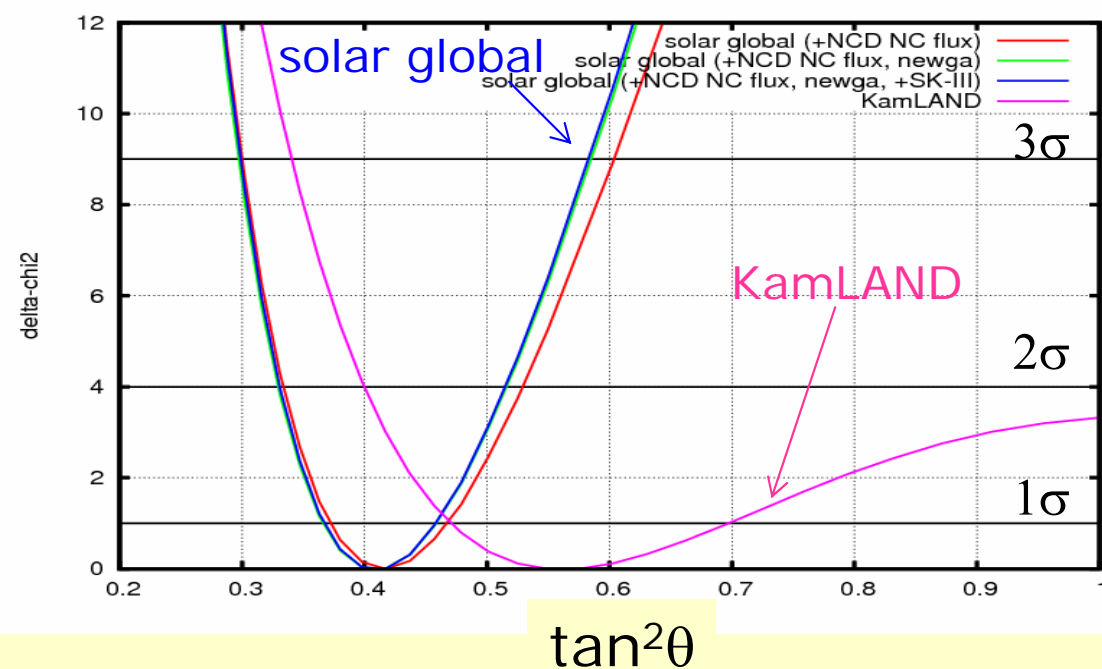
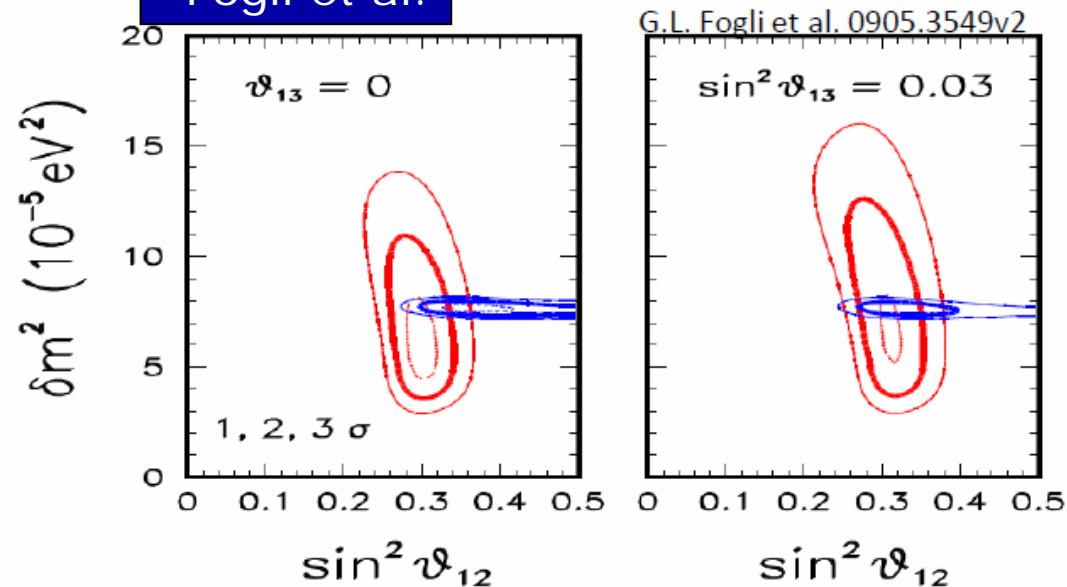
KamLAND paper



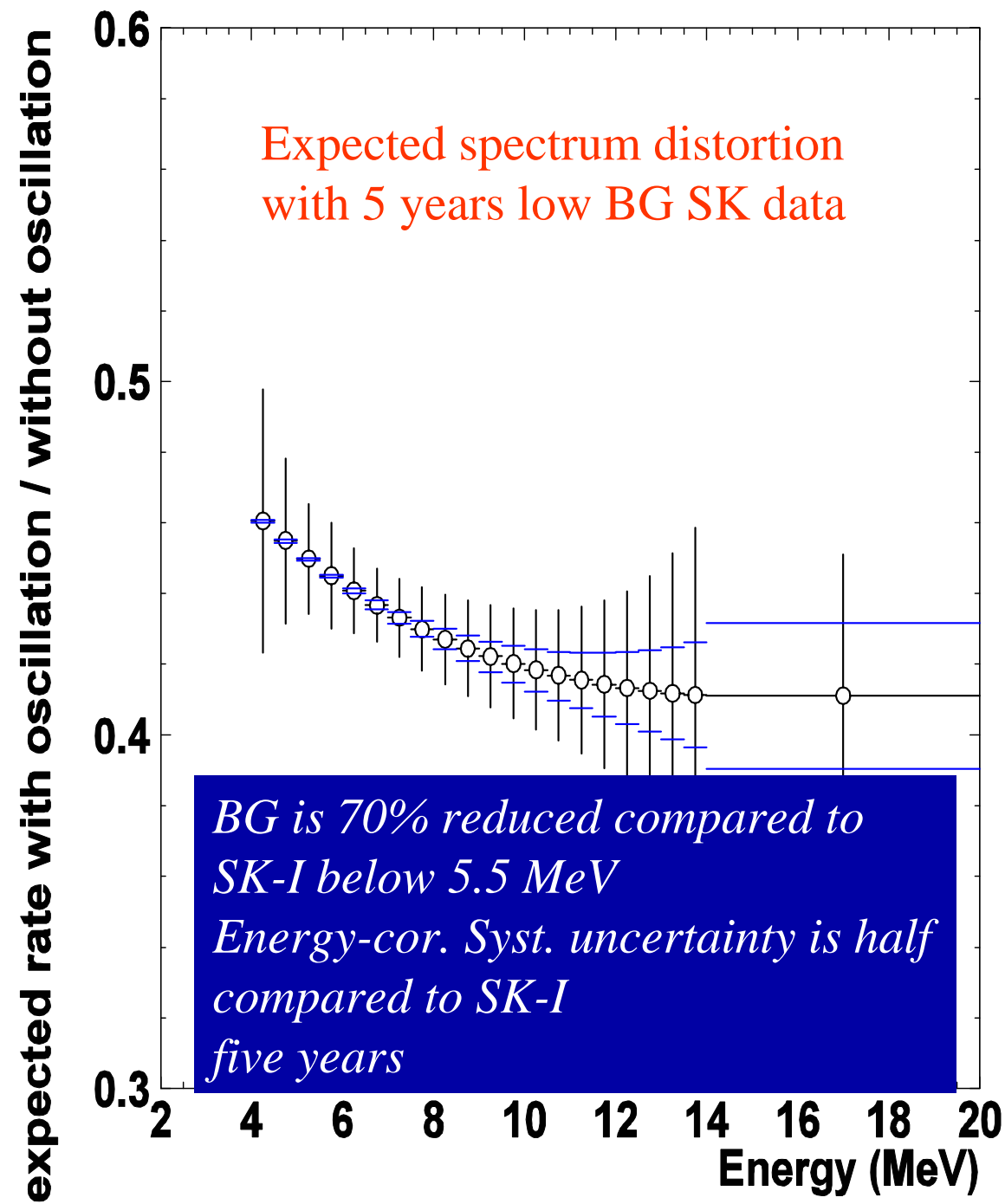
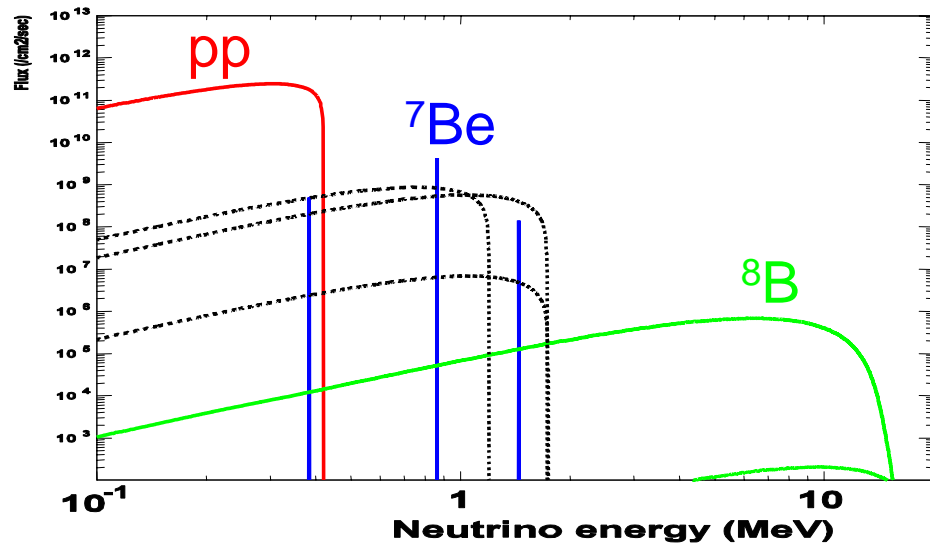
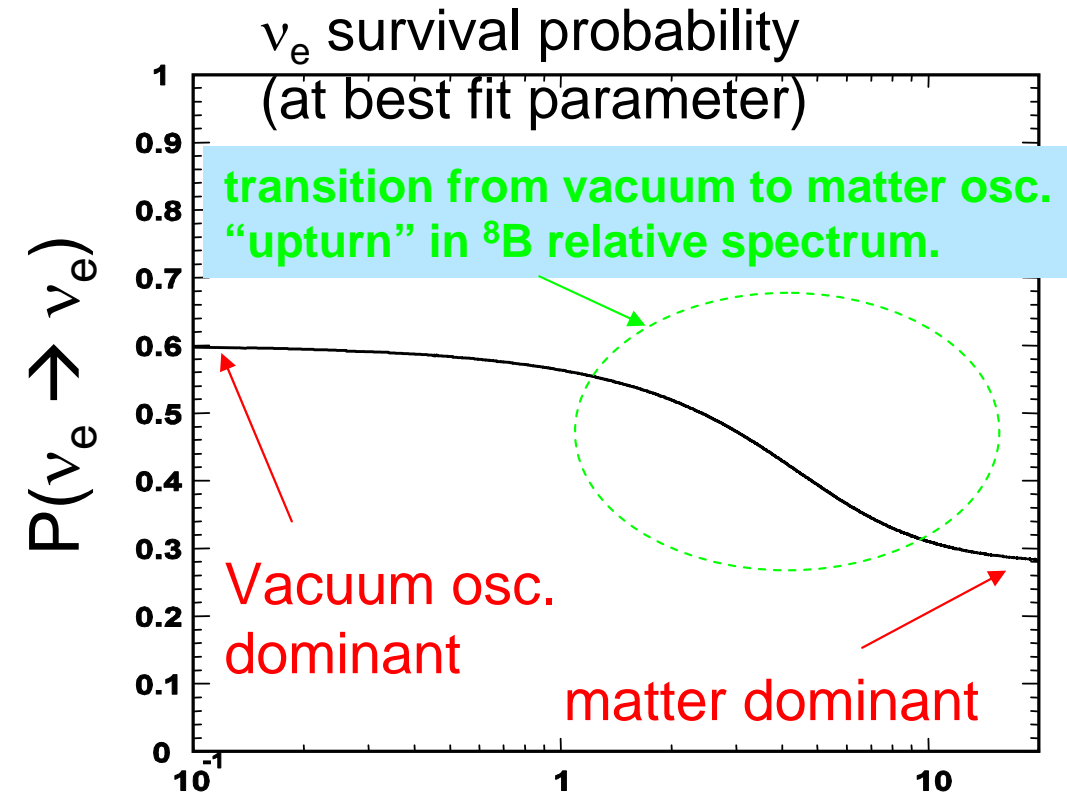
Latest global analysis



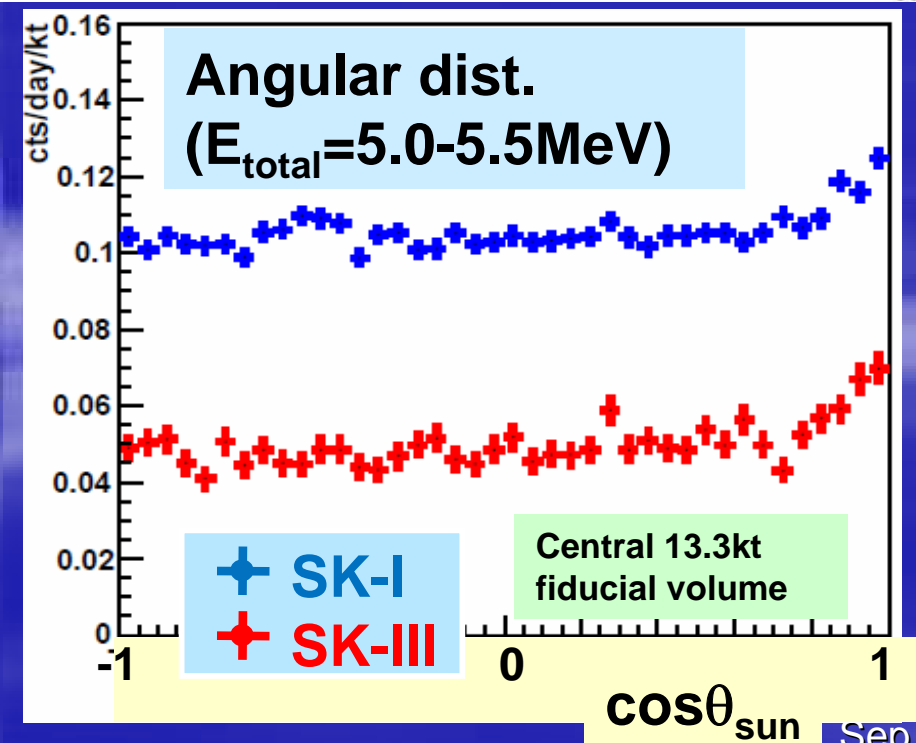
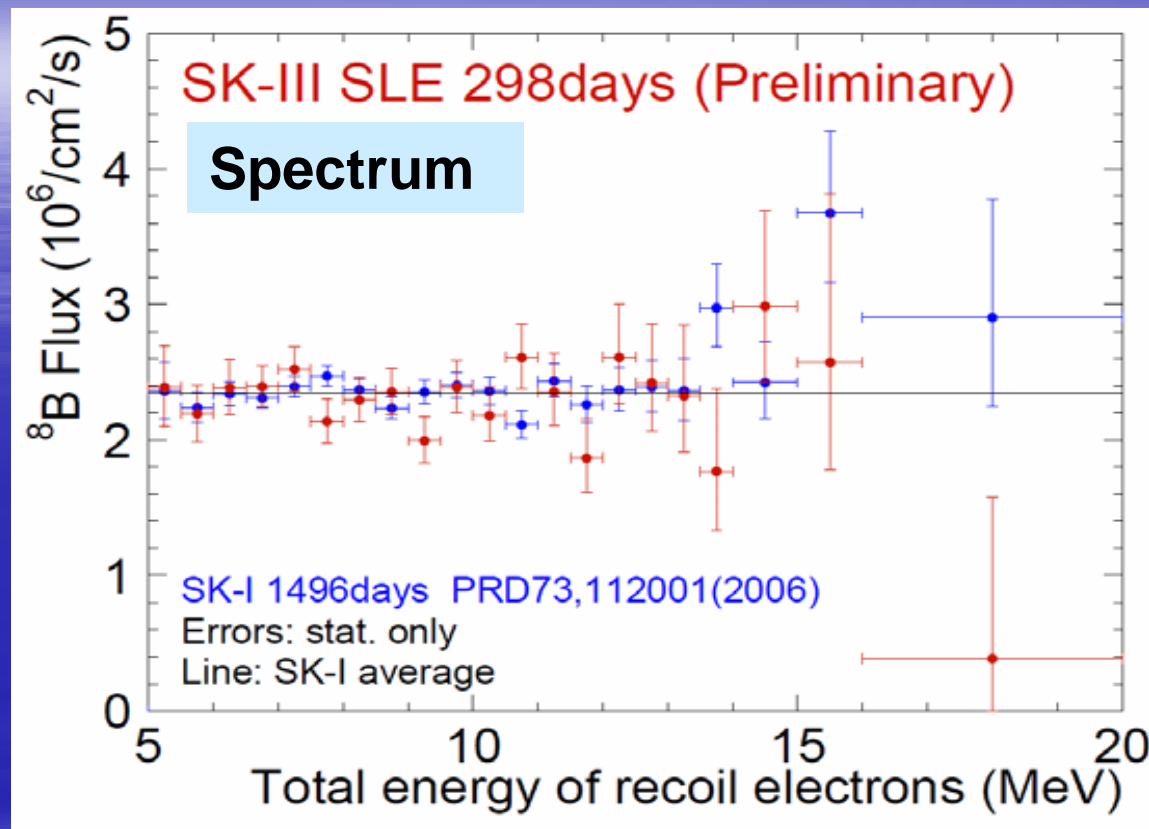
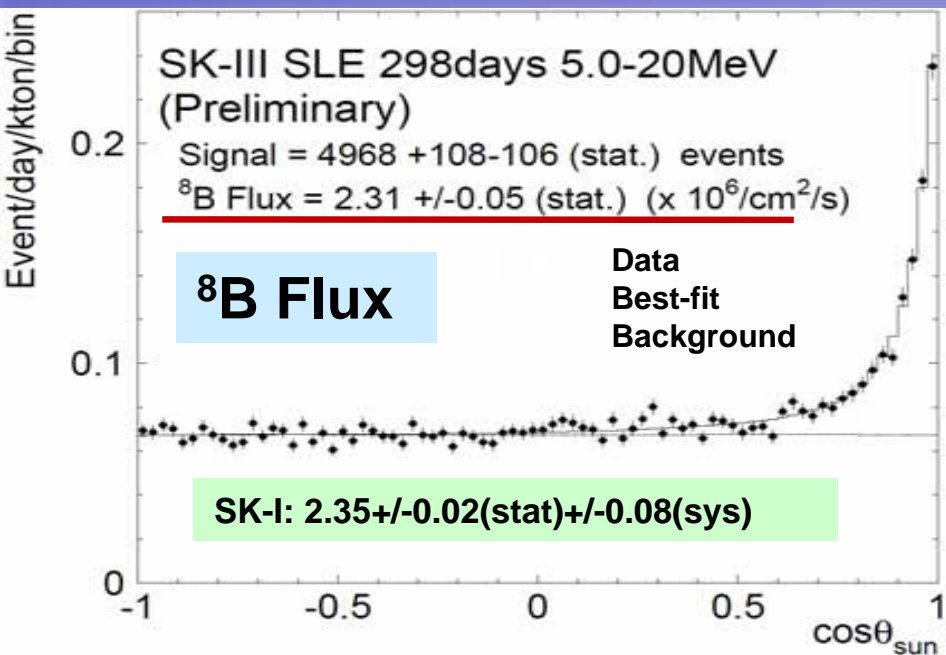
Fogli et al.



Solar Neutrino Future Prospects in SK



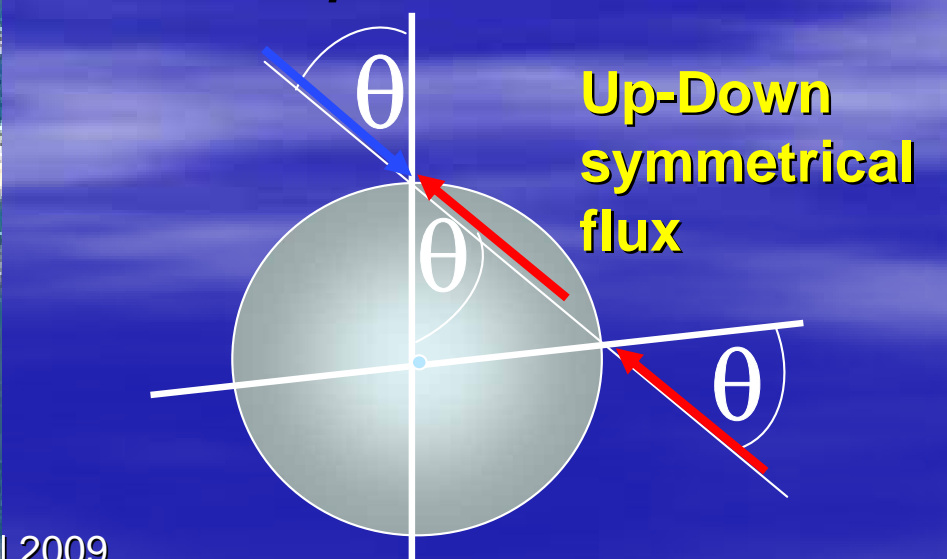
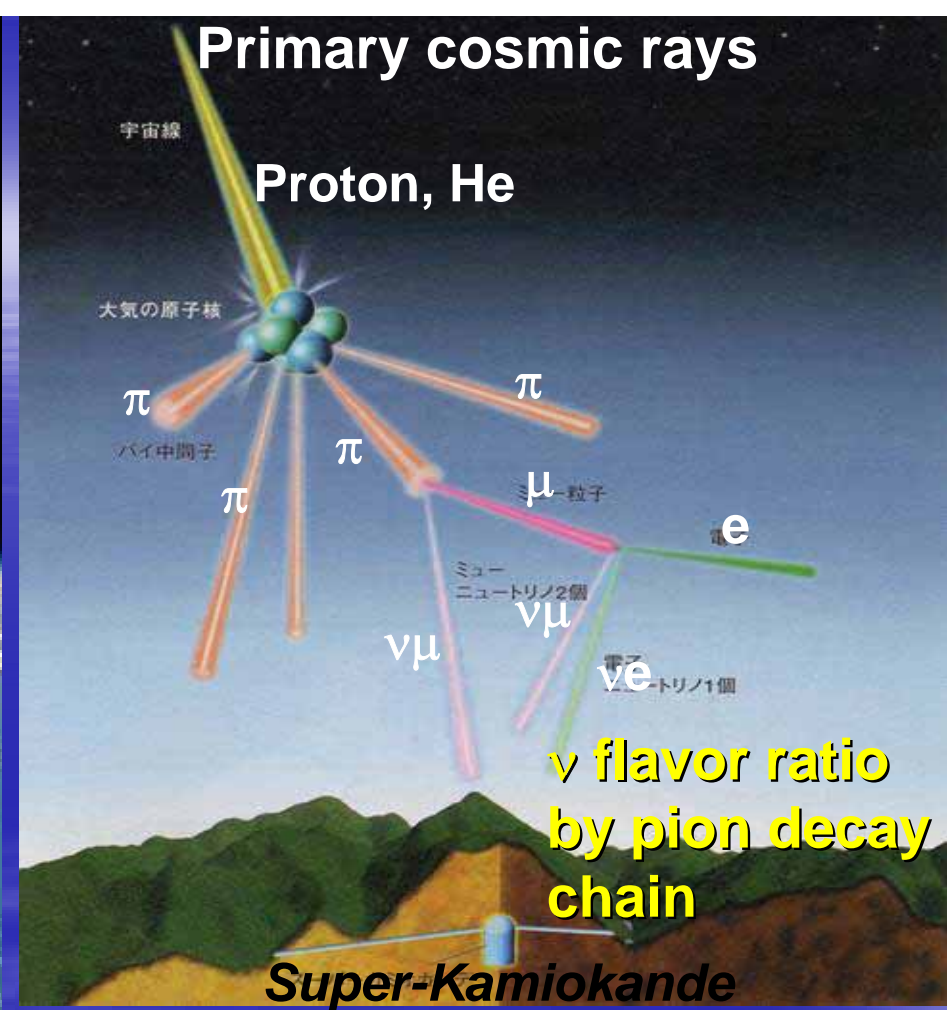
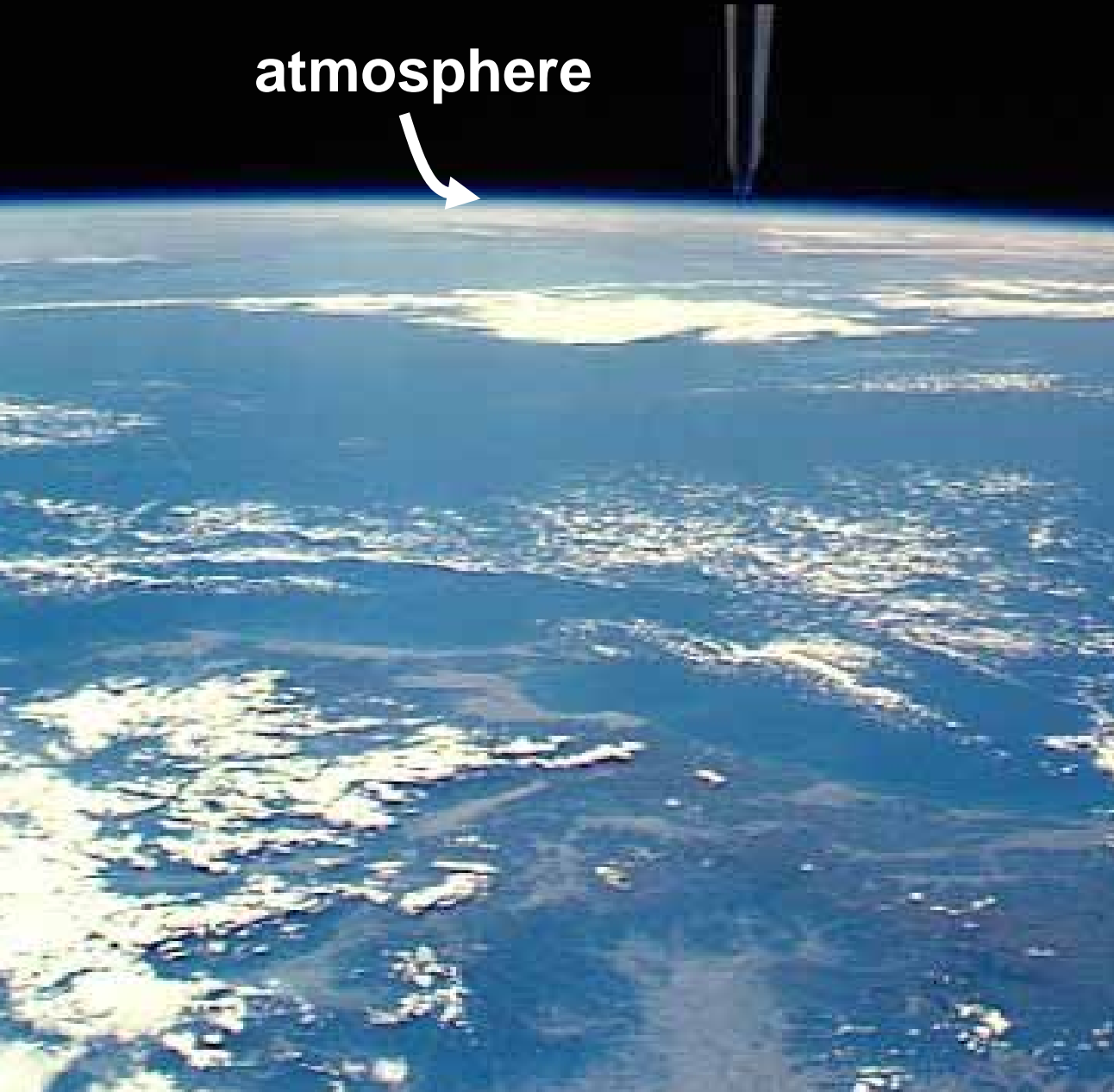
recovered original threshold (SK-III)



- We tuned reconstruction program and analysis programs. (almost done)
- Better angular resolution: 10% better than SK-I
- Lower BG rate than before is obtained in SK-III low energy regions

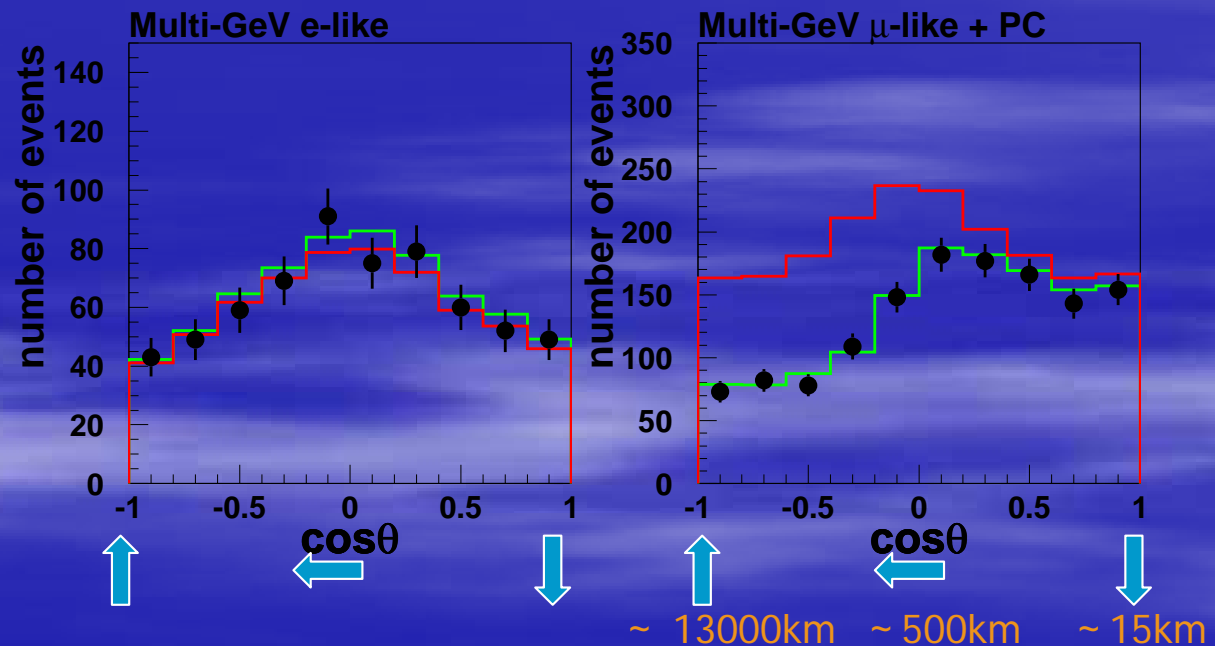
Atmospheric ν 's

atmosphere



discovery of neutrino oscillation

- 1998, Atmospheric neutrino observation at Super-Kamiokande
 - deficit of upward going muon (neutrinos)
 - electron (neutrino) as is expected
 - consistent with pure $\nu_{\mu} \rightarrow \nu_{\tau}$
 - $\sin^2 \theta_{23} > 0.82, 5 \times 10^{-4} < \Delta m^2_{23} < 6 \times 10^{-3} \text{ eV}^2$

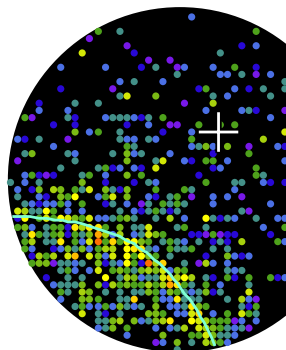


**NOW: More than 28,000 events have been recorded.
Provide evidence and still provide largest statistics**

Particle ID and the number of Cherenkov rings

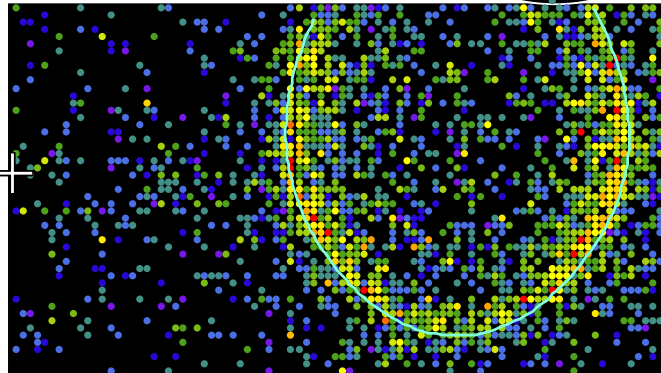
Super-Kamiokande

Run 5704 Event 3551590
 98-03-17:07:14:39
 Inner: 3397 hits, 7527 pE
 Outer: 0 hits, 0 pE (in-time)
 Trigger ID: 0x07
 D wall: 1089.6 cm
 FC e-like, $p = 923.2$ MeV/c

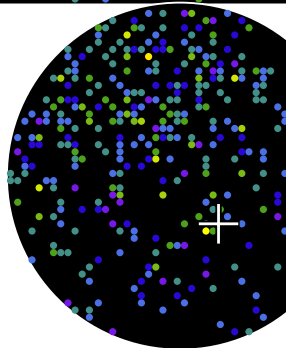


Charge (pe)

- >15.0
- 13.1-15.0
- 11.4-13.1
- 9.8-11.4
- 8.2- 9.8
- 6.9- 8.2
- 5.6- 6.9
- 4.5- 5.6
- 3.5- 4.5
- 2.6- 3.5
- 1.9- 2.6
- 1.2- 1.9
- 0.8- 1.2
- 0.4- 0.8
- 0.1- 0.4
- < 0.1

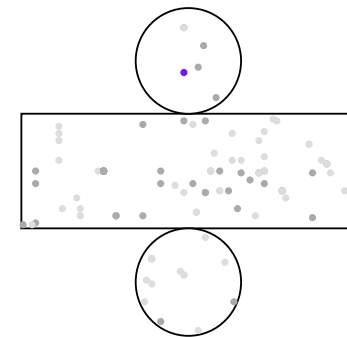
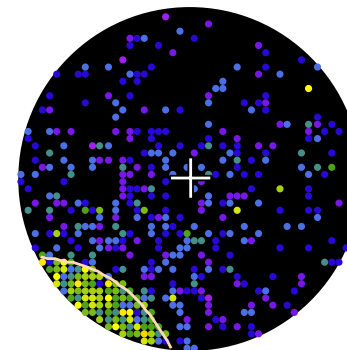


Electron-like ring
(diffused ring)



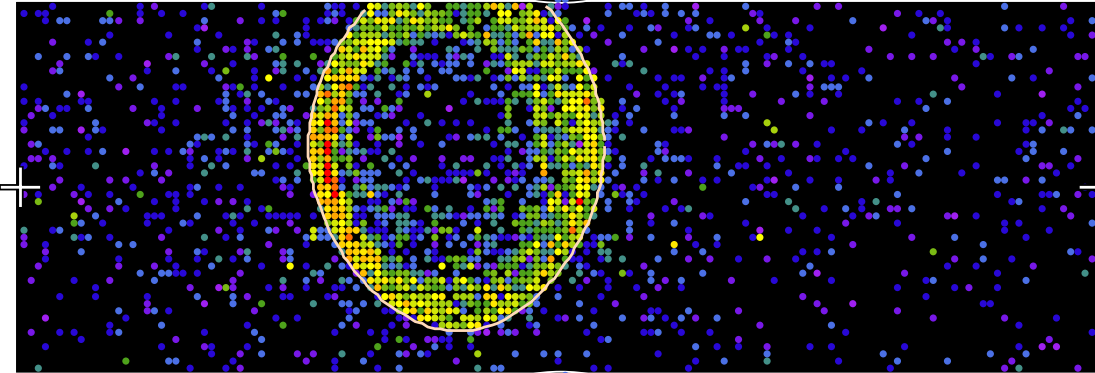
Super-Kamiokande

Run 3962 Sub 125 Ev 965982
 97-05-01:15:32:29
 Inner: 2887 hits, 9607 pE
 Outer: 1 hits, 0 pE (in-time)
 Trigger ID: 0x03
 D wall: 1690.0 cm
 FC mu-like, $p = 1323.6$ MeV/c

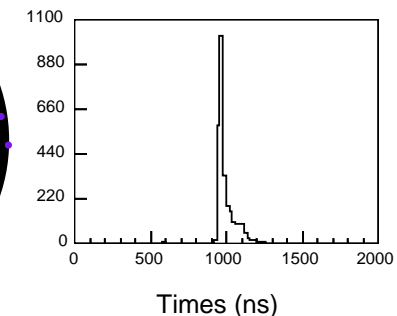
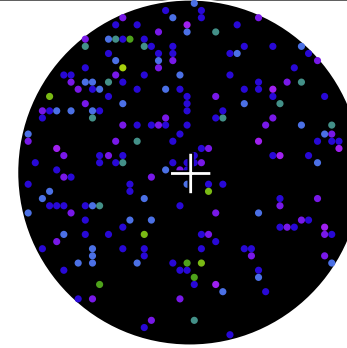


Charge (pe)

- >26.7
- 23.3-26.7
- 20.2-23.3
- 17.3-20.2
- 14.7-17.3
- 12.2-14.7
- 10.0-12.2
- 8.0-10.0
- 6.2- 8.0
- 4.7- 6.2
- 3.3- 4.7
- 2.2- 3.3
- 1.3- 2.2
- 0.7- 1.3
- 0.2- 0.7
- < 0.2



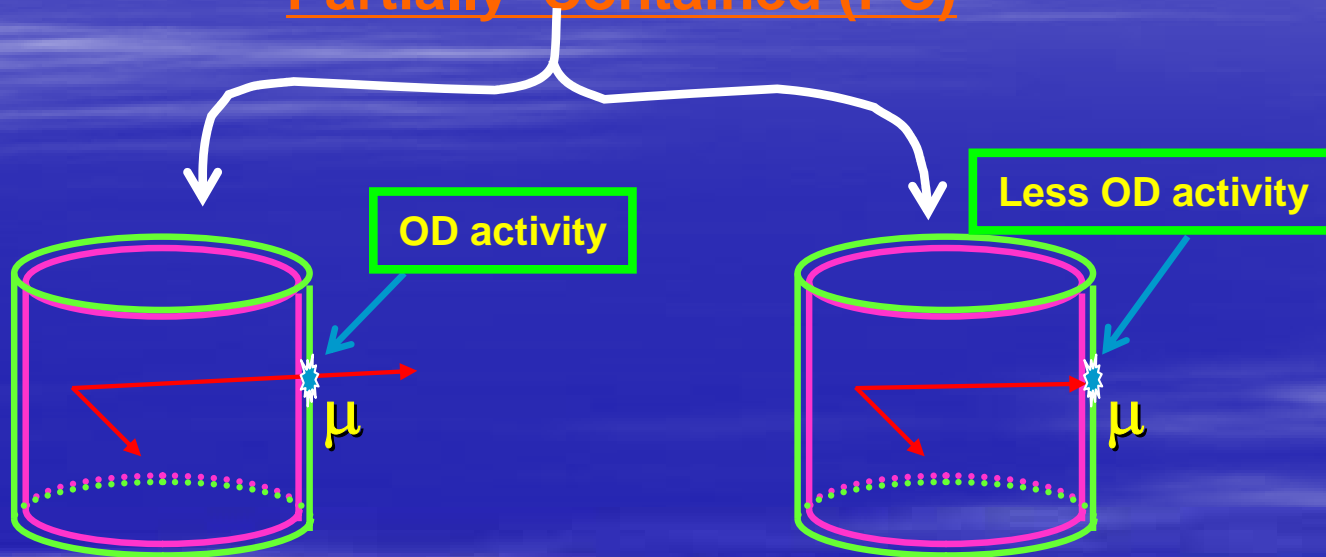
muon-like ring
(sharp edge)



Event category 2 (PC and $up\mu$)

Particle ID and # of Cherenkov rings are not used.

Partially Contained (PC)



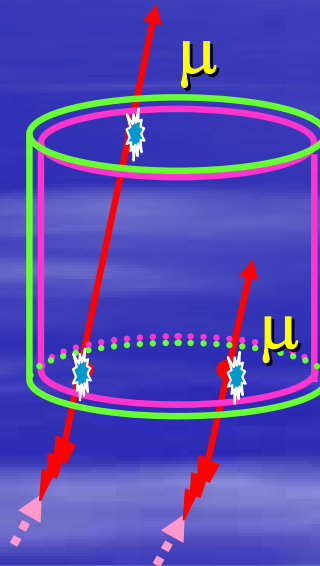
OD through going PC

OD stopping PC

- discrimination of PC stop and through by deposited energy in the outer detector
- PC stop is a kind of FC events and energy can be reconstructed.

Upward μ

Upward through-going μ

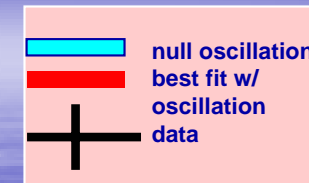
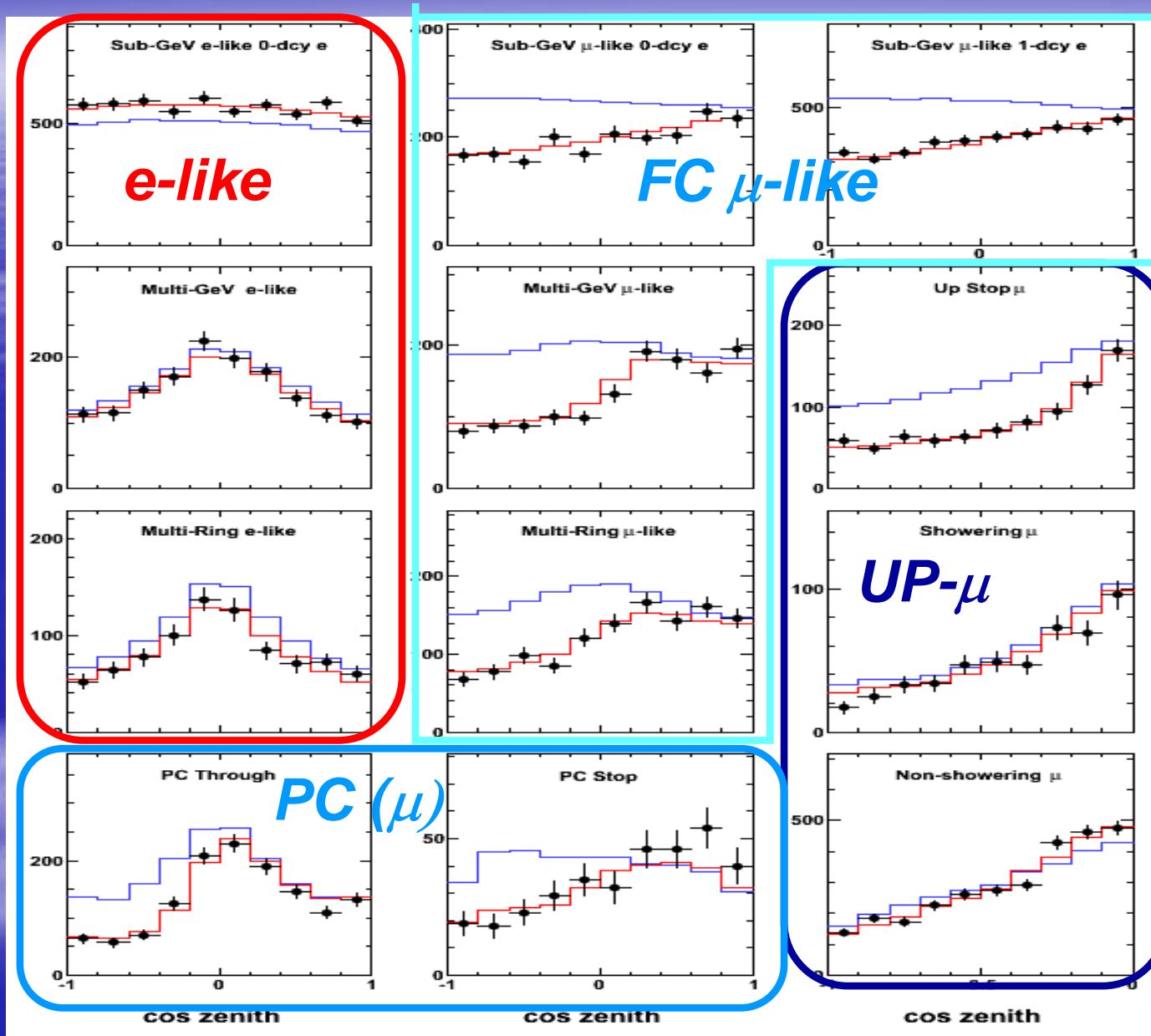


Upward stopping μ

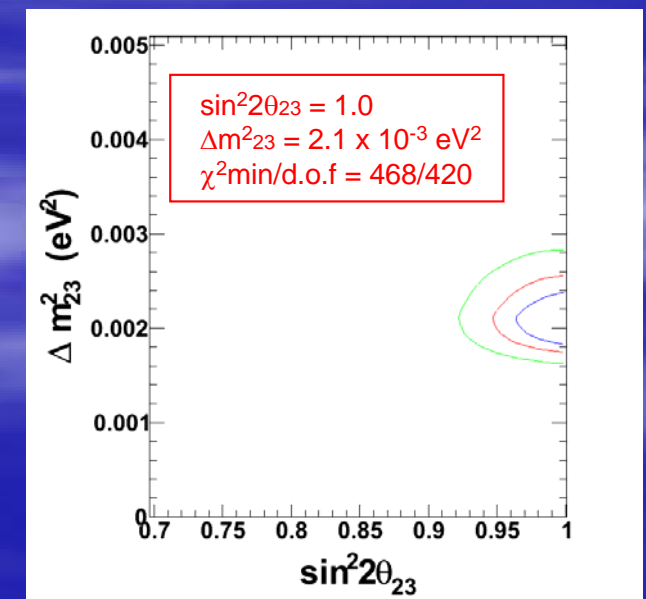
- target is rock (water for FC and PC)
- different energy scale and detection technique

2 flavor oscillation analysis

zenith angle (SK-I +II+III combined)

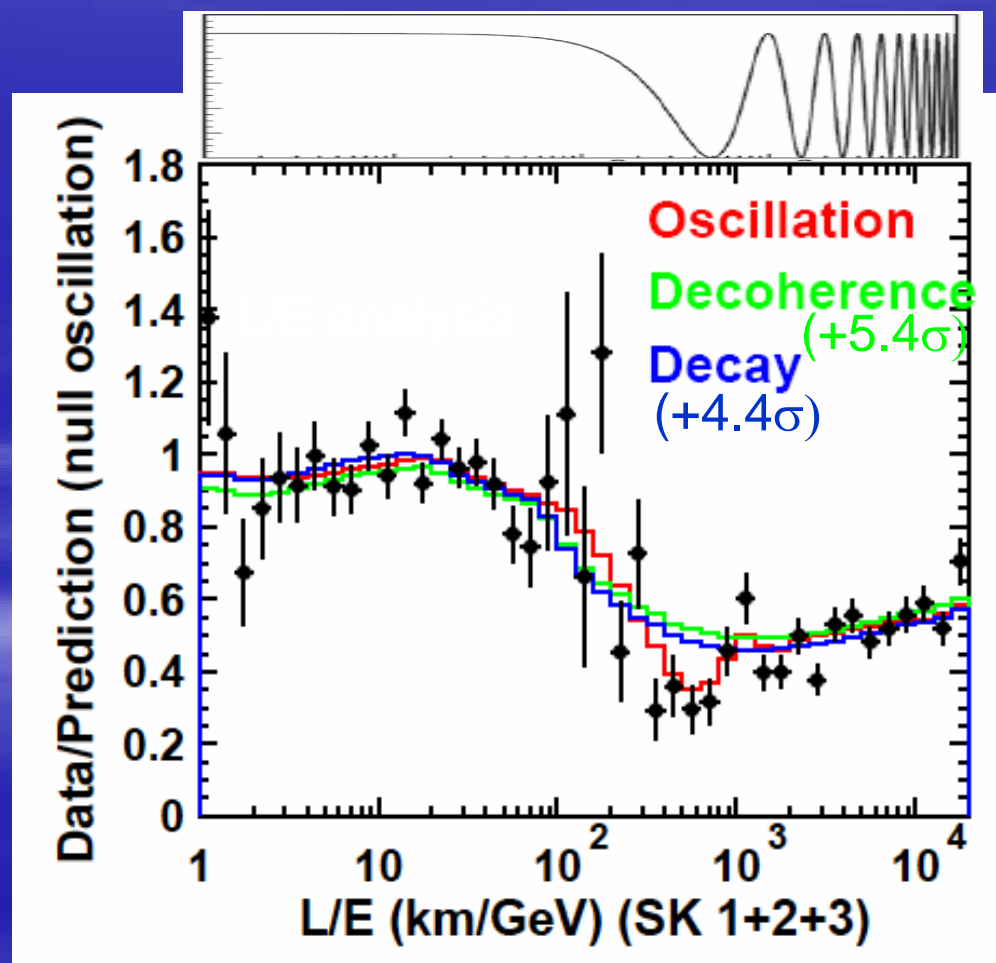
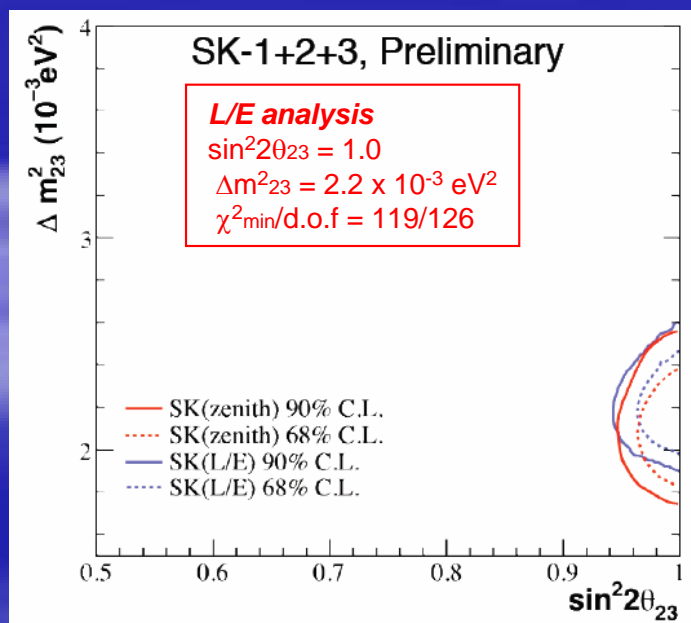


- Latest results
 - SKI+II+III combined
 - 2806days (173ktyr) for FC+PC
 - 24841 events
 - 3109days for UP- μ
 - 4238 events

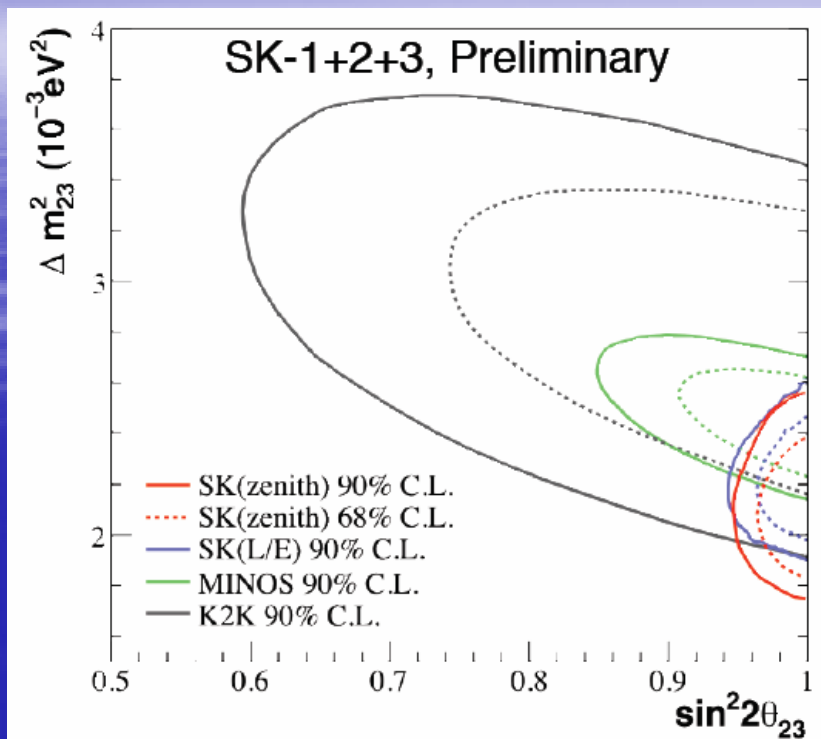


L/E analysis (SK-I +II+III)

- Zenith analysis
 - Use almost all sub-samples, binned by zenith angle
- L/E analysis
 - Select events with good L/E ($\Delta(L/E) < 70\%$)
 - Binned by L/E
 - Position of the dip $(L/E)^{-1} \sim \Delta m^2$
 - Aim to directly determine Δm^2
 - Do see the dip at $L/E \sim 500$



Compilation of measurements



- $\Delta m^2_{23} = 2.11^{+0.11}_{-0.19} \times 10^{-3} \text{ eV}^2$ (68% ($\Delta\chi^2=1$)) (atm ν : zenith)
- $\Delta m^2_{23} = 2.19^{+0.14}_{-0.13} \times 10^{-3} \text{ eV}^2$ (68%) (atm ν : L/E)
- $\Delta m^2_{23} = 2.43^{+0.13}_{-0.13} \times 10^{-3} \text{ eV}^2$ (68%) (MINOS, PRL101(08))

→ Δm^2 : comparable accuracies

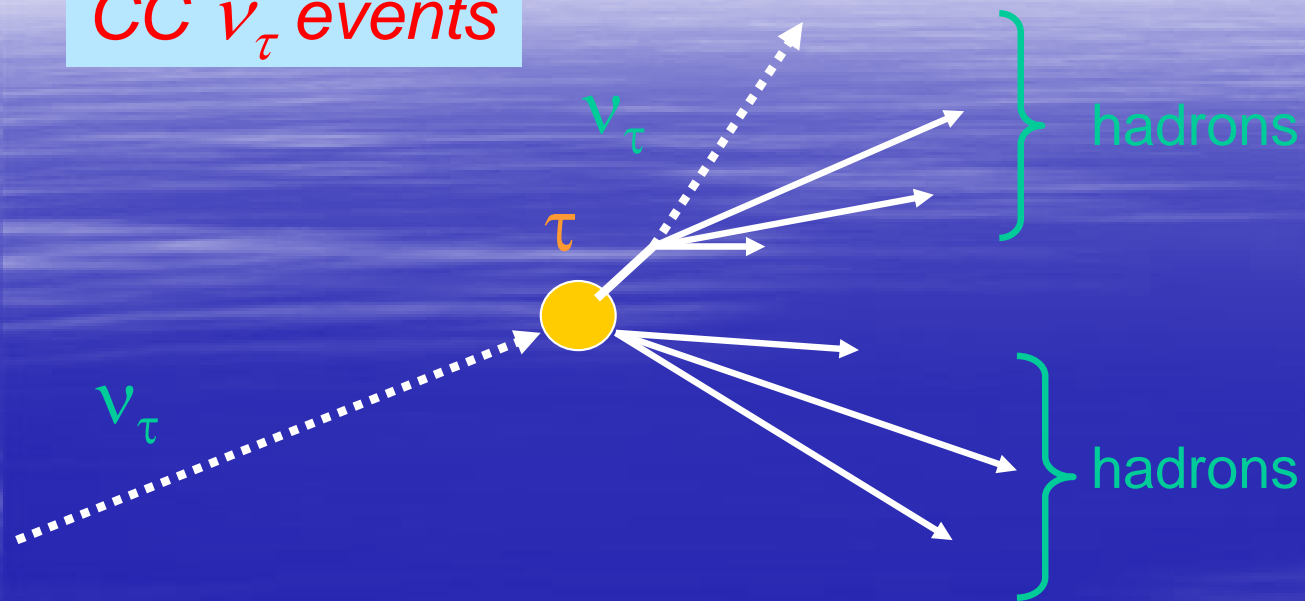
- $\sin^2 2\theta > 0.96$ (90% ($\Delta\chi^2=2.7$)) (atm ν)
- $\sin^2 2\theta > 0.90$ (90%) (MINOS)

→ $\sin^2 2\theta$: better by atm ν

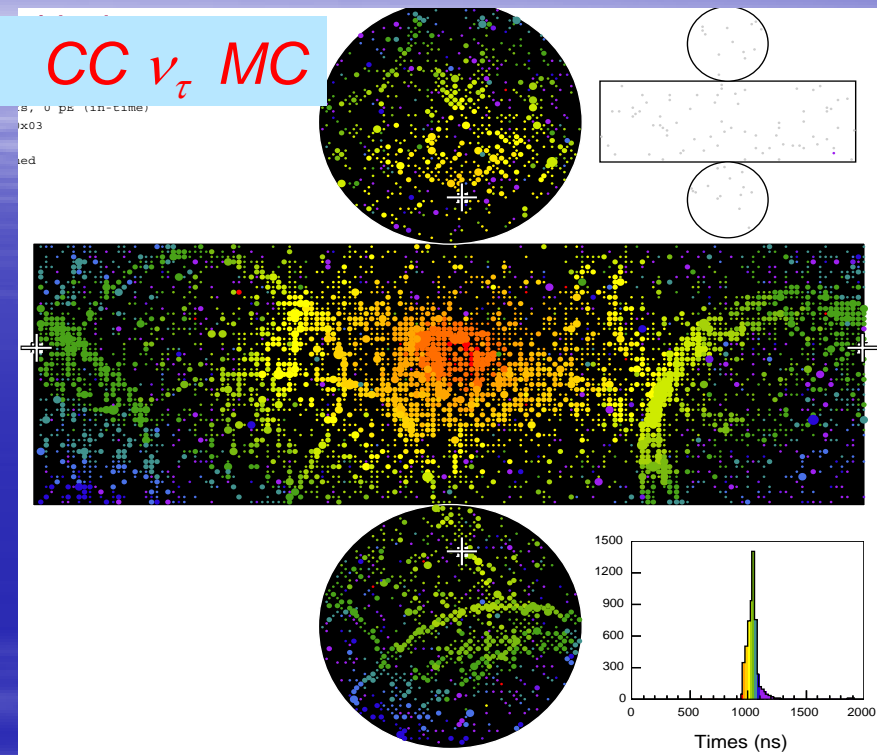
- future: - atm ν $1/\sqrt{\text{statistics}}$
- LBL E will be probably better

Search for CC ν_τ events

CC ν_τ events



CC ν_τ MC



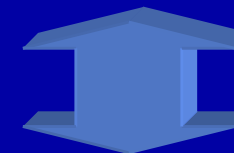
Signature of CC ν_τ events

- Higher multiplicity of Cherenkov rings
- More μ e decay signals
- Spherical event pattern



Likelihood and neural network analysis

Only ~ 1.0 CC ν_τ FC
events/kton \cdot yr

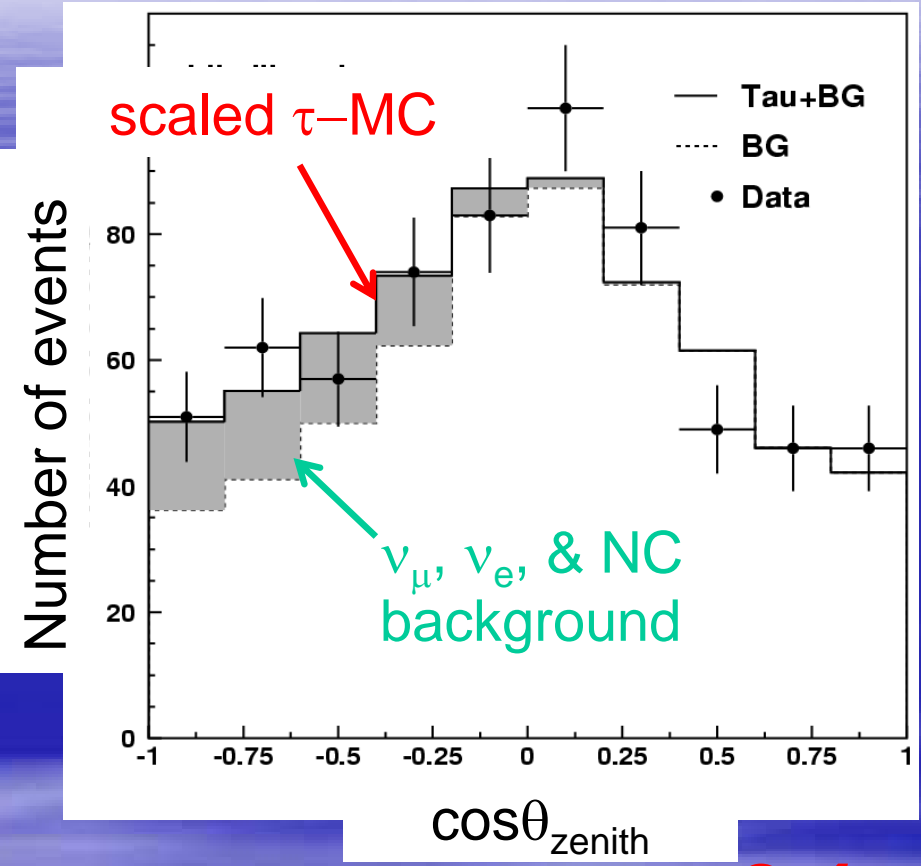


(BG (other ν events) \sim
130 ev./kton \cdot yr)

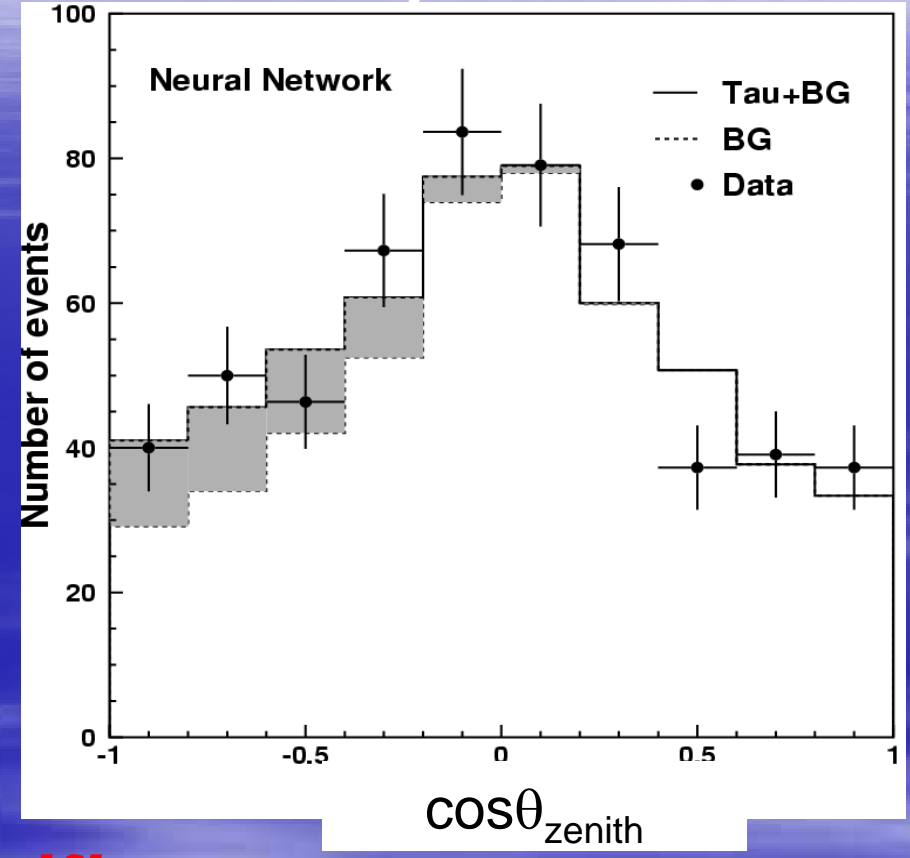
OLD
(to be updated)

Zenith angle dist. and fit results

Likelihood analysis



NN analysis



2.4 σ significance

Fitted # of τ events

Expected # of τ events

138 \pm 48(stat) +15 / -32(syst)	134 \pm 48(stat) +16 / -27(syst)
78 \pm 26(syst)	78 \pm 27 (syst)

Soon; 1, combine both analyses to improve signal/BG
 2, add SK-II + III (double statistics)

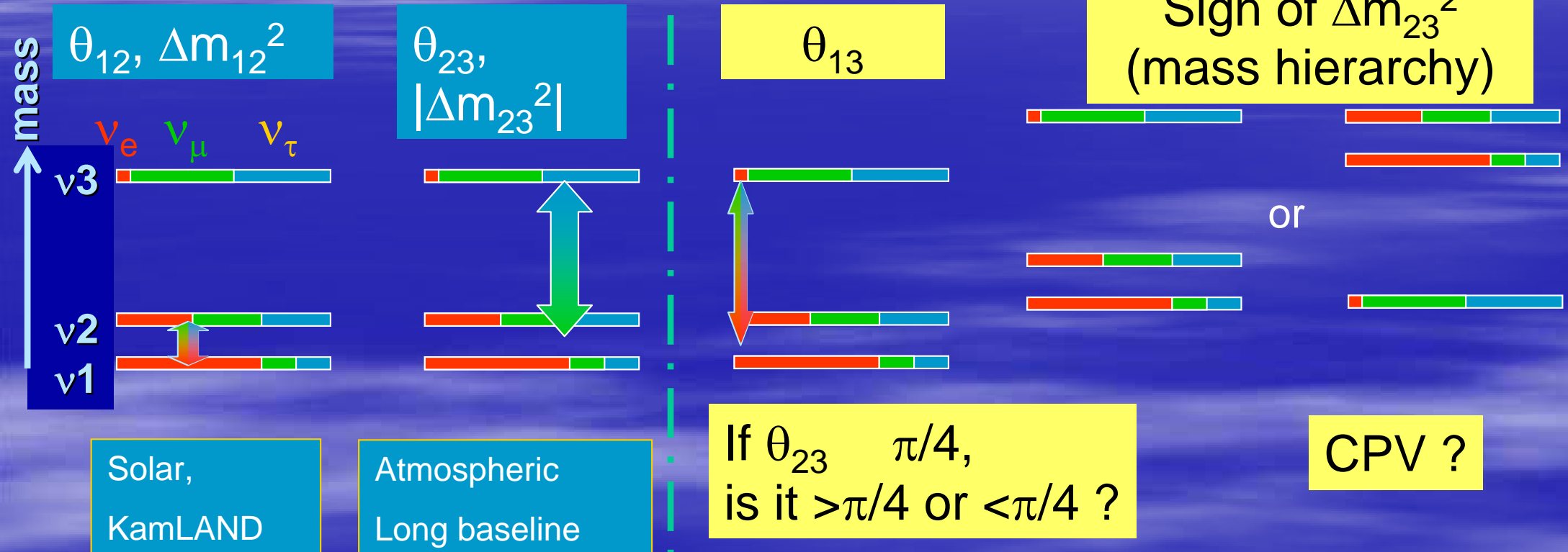
Remaining problems for atm- ν

ν mass and mixing parameters:

$$\theta_{12}, \theta_{23}, \theta_{13}, \delta, \Delta m_{12}^2, \Delta m_{13}^2 (= \Delta m_{23}^2)$$

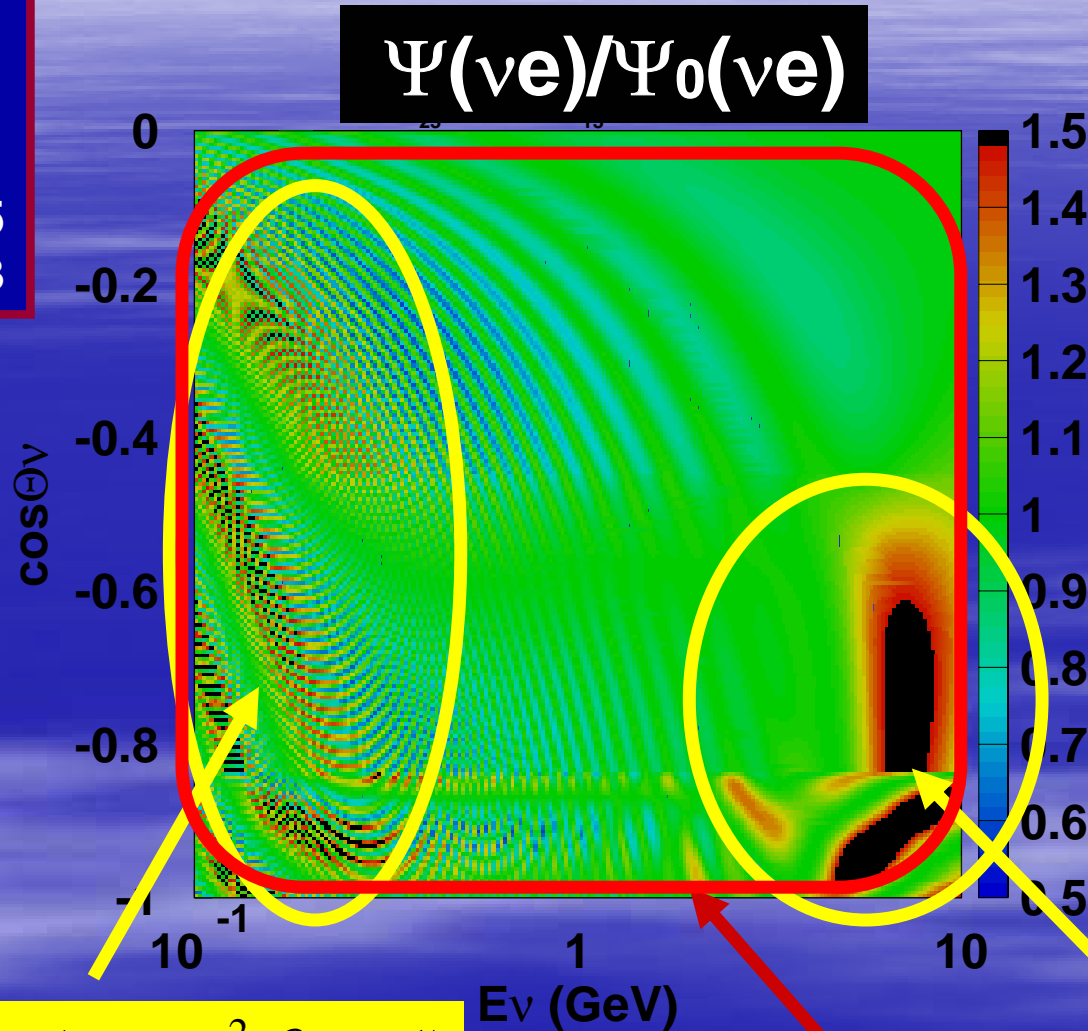
Known:

Unknown:



Global picture of oscillation effects

$s^2\theta_{12}=0.825$
 $s^2\theta_{23}=0.4$
 $s^2\theta_{13}=0.04$
 $\delta_{CP}=45^\circ$
 $\Delta m^2_{12}=8.3e-5$
 $\Delta m^2_{23}=2.5e-3$



Pares and Smirnov hep - ph/0309312 (at Sub - GeV range)

$$\frac{\Psi(\nu_e)}{\Psi_0(\nu_e)} - 1 \cong P_2(r \cdot c_{23}^2 - 1) \quad \text{LMA}$$

$$- r \cdot \tilde{s}_{13} \cdot \tilde{c}_{13}^2 \cdot \sin 2\mathcal{G}_{23}(\cos \delta_{CP} \cdot R_2 - \sin \delta_{CP} \cdot I_2) \quad \text{interference}$$

$$+ 2\tilde{s}_{13}^2(r \cdot s_{23}^2 - 1) \quad \mathcal{G}_{13} \text{ resonance}$$

r : μ/e flux ratio (~ 2 at low energy)

$P_2 = |A_{e\mu}|^2$: 2ν transition probability $\nu_e \rightarrow \nu_{\mu\tau}$ in matter

$R_2 = \text{Re}(A_{ee}^* A_{e\mu})$

$I_2 = \text{Im}(A_{ee}^* A_{e\mu})$

A_{ee} : survival amplitude of the 2ν system

$A_{e\mu}$: transition amplitude of the 2ν system

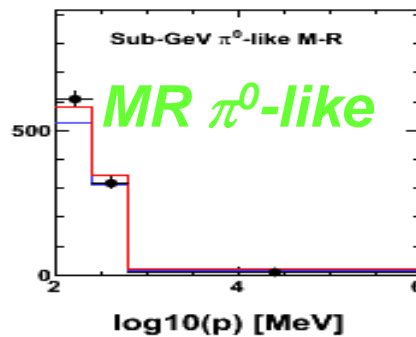
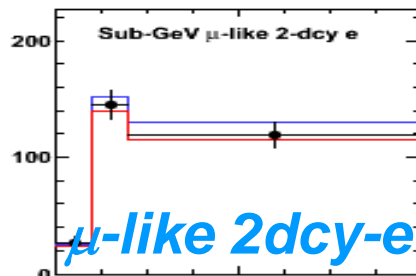
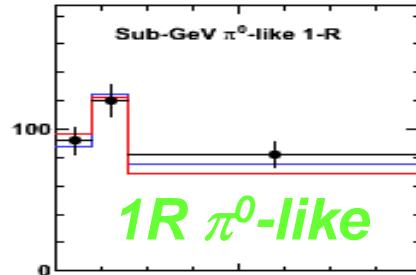
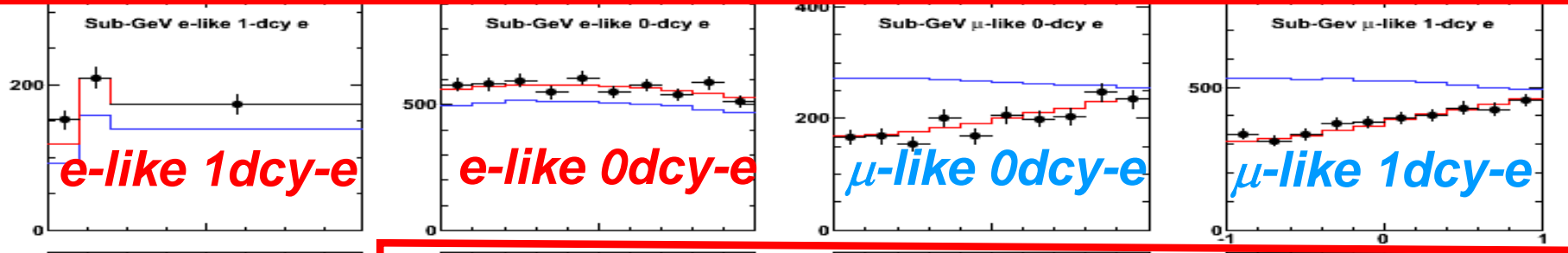
LMA $\propto (r \cdot \cos^2 \mathcal{G}_{23} - 1)$

interference (CPV)

\mathcal{G}_{13} resonance

$\propto \sin^2 \mathcal{G}_{13}^m (r \cdot \sin^2 \mathcal{G}_{23} - 1)$

new sub-samples (SK-I +II+III combined)

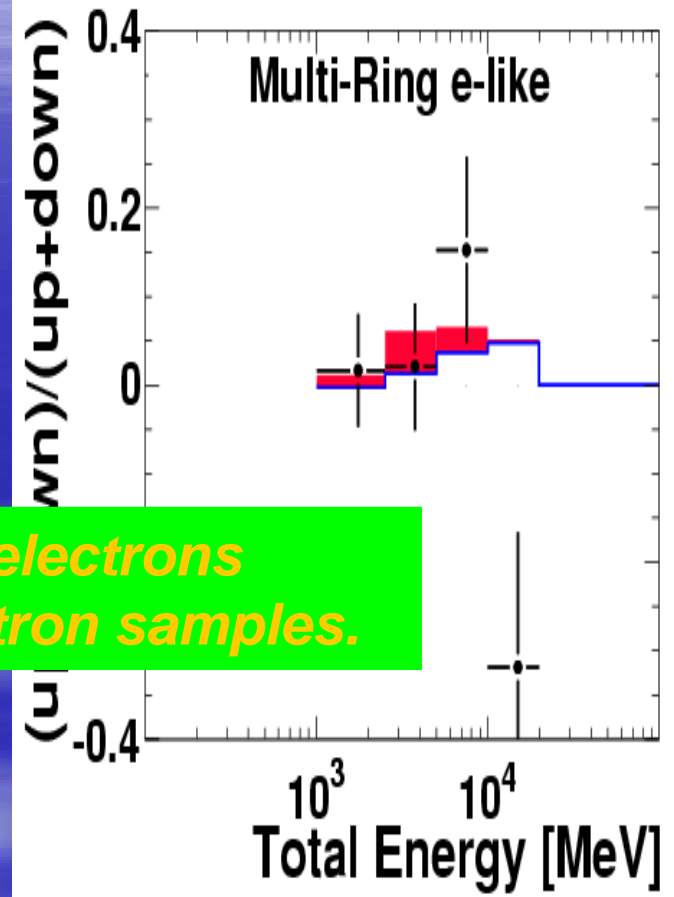
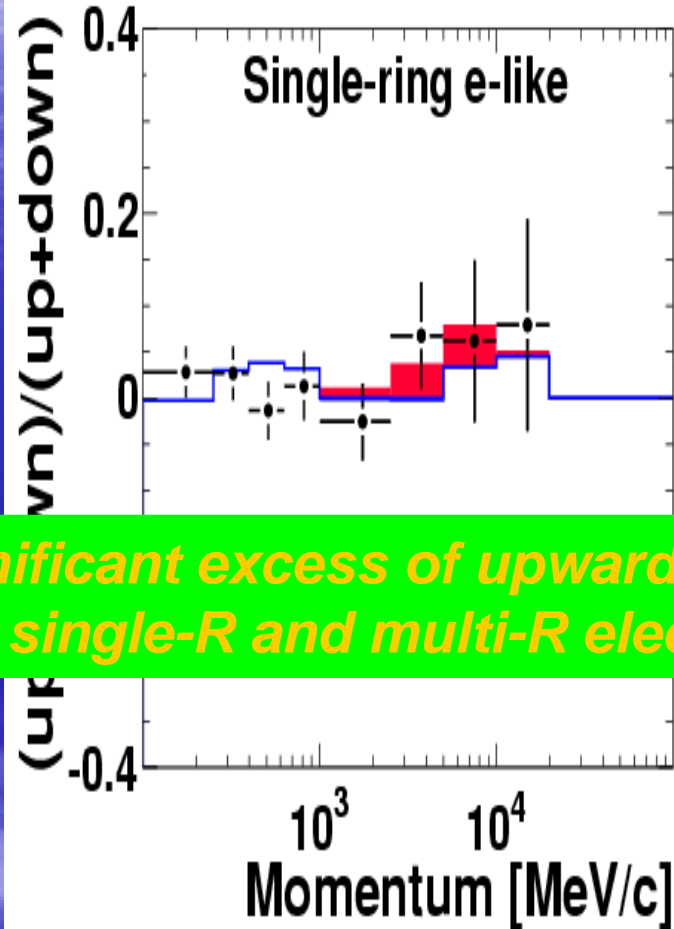
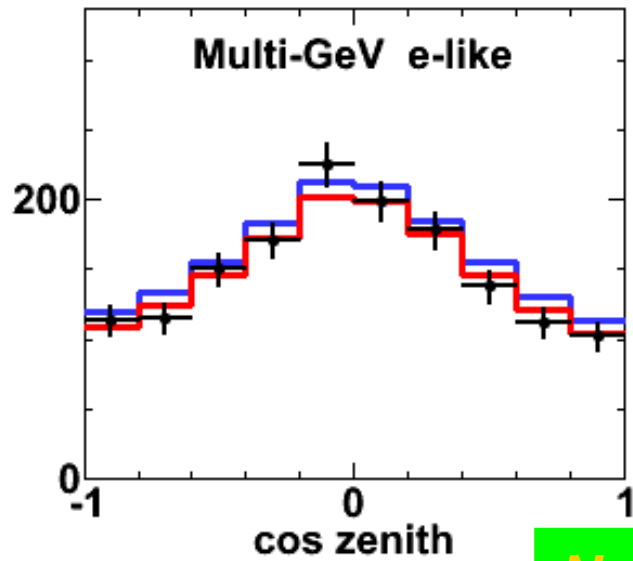


- Sub-GeV samples are further sub-divided
 - by decay-e and π^0 -likelihood
 - improve sensitivity to oscillation effect in low energy

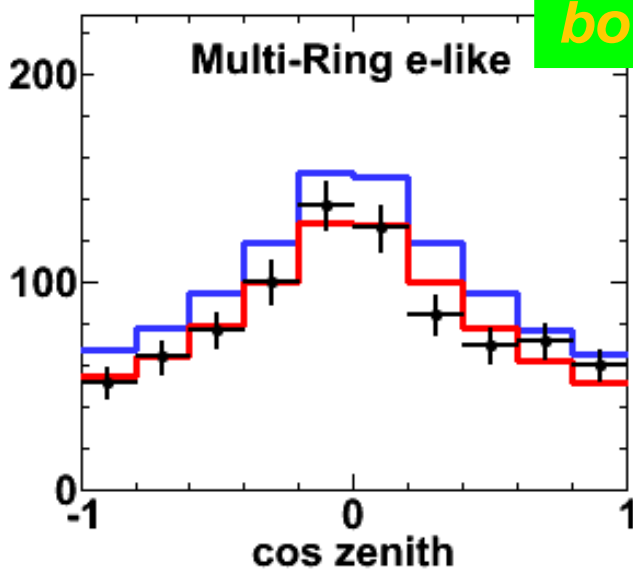
e-like	FC sub-GeV single-ring e-like			FC sub-GeV single-ring e-like
	0-DKe	1-DKe	π^0 -like	
number of MC events	2663.2	210.9	191.8	2996.4
Q.E.	77.7 %	3.8 %	10.6 %	70.6 %
CC single meson	12.4 %	50.3 %	7.0 %	15.2 %
$\nu_e + \bar{\nu}_e$ multi π	1.0 %	9.7 %	1.8 %	1.7 %
coherent π	1.3 %	8.5 %	0.5 %	1.7 %
CC $\nu_\mu + \bar{\nu}_\mu$	0.6 %	15.2 %	7.0 %	2.0 %
NC	6.8 %	11.2 %	72.0 %	8.7 %

mu-like	FC sub-GeV single-ring mu-like			FC sub-GeV single-ring mu-like
	0-DK μ	1-DK μ	2-DK μ	
number of MC events	1412.4	2745.4	164.3	4297.8
Q.E.	71.3 %	78.5 %	5.8 %	74.7 %
CC single meson	12.9 %	15.5 %	65.7 %	16.7 %
$\nu_\mu + \bar{\nu}_\mu$ multi π	1.1 %	1.5 %	14.9 %	1.9 %
coherent π	0.8 %	1.5 %	8.6 %	1.6 %
CC $\nu_e + \bar{\nu}_e$	1.8 %	<0.1 %	<0.1 %	0.7 %
NC	11.8 %	2.6 %	3.3 %	4.3 %

θ_{13} search (SK-I +II+III)

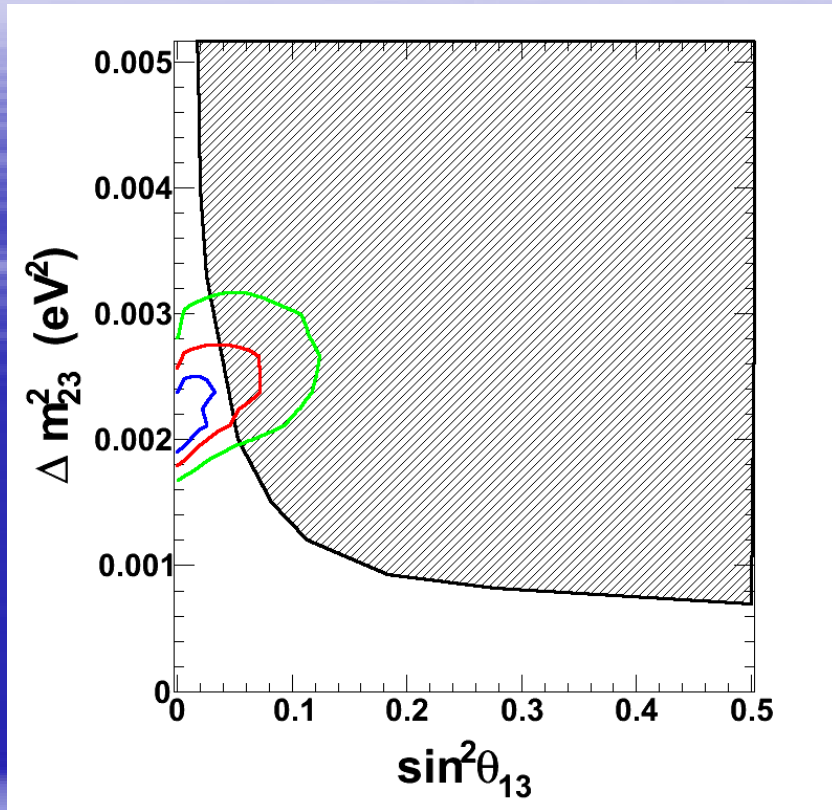


No significant excess of upward electrons both in single-R and multi-R electron samples.

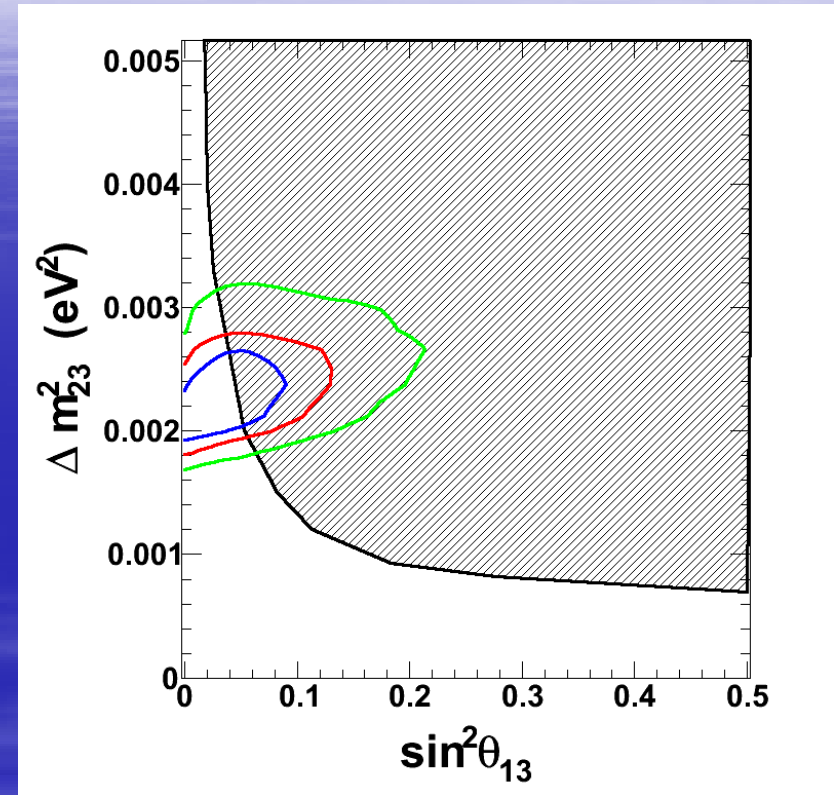


Current constraint on θ_{13} (SK-I +II+III)

$\Delta m^2 > 0$



$\Delta m^2 < 0$



No evidence for nonzero θ_{13} so far

CHOOZ reactor: $\sin^2\theta_{13} < 0.04$ @ 90%CL

SK atm- ν : $\sin^2\theta_{13} < 0.066$ (normal hierarchy)
 < 0.131 (inverted hierarchy)

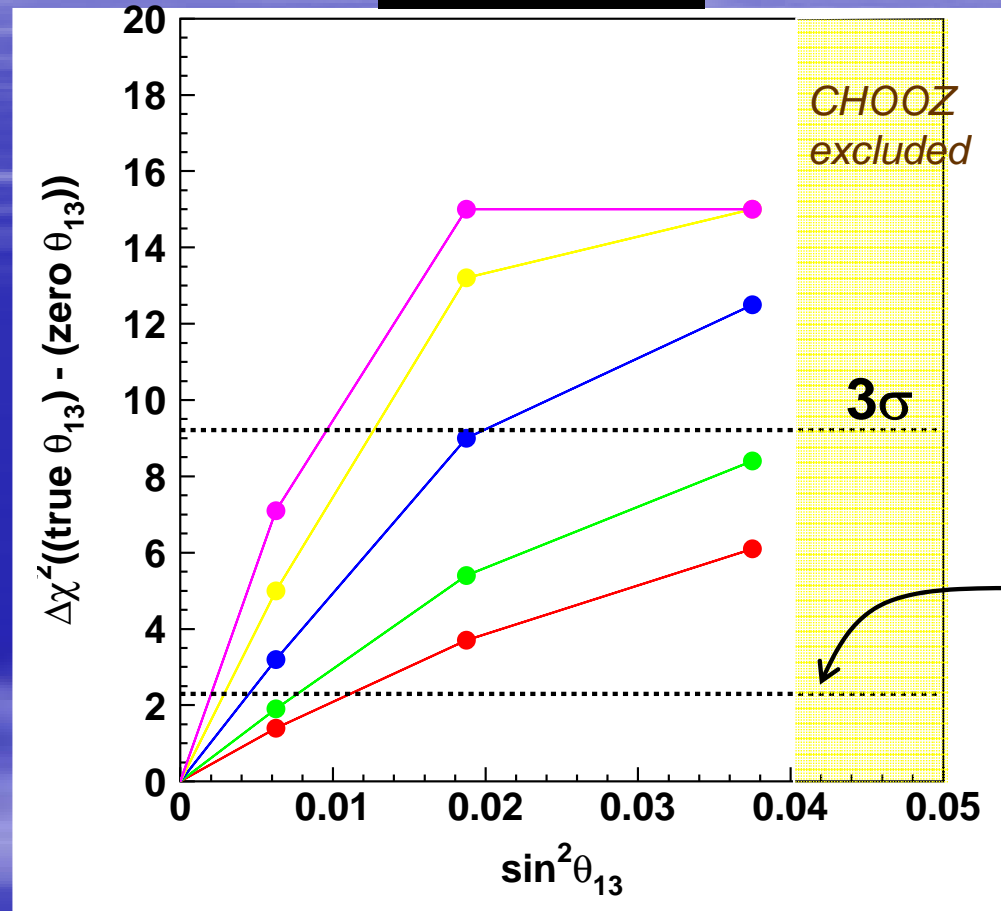
LBL K2K: $\sin^2\theta_{13} < 0.075$ PRL93(04)

MINOS: $\sin^2\theta_{13} < 0.073$ PRL101(08)

Significance for nonzero θ_{13}

$s^2_{21}\theta_{12}=0.825$
 $s^2_{23}\theta_{23}=0.40 \sim 0.60$
 $s^2_{13}\theta_{13}=0.00\sim 0.04$
 $\delta_{cp}=45^\circ$
 $\Delta m^2_{12}=8.3e-5$
 $\Delta m^2_{23}=2.5e-3$

20yrs SK



$s^2_{23}=0.40$
 0.45
 0.50
 0.55
 0.60

\mathcal{G}_{13} resonance
 $\propto \sin^2 \mathcal{G}_{13}^m (r \cdot \sin^2 \mathcal{G}_{23} - 1)$

**3σ for 80yrs SK
 ~4yrs HK**

Positive signal for nonzero θ_{13} can be seen if θ_{13} is near the CHOOZ limit and $s^2_{23} > 0.5$

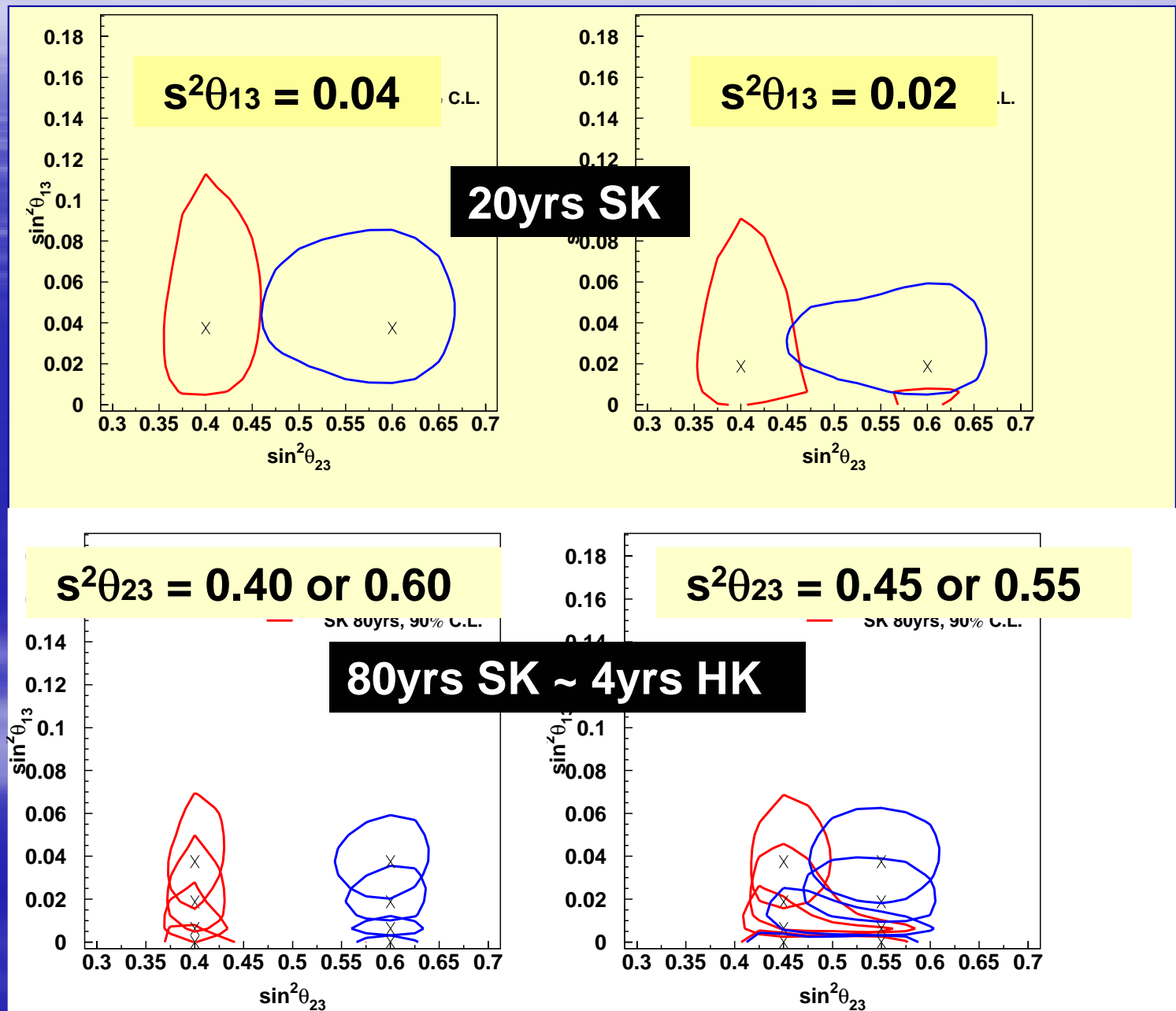
Discrimination of θ_{23} octant

SK-I+II+III
 $0.441 < \sin^2 \theta_{23} < 0.561$ (68%)

$s^2 \theta_{12} = 0.825$
 $s^2 \theta_{23} = 0.4$ or 0.6
 $s^2 \theta_{13} = 0.00 \sim 0.04$
 $\delta_{CP} = 45^\circ$
 $\Delta m^2_{12} = 8.3e-5$
 $\Delta m^2_{23} = 2.5e-3$

With 20yrs SK,
 discrimination
 may possible for
 $\sin^2 \theta_{13} > 0.02$

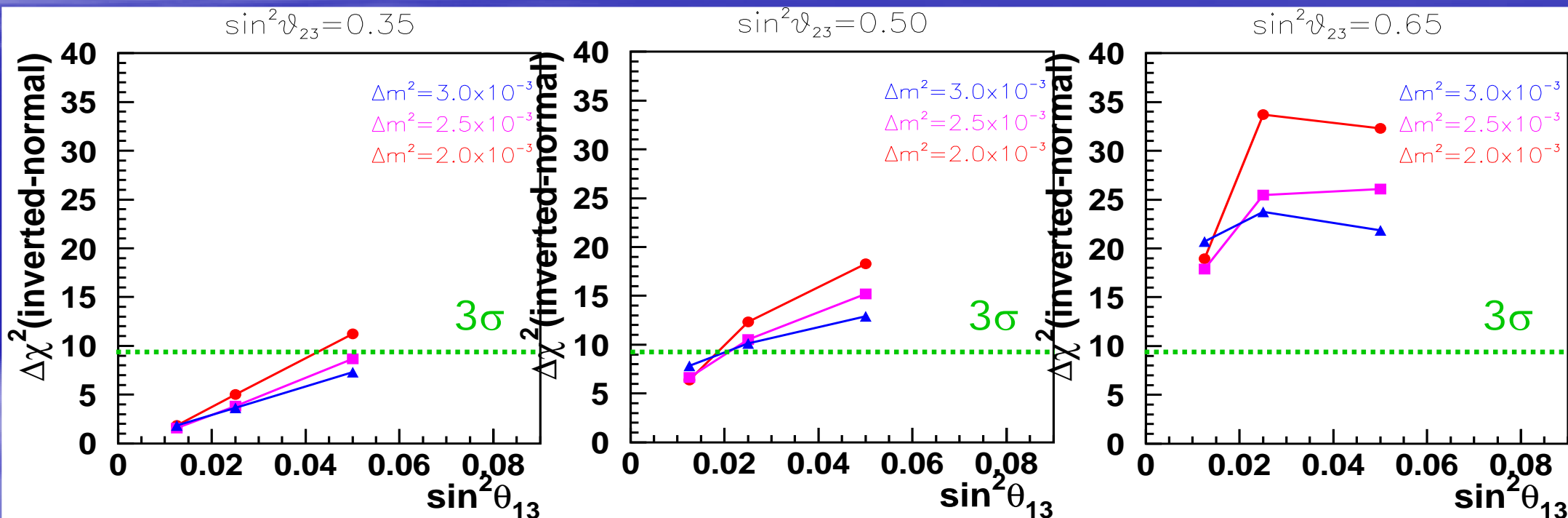
More statistics
 will increase
 power.



Discrimination of mass hierarchy (1)

1.8 Mton·yr = 3.3yr HK

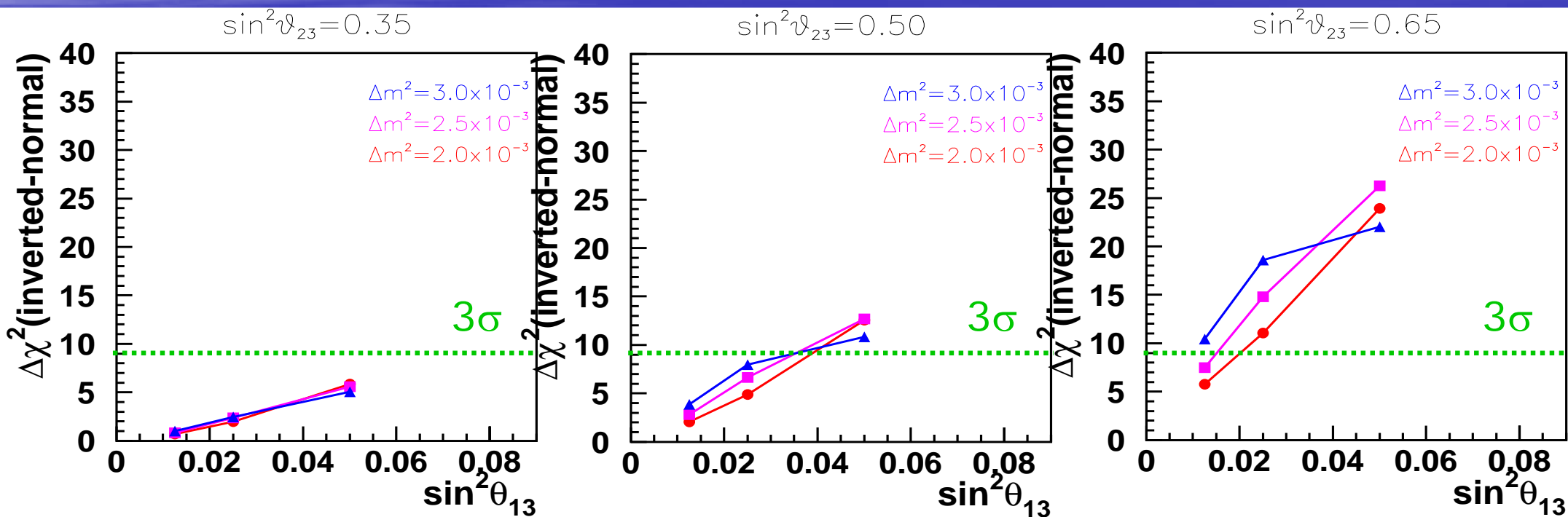
True : normal mass hierarchy



Discrimination of mass hierarchy (2)

1.8 Mton·yr = 3.3yr HK

True : inverted mass hierarchy

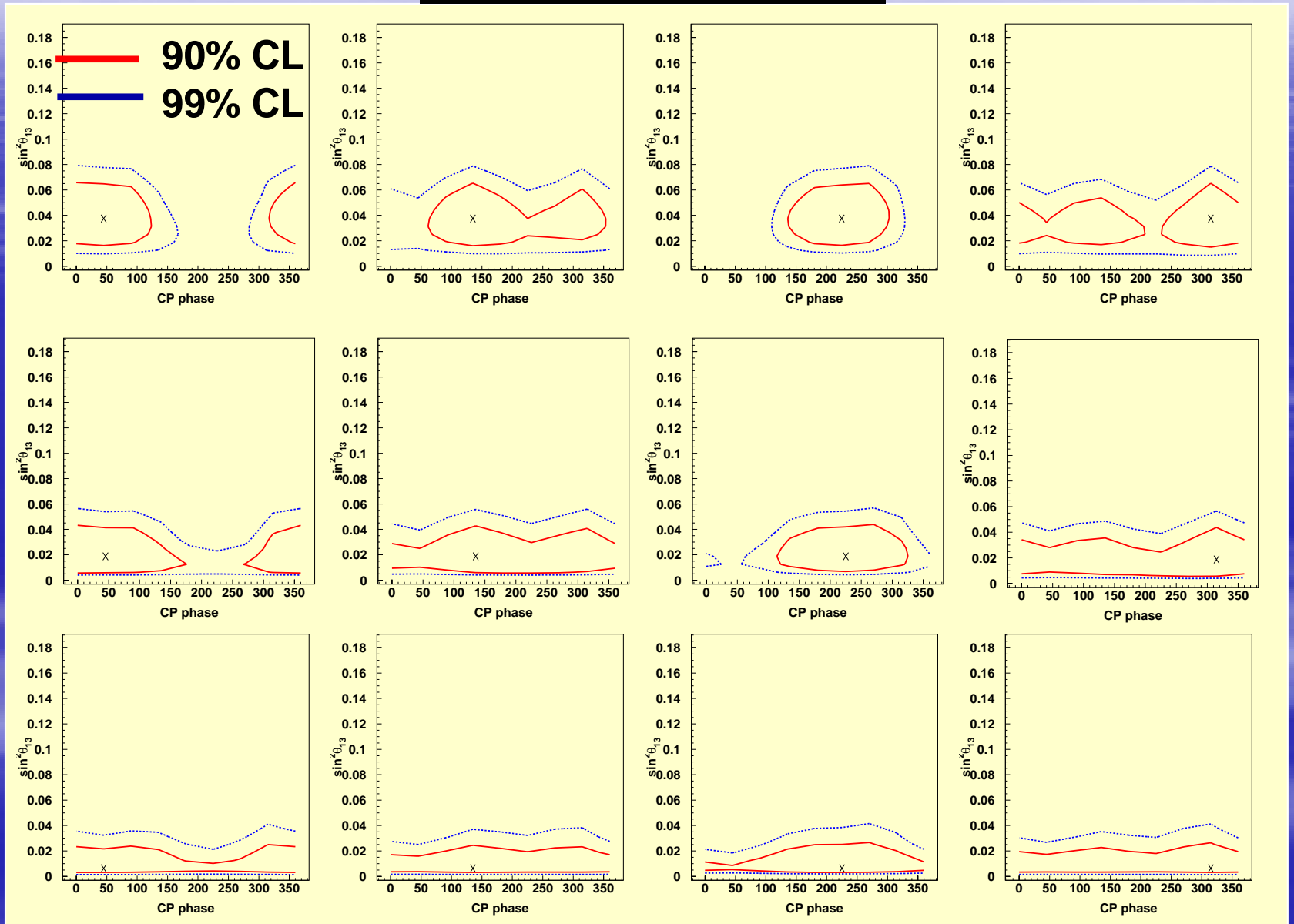


δ_{CP} sensitivity

80yrs SK ~ 4yrs HK

$s^2 2\theta_{12} = 0.825$
 $s^2 \theta_{23} = 0.5$
 $s^2 \theta_{13} = 0.00 \sim 0.04$
 $\delta_{CP} = 0^\circ \sim 360^\circ$
 $\Delta m^2_{12} = 8.3e-5$
 $\Delta m^2_{23} = 2.5e-3$

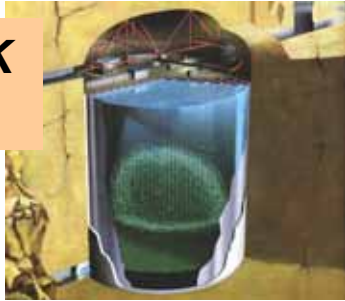
$\sin^2 \theta_{13}$ ↑
CP phase →



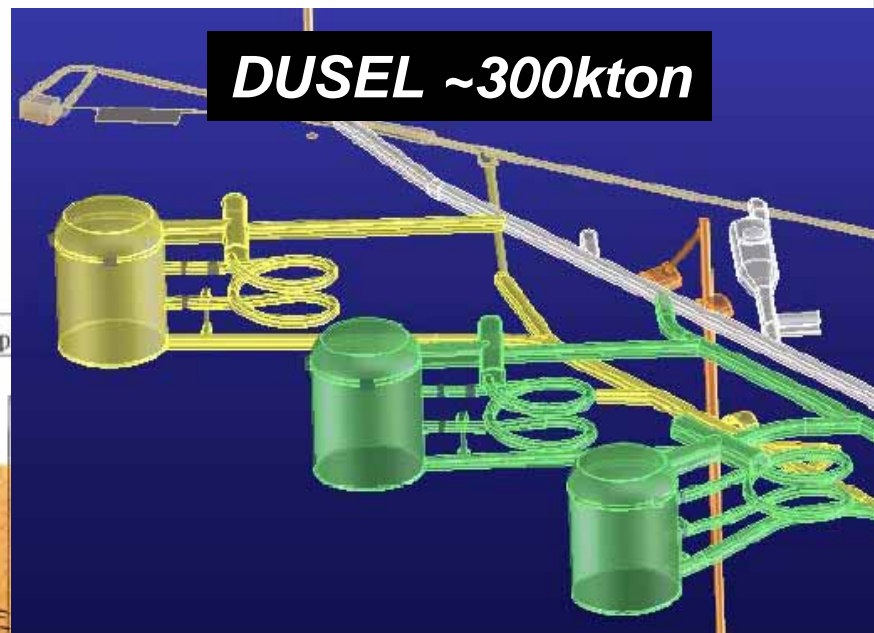
CP phase could be seen if θ_{13} is close to the CHOOZ limit.

Megaton scale water Cherenkov detectors

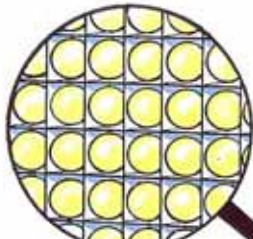
**Super-K
22kton**



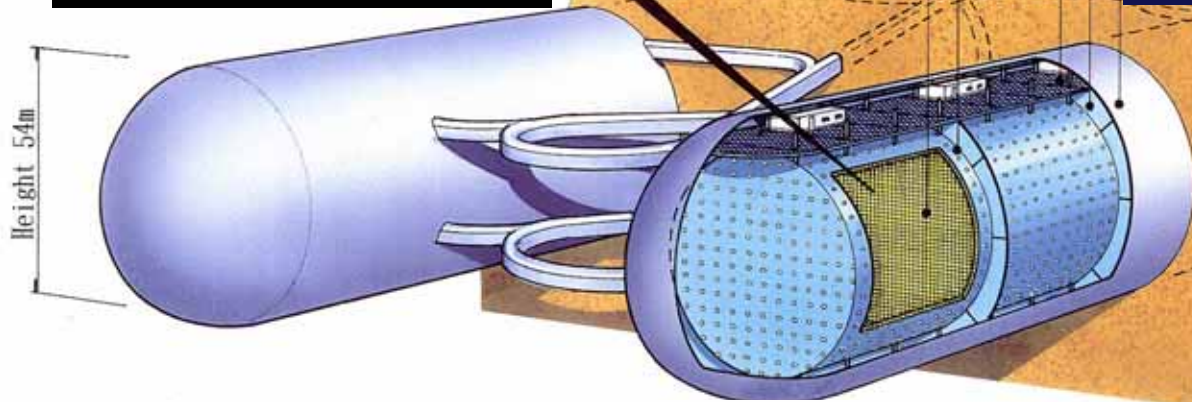
DUSEL ~300kton



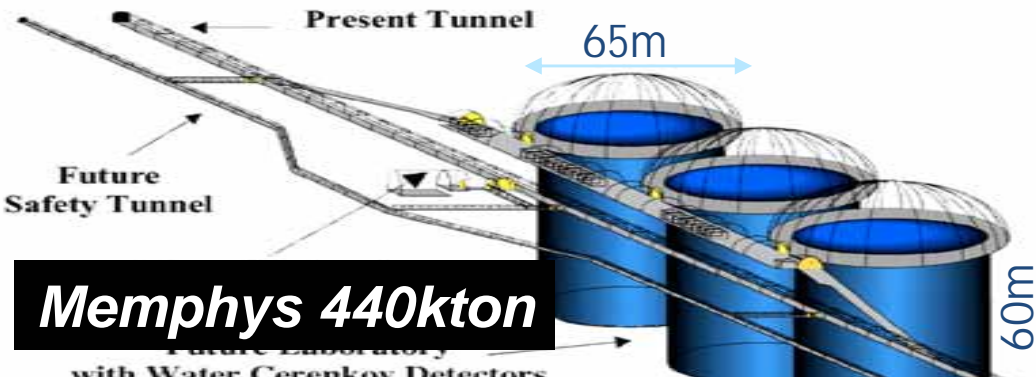
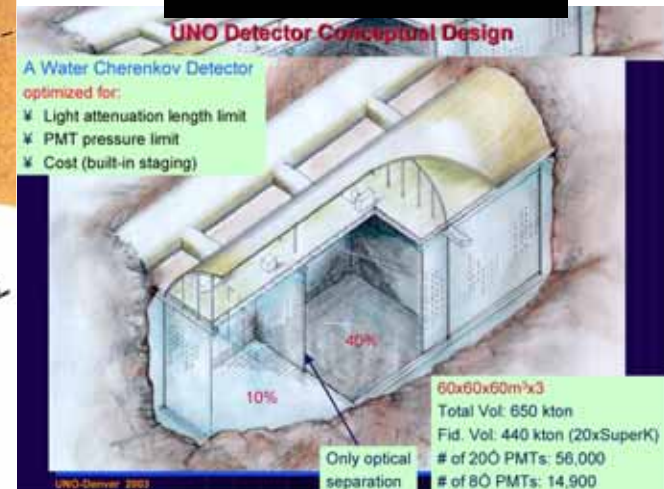
Hyper-K 540kton



Outer Detector
Inner Detector
Access Drift



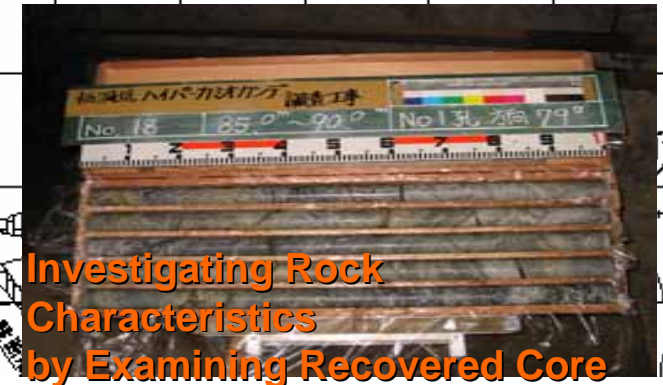
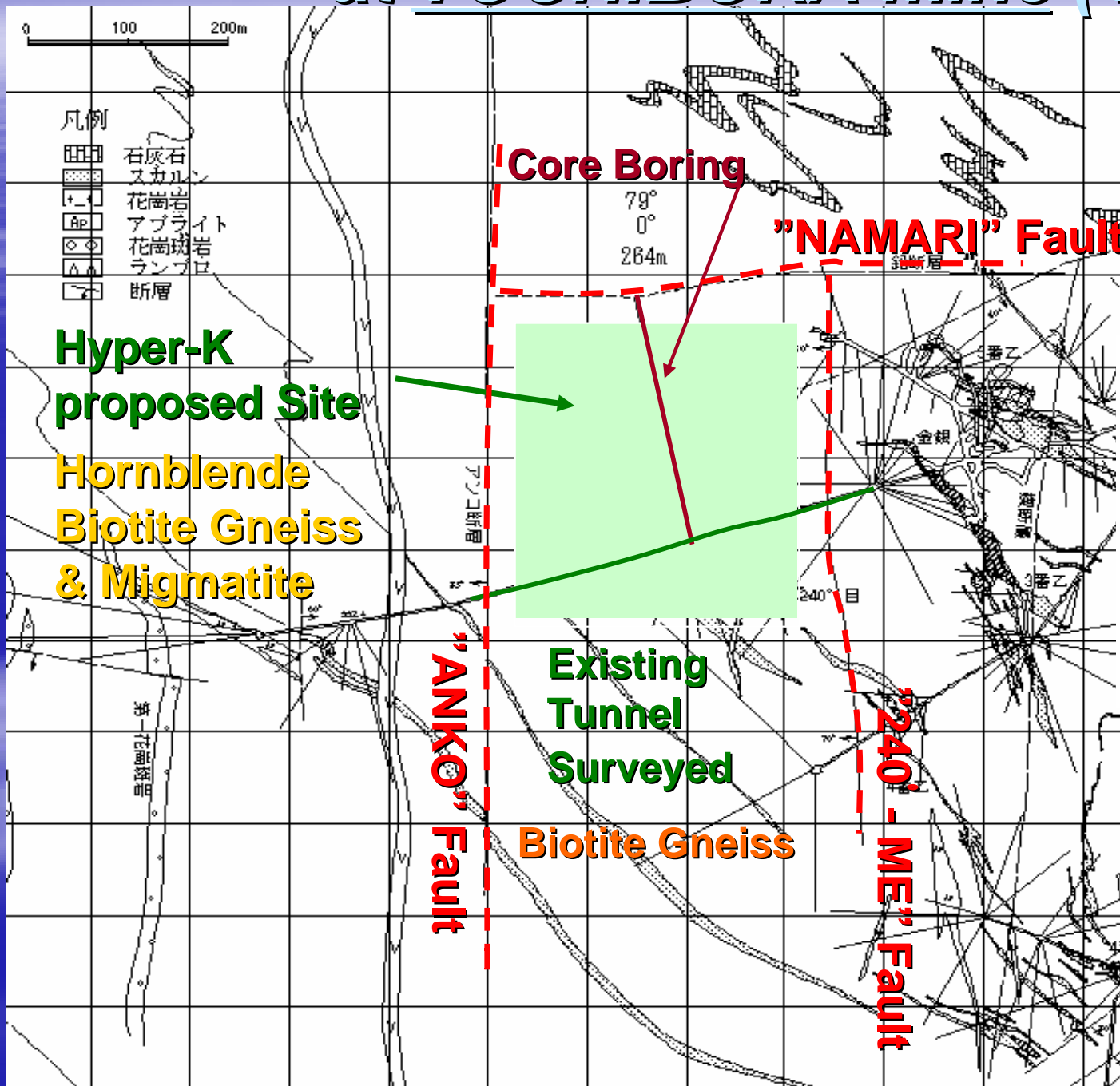
UNO 440kton



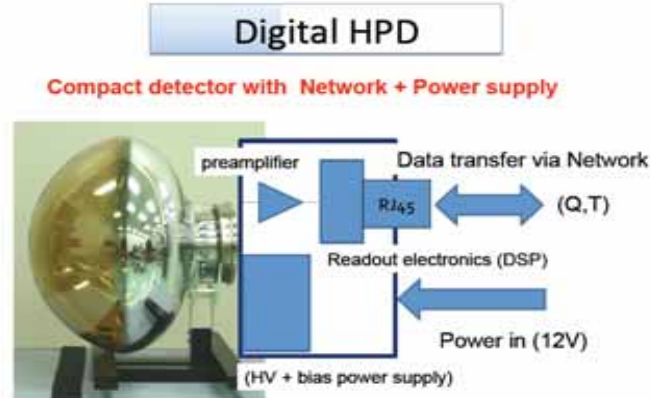
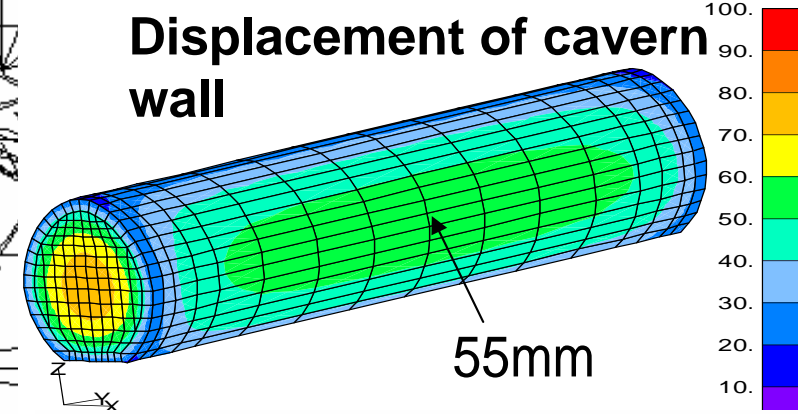
Memphys 440kton

Future Laboratory with Water Cherenkov Detectors

Geological Map of Proposed Site at TOCHIBORA mine (+ 550mEL)



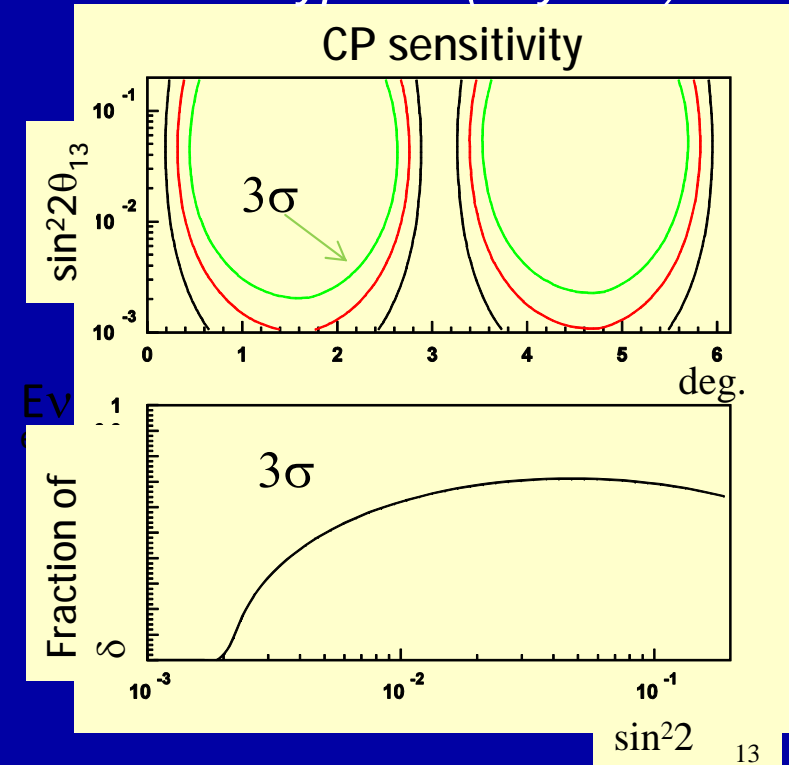
Investigating Rock Characteristics by Examining Recovered Core



Physics goals

- CPV with accelerator ν (LBLE)
- proton decay searches
 - $\sim 10^{35}$ years for $p \rightarrow e^+ \pi^0$
- precise meas. of atmospheric ν
 - δ , θ_{13} , mass hierarchy ($\sin^2 \theta_{13} > \sim 0.01$)
 - θ_{23} octant
- supernova ν
 - mechanism of stellar collapse
 - mass hierarchy?
- solar ν
 - day-night flux (matter effect)
 - Hep ν

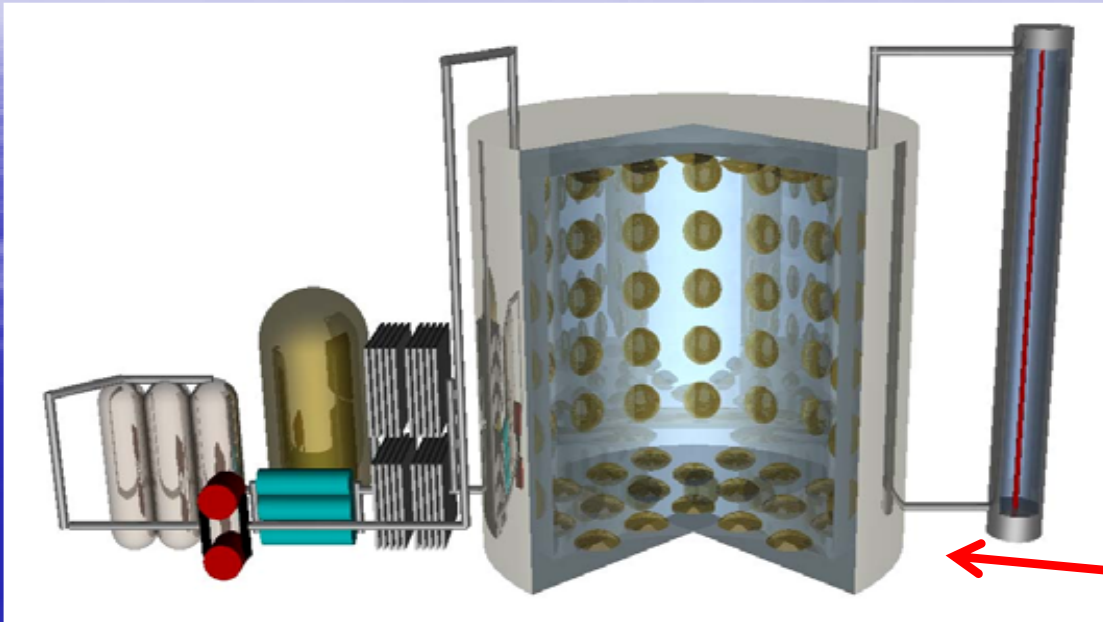
JPARC upgraded 1.66MW beam
+ 540kton Hyper-K (10years)



Summary

- Super-K is now back on solar ν business
 - Aiming to reach 4 MeV (new electronics, lower BG, better reconstruction, smaller systematics)
- Atm- ν is providing information on $\nu_{\mu} \rightarrow \nu_{\tau}$
 - $\Delta m_{23}^2 = 2.19^{+0.14}_{-0.13} \times 10^{-3} \text{ eV}^2$ (68% by L/E study)
 - $\sin^2 2\theta_{23} > 0.96$ @ 90% CL
 - 2.4 σ level ν_{τ} significance
- LBLE era
 - MINOS, T2K, Nova, LBLE with Mton-size...
 - Atm- ν could also give us information on θ_{13} , octant of θ_{23} , mass hierarchy, and CP phase in future (especially in the case of $\sin^2 \theta_{13} > \sim 0.01$)

Gadolinium test tank (under construction)



- Hall excavation (Sep.-2009 to Dec.)
- 200ton tank with PMTs
- Test the effect of Gd on light attenuation length, purification method, corrosion of detector materials...

