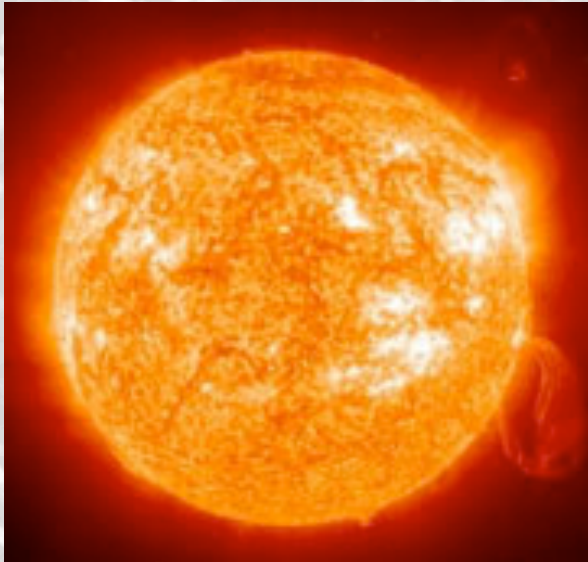




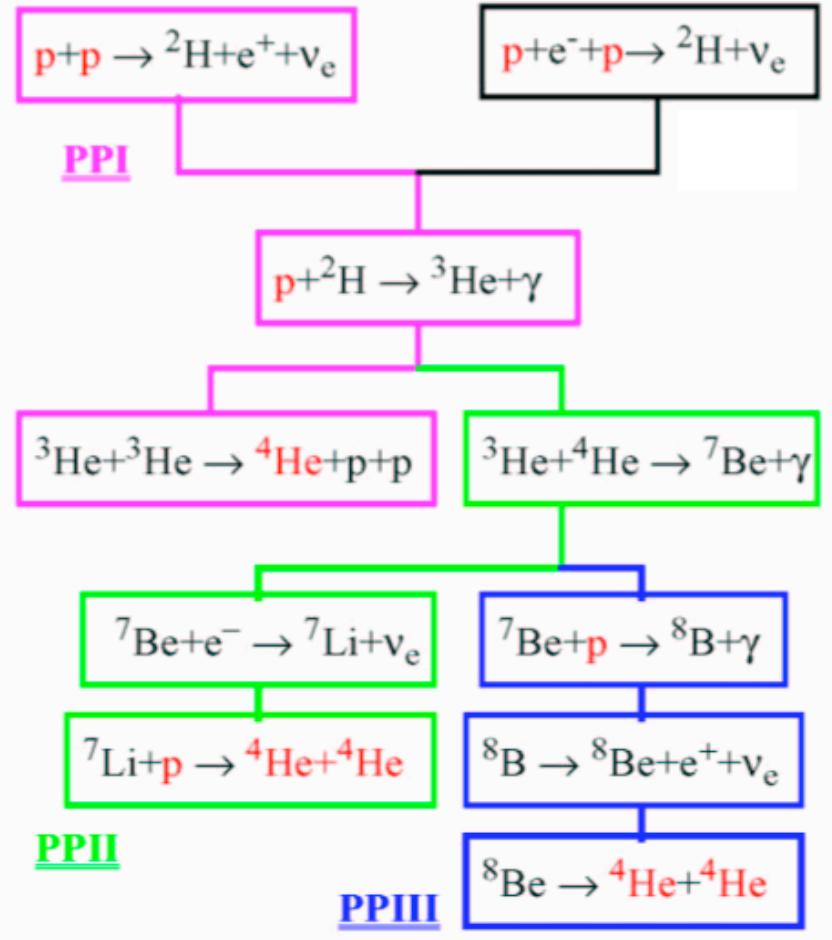
# **Experimental evidence of electron neutrino oscillations and validation of MSW-LMA model with Borexino**

**Margherita Buizza Avanzini**  
for the Borexino Coll.

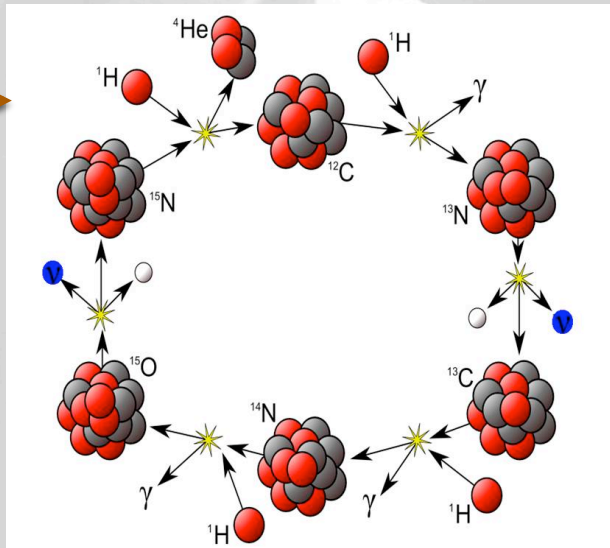
# BEFORE SOLAR NEUTRINOS... THE SUN BURNING



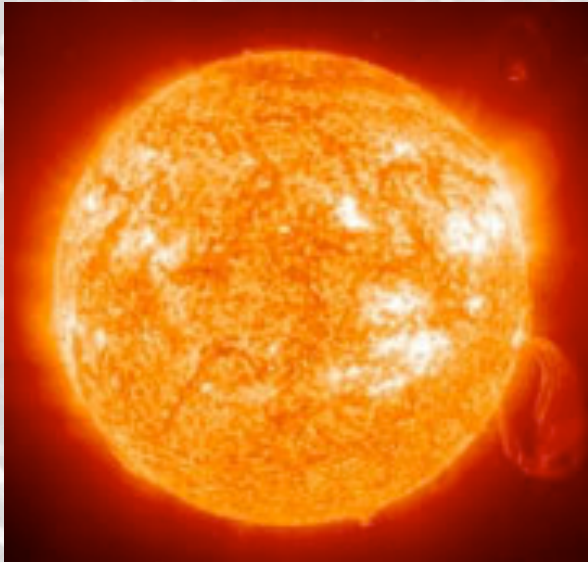
**98% of energy  
through the pp  
chain**



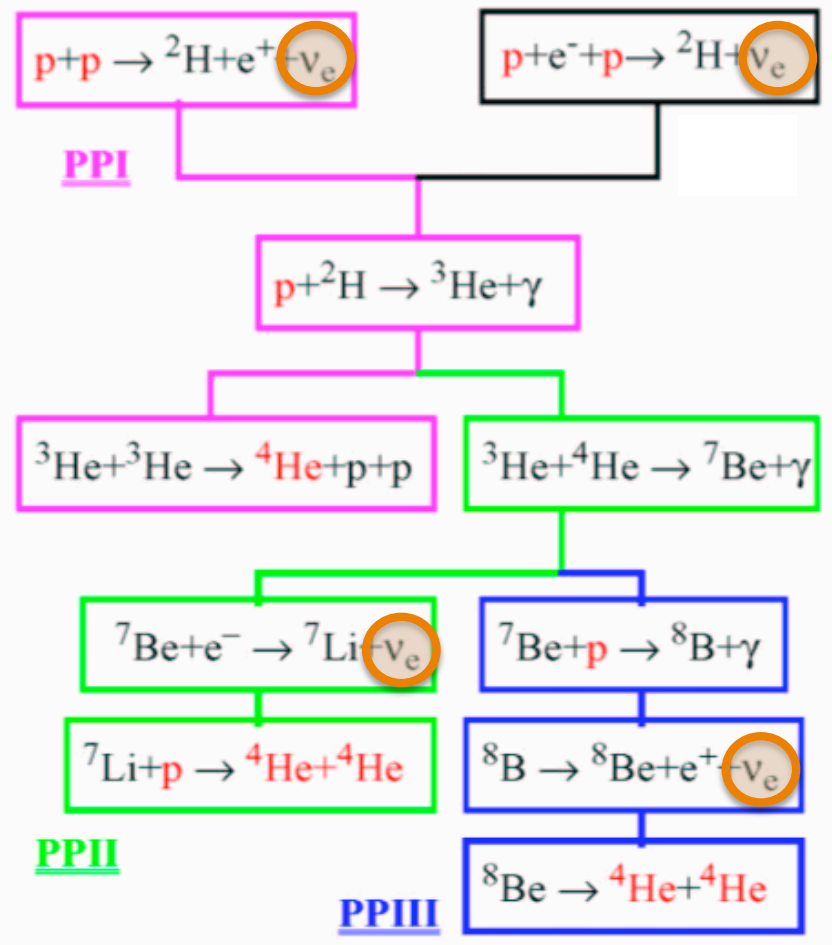
**CNO? Just  
2% of  
energy...**



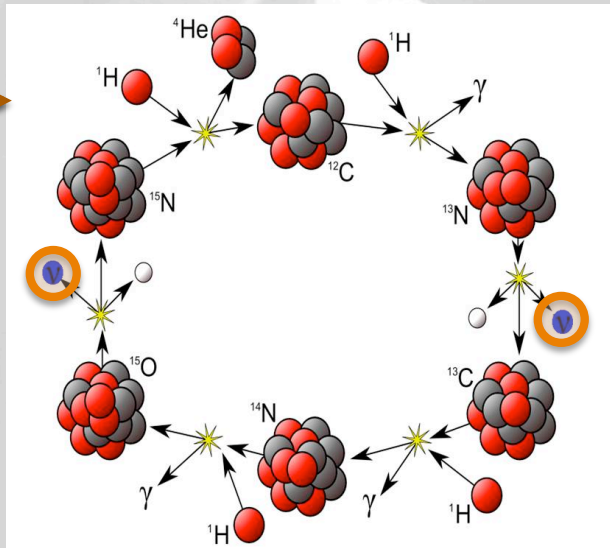
# BEFORE SOLAR NEUTRINOS... THE SUN BURNING



**98% of energy  
through the pp  
chain**



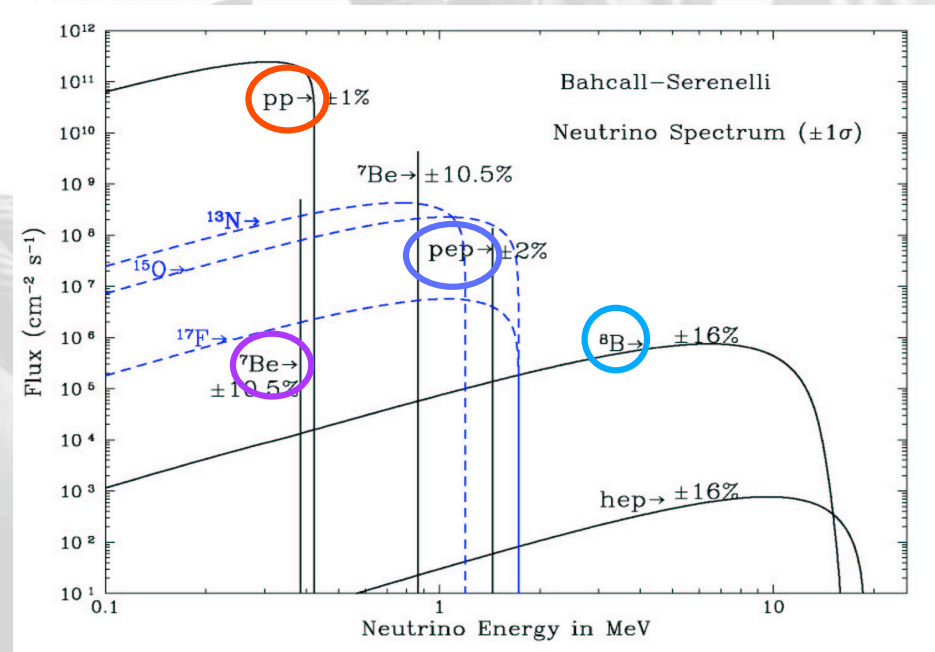
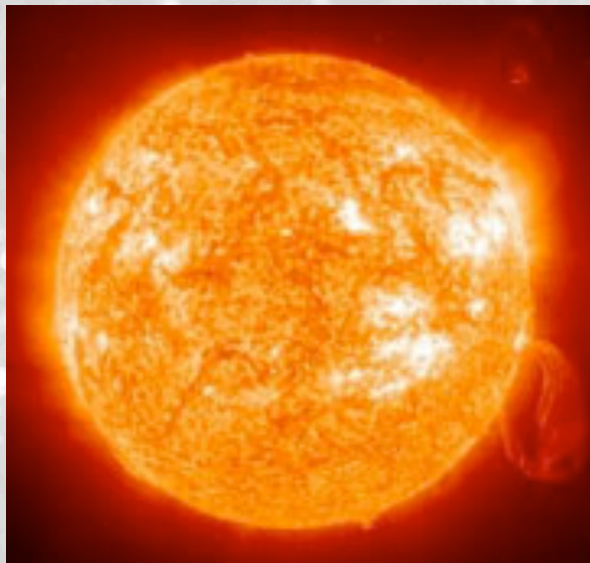
**CNO? Just  
2% of  
energy...**



But what about neutrinos?!

# SOLAR NEUTRINOS

Standard Solar Model (BPS09) predicts **fluxes** and **spectra** of  $\nu_e$ :

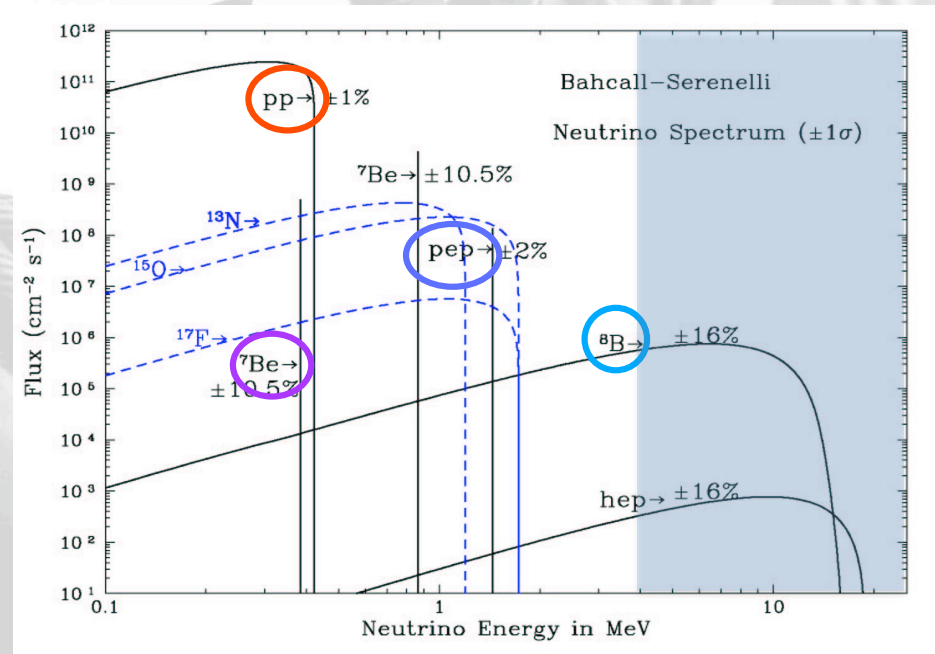
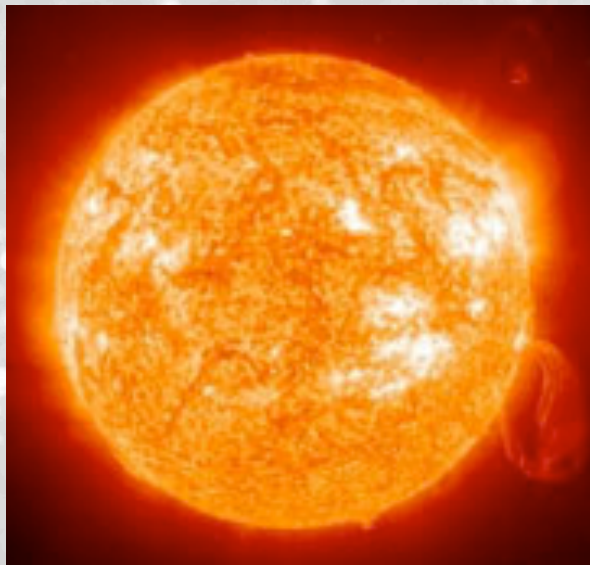


Real time Cherenkov experiments can investigate only the final part of  $^8\text{B}$  spectrum.

Solar neutrino experiments helped in clarifying the neutrino physics:  
**neutrino oscillations.**

# SOLAR NEUTRINOS

Standard Solar Model (BPS09) predicts **fluxes** and **spectra** of  $\nu_e$ :

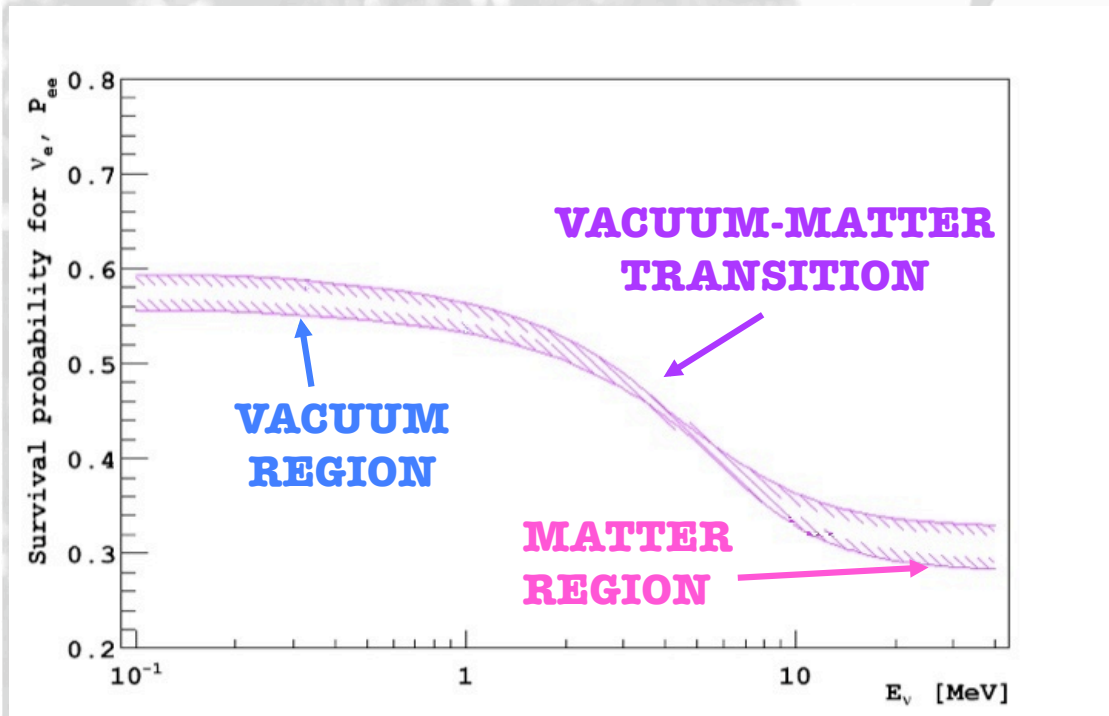


Real time Cherenkov experiments can investigate only the final part of  $^8\text{B}$  spectrum.

Solar neutrino experiments helped in clarifying the neutrino physics:  
**neutrino oscillations.**

# THE MSW-LMA SOLUTION

The **survival probability** of  $\nu_e$  for the MSW-LMA solution:



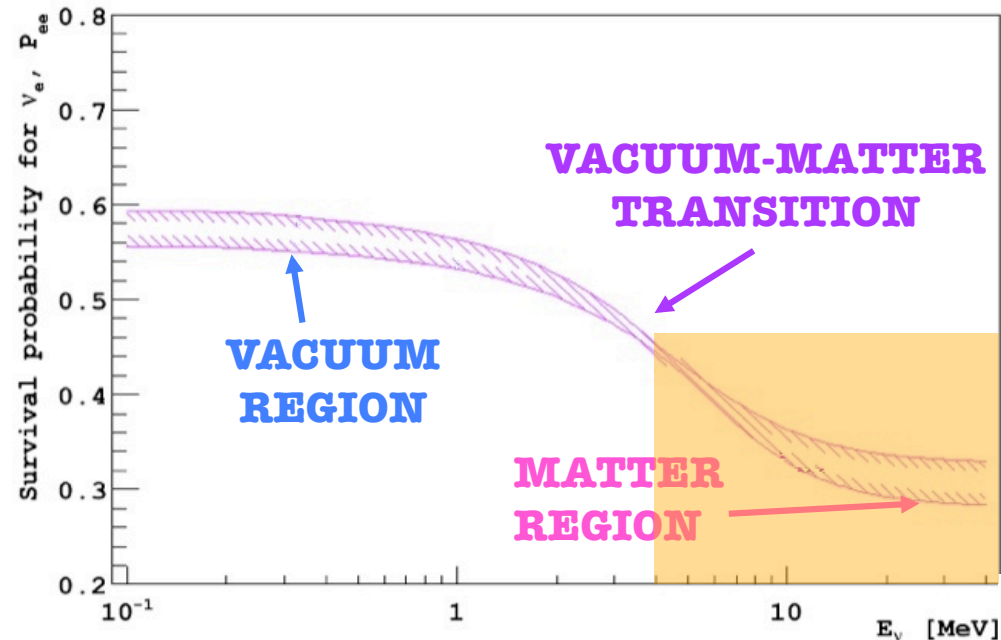
**VACUUM REGION:** low energy, oscillations like in the vacuum.

**VACUUM-MATTER TRANSITION:** in between

**MATTER REGION:** high energy, oscillation affected by the matter effect (**MSW effect**)

# THE MSW-LMA SOLUTION

The **survival probability** of  $\nu_e$  for the MSW-LMA solution:



**VACUUM REGION:** low energy, oscillations like in the vacuum.

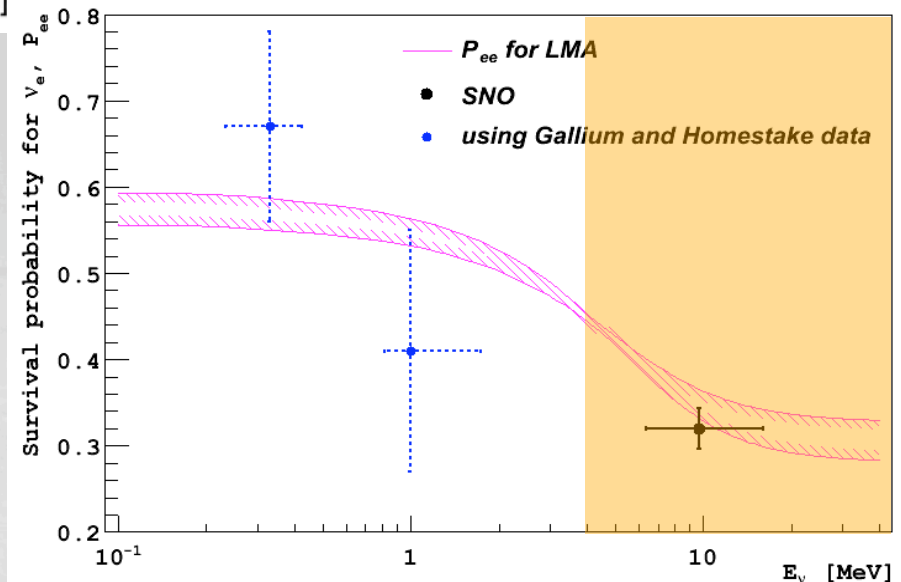
**VACUUM-MATTER TRANSITION:** in between

**MATTER REGION:** high energy, oscillation affected by the matter effect (**MSW effect**)

## WITHOUT BOREXINO

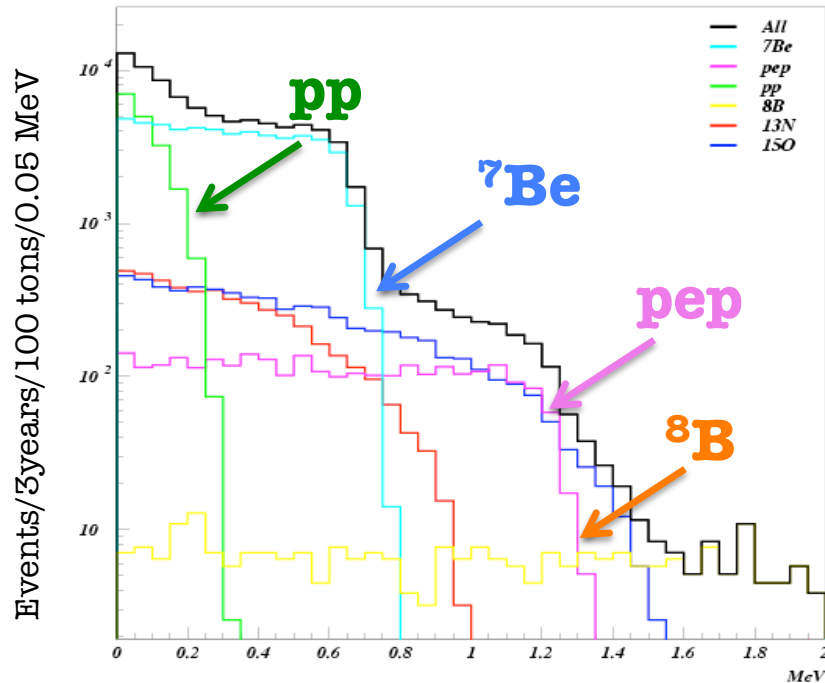
**$^8\text{B}$** : real time experiment with energy threshold of detected electron at about **4 MeV** ( $\approx 10^{-4}$  of the total flux)  $\rightarrow$  **matter region**.

**What new with Borexino???**



# THE BOREXINO EXPERIMENT

## MONTECARLO SPECTRA in BX



First real time experiment for **low energy solar neutrinos!**

✓ Main goal:  ${}^7\text{Be}$  flux measurement

✓ But also: measure of  ${}^8\text{B}$  neutrinos!!!

✓ And for the future... pep and pp

For the first time in real time:

- **Test of Standard Solar Model:** precise measurement of  ${}^7\text{Be}$   $\nu$  flux!
- **Test of MSW-LMA solution** at low energy!
- BOREXINO can measure ALSO  ${}^8\text{B}$  neutrinos!
- Energy threshold for  ${}^8\text{B}$  neutrinos down to **3 MeV** (electron energy)!







# THE BOREXINO DETECTOR

## Scintillator:

270 t PC+PPO in a 125  $\mu\text{m}$  thick nylon vessel

## Nylon vessels:

Inner: 4.25 m  
Outer: 5.50 m

Software cut at 3m,  
defining the Fiducial  
Volume (100tons)

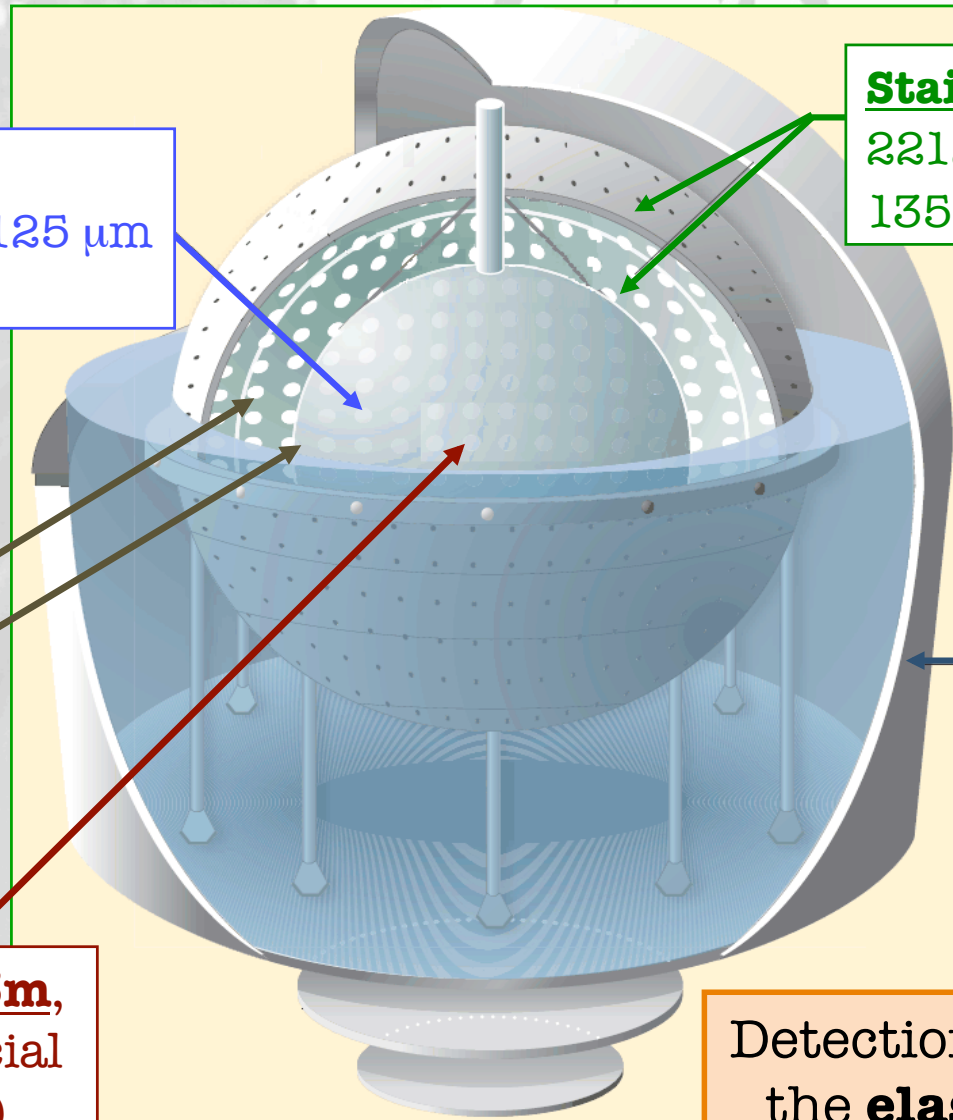
## Stainless Steel Sphere:

2212 photomultipliers  
1350  $\text{m}^3$

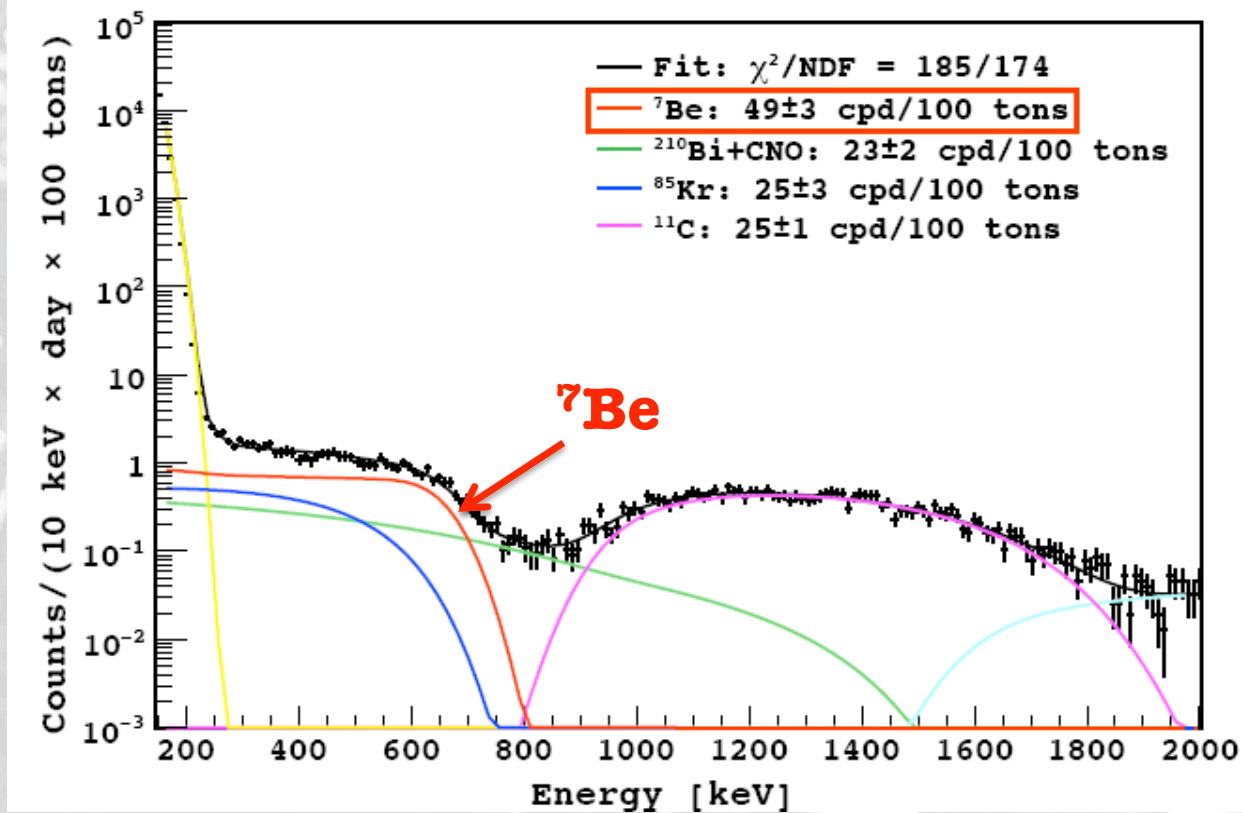
## Water Tank:

$\gamma$  and  $n$  shield  
 $\mu$  water Ch detector  
208 PMTs in water  
2100  $\text{m}^3$

Detection of solar  $\nu$  through  
the **elastic scattering** on  
electrons in high purified  
scintillator (300tons).



# RESULTS ON THE ${}^7\text{Be}$ $\nu$ MEASUREMENT (192d)



- Fit in the region 0.2–2 MeV
- $\alpha$  removal
- ${}^{210}\text{Bi}$ ,  ${}^{11}\text{C}$  and  ${}^{85}\text{Kr}$  as free parameters
- Other neutrino fluxes are fixed

Borexino Collaboration  
PRL 101 (2008)

$$R_{7\text{Be}} = 49 \pm 3_{\text{stat}} \pm 4_{\text{sys}} \text{ c/d/100 tons}$$

# $^8\text{B}$ NEUTRINO FLUX MEASUREMENT (1)

## EXPECTED RATE SSM+MSW-LMA

Whole energy spectrum: ~ **0.5 c/d/100 tons**

Above **5 MeV**: ~ **0.14 c/d/100 tons**

BUT we can decrease the threshold down to **3 MeV !!!**

# $^8\text{B}$ NEUTRINO FLUX MEASUREMENT (1)

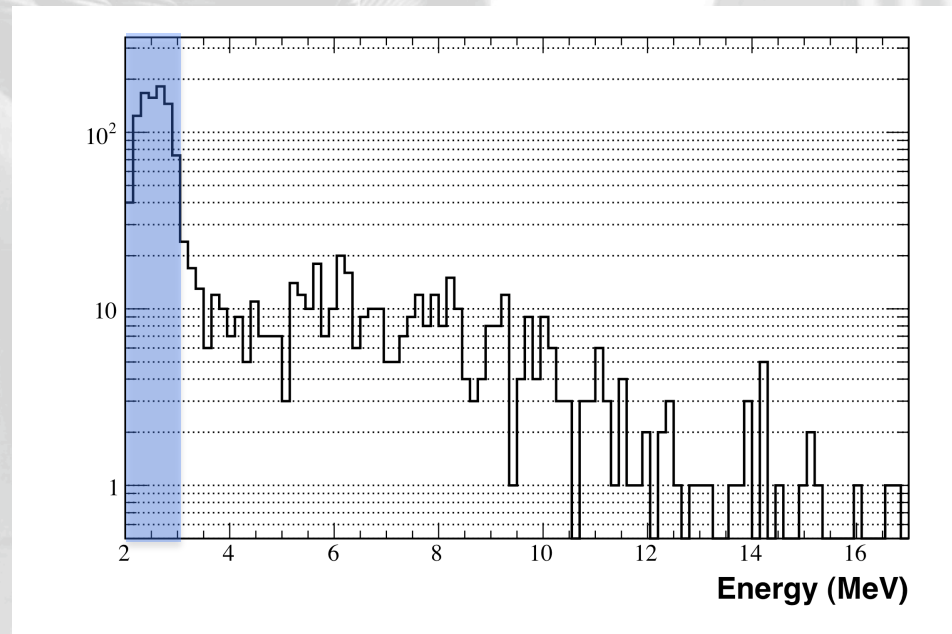
## EXPECTED RATE SSM+MSW-LMA

Whole energy spectrum:  $\sim 0.5$  c/d/100 tons

Above **5 MeV**:  $\sim 0.14$  c/d/100 tons

BUT we can decrease the threshold down to **3 MeV !!!**

Why the energy threshold at  
3MeV??



# $^8\text{B}$ NEUTRINO FLUX MEASUREMENT (1)

## EXPECTED RATE SSM+MSW-LMA

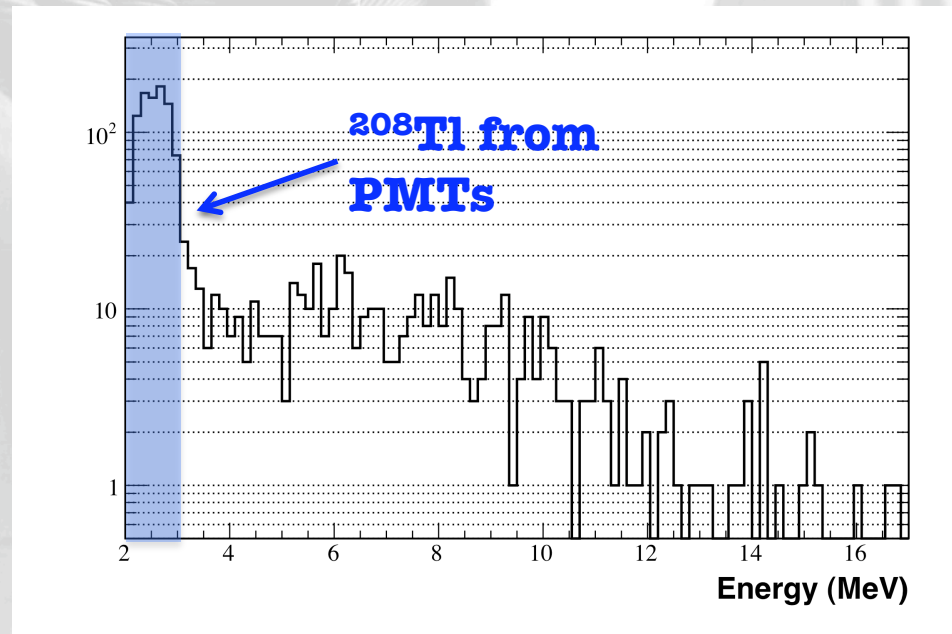
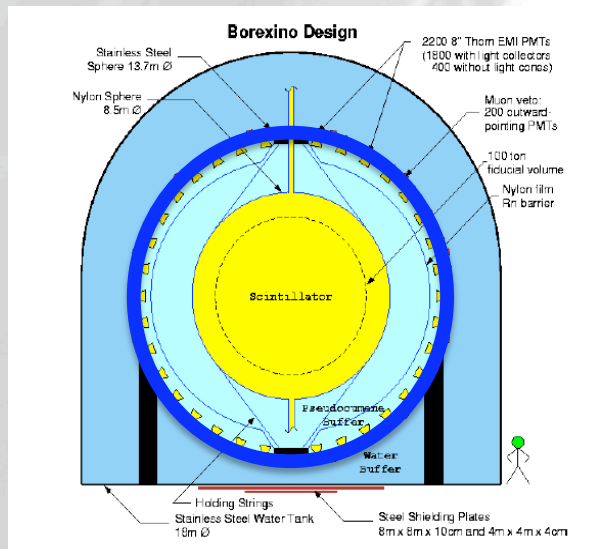
Whole energy spectrum:  $\sim 0.5$  c/d/100 tons

Above **5 MeV**:  $\sim 0.14$  c/d/100 tons

BUT we can decrease the threshold down to **3 MeV !!!**

Why the energy threshold at  
3MeV??

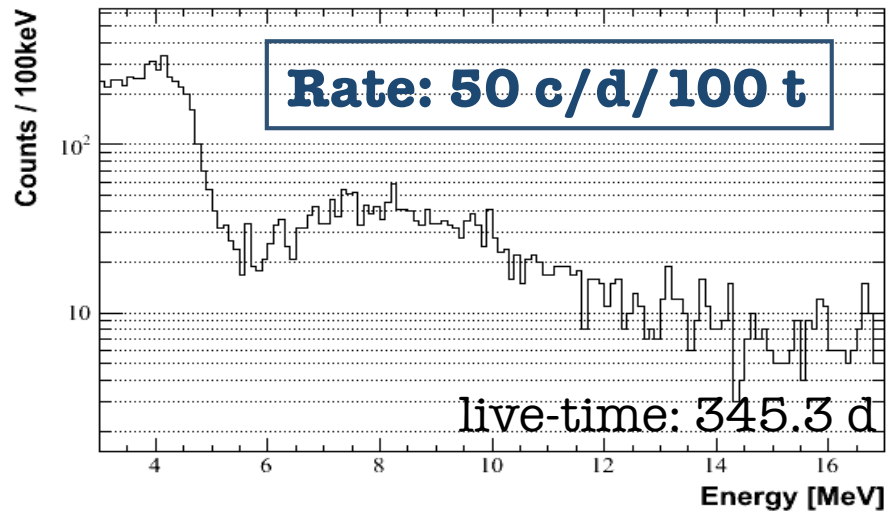
2.6MeV gammas from  $^{208}\text{Tl}$  coming  
from the PMTs reaching the FV



Setting the  
threshold at 3MeV

# BACKGROUND SOURCES (above 3MeV)

## DATA SPECTRUM BEFORE THE ANALYSIS

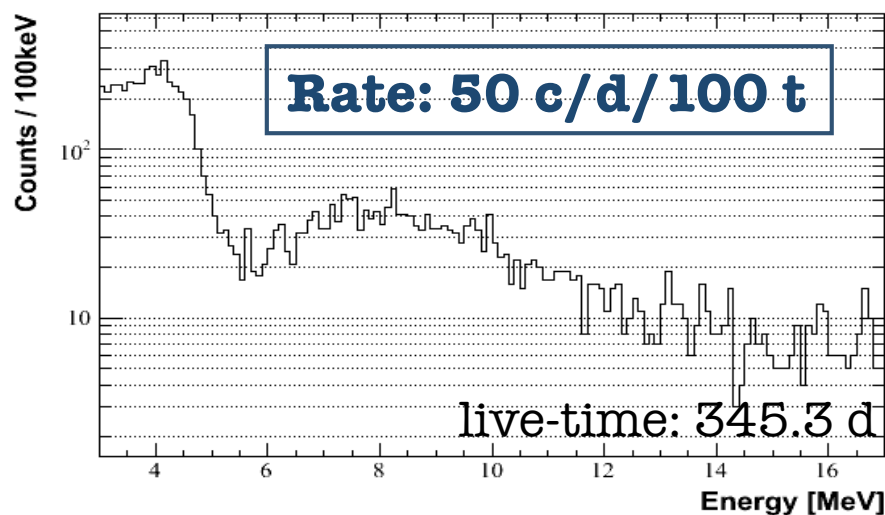


**Signal/Bkg  
ratio < 1/150 !!!**

**HARD GOAL!**

# BACKGROUND SOURCES (above 3MeV)

## DATA SPECTRUM BEFORE THE ANALYSIS



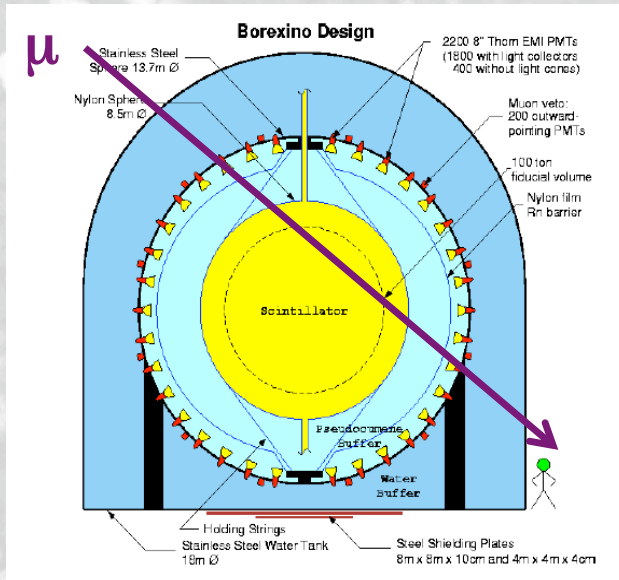
**Signal/Bkg  
ratio < 1/150 !!!**

**HARD GOAL!**

- **Cosmic muons**
- High energy  $\gamma$  (4.9 MeV) from **neutron capture** on  $^{12}\text{C}$
- **Cosmogenic isotopes**
- $^{214}\text{Bi}$  and  $^{208}\text{Tl}$  from **vessel contamination** of  $^{238}\text{U}$  and  $^{232}\text{Th}$
- $^{214}\text{Bi}$  and  $^{208}\text{Tl}$  from **internal contamination** of  $^{238}\text{U}$  and  $^{232}\text{Th}$

**Ad hoc technique for each kind of background !!!**

# BACKGROUND SOURCES (above 3MeV)

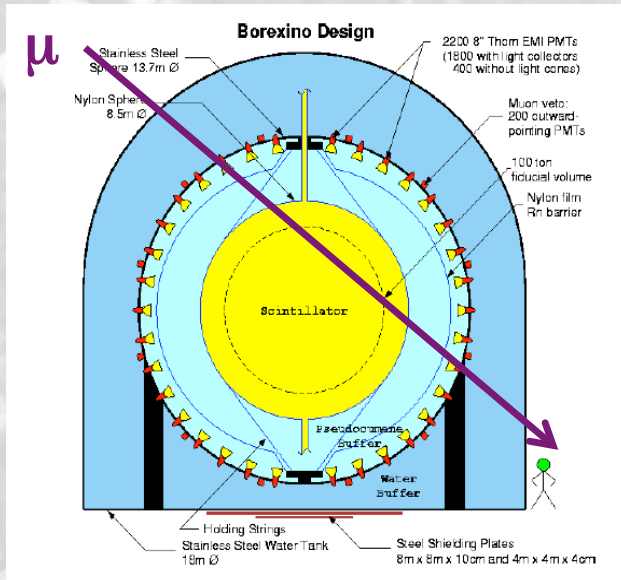


## MUONS AND COSMOGENIC NEUTRONS

Study of the pulse shape in the scintillator  $\neq$  point-like events!!  
+ rejection of gamma from n capture, in coincidence with the  $\mu$



# BACKGROUND SOURCES (above 3MeV)



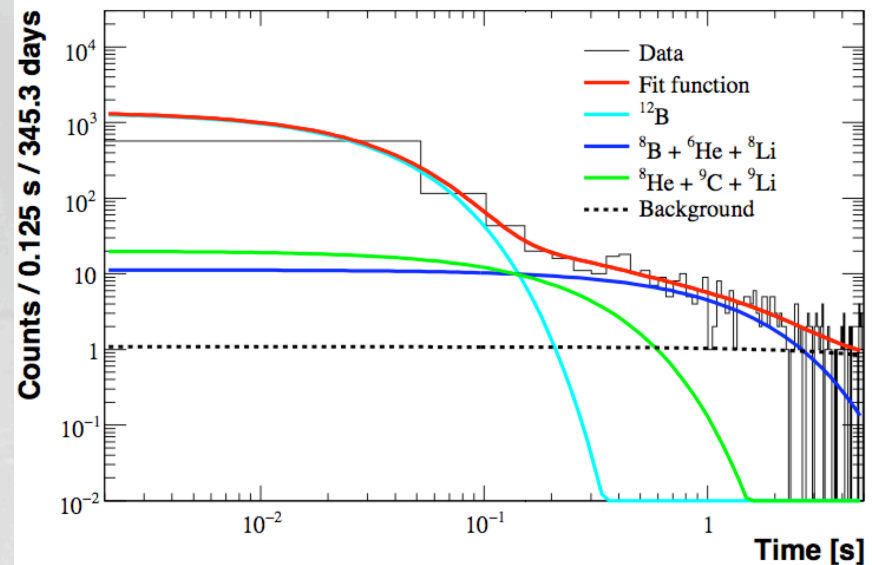
## MUONS AND COSMOGENIC NEUTRONS

Study of the pulse shape in the scintillator  $\neq$  point-like events!!  
+ rejection of gamma from n capture, in coincidence with the  $\mu$

## COSMOGENIC ISOTOPES

Study of cosmogenic isotopes through the coincidence with the father muon.

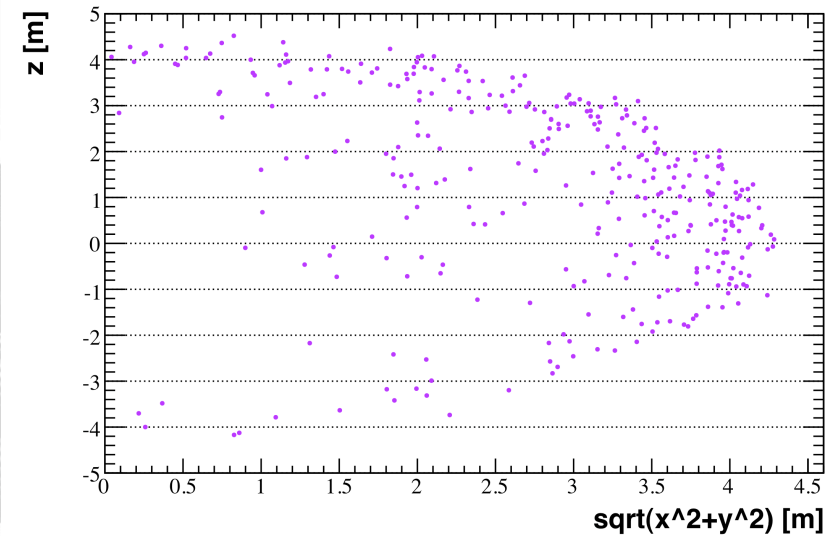
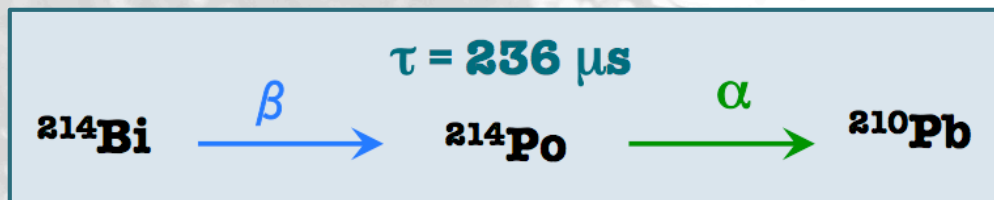
Rate measurements in agreement with Kamland results (arXiv:0907.0066)



# BACKGROUND SOURCES (above 3MeV)

## $^{214}\text{Bi}$ CONTAMINATION

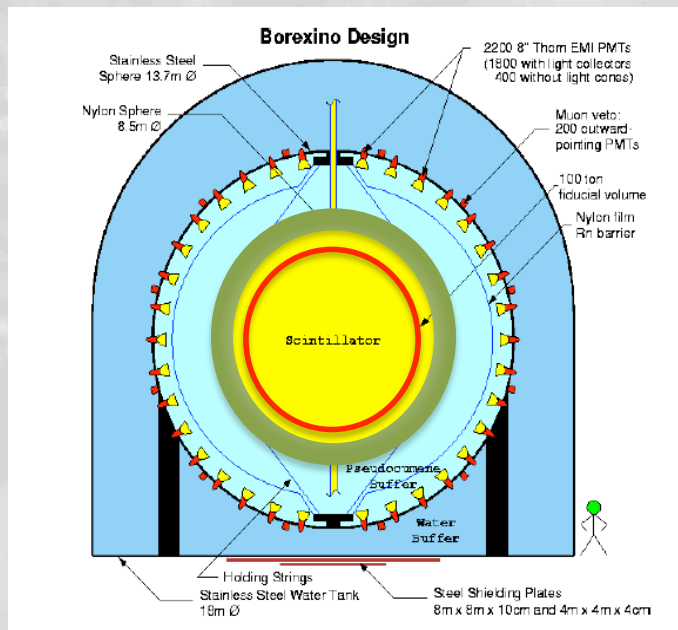
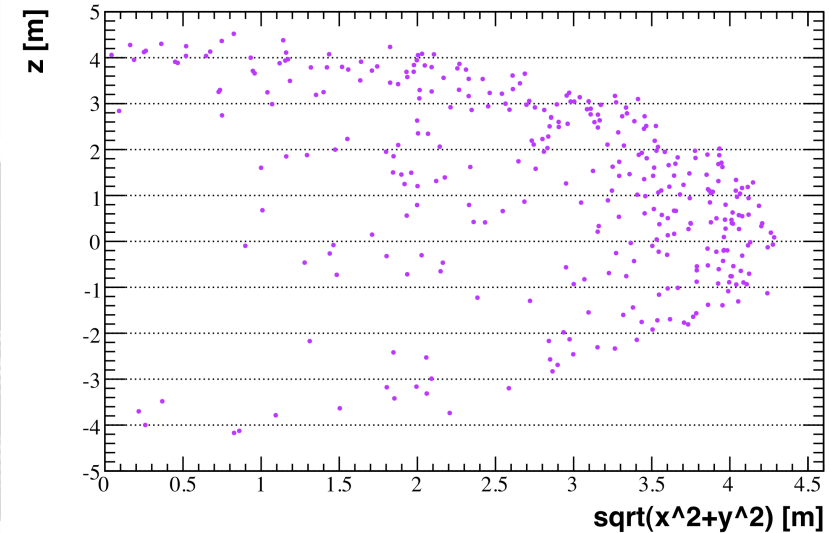
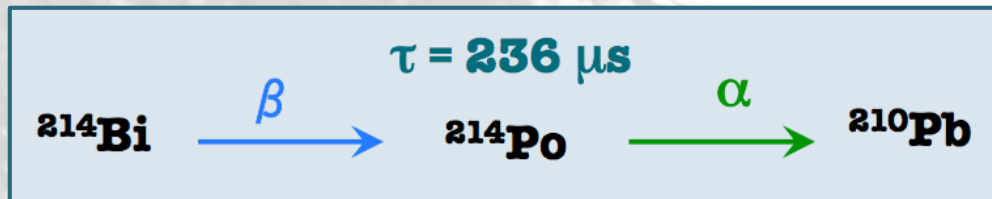
Searching for fast coincidences in  
time and space between



# BACKGROUND SOURCES (above 3MeV)

## $^{214}\text{Bi}$ CONTAMINATION

Searching for fast coincidences in time and space between



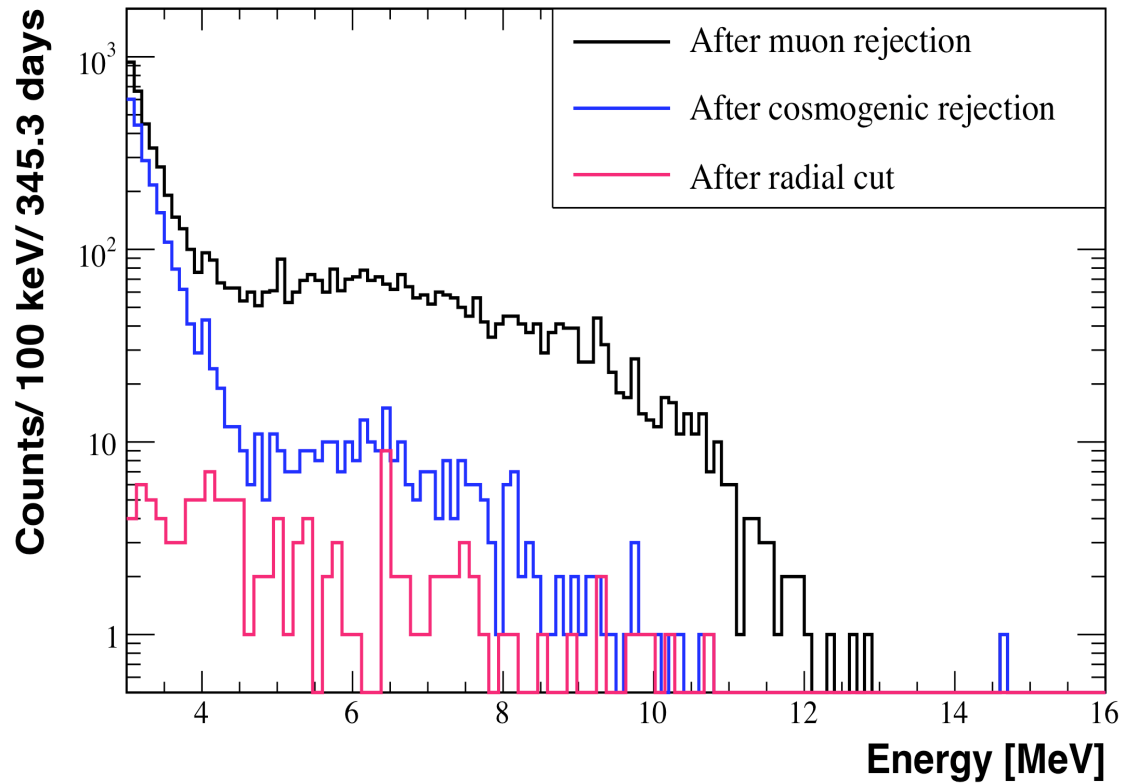
## FIDUCIAL VOLUME CUT

Radon contamination from the vessel... diffusion and decay in  $^{214}\text{Bi}$  and  $^{208}\text{Tl}$

→ **Necessity of a radial cut at 3m**

# $^8\text{B}$ NEUTRINO FLUX MEASUREMENT (2)

## ENERGY SPECTRUM OF OUR DATA

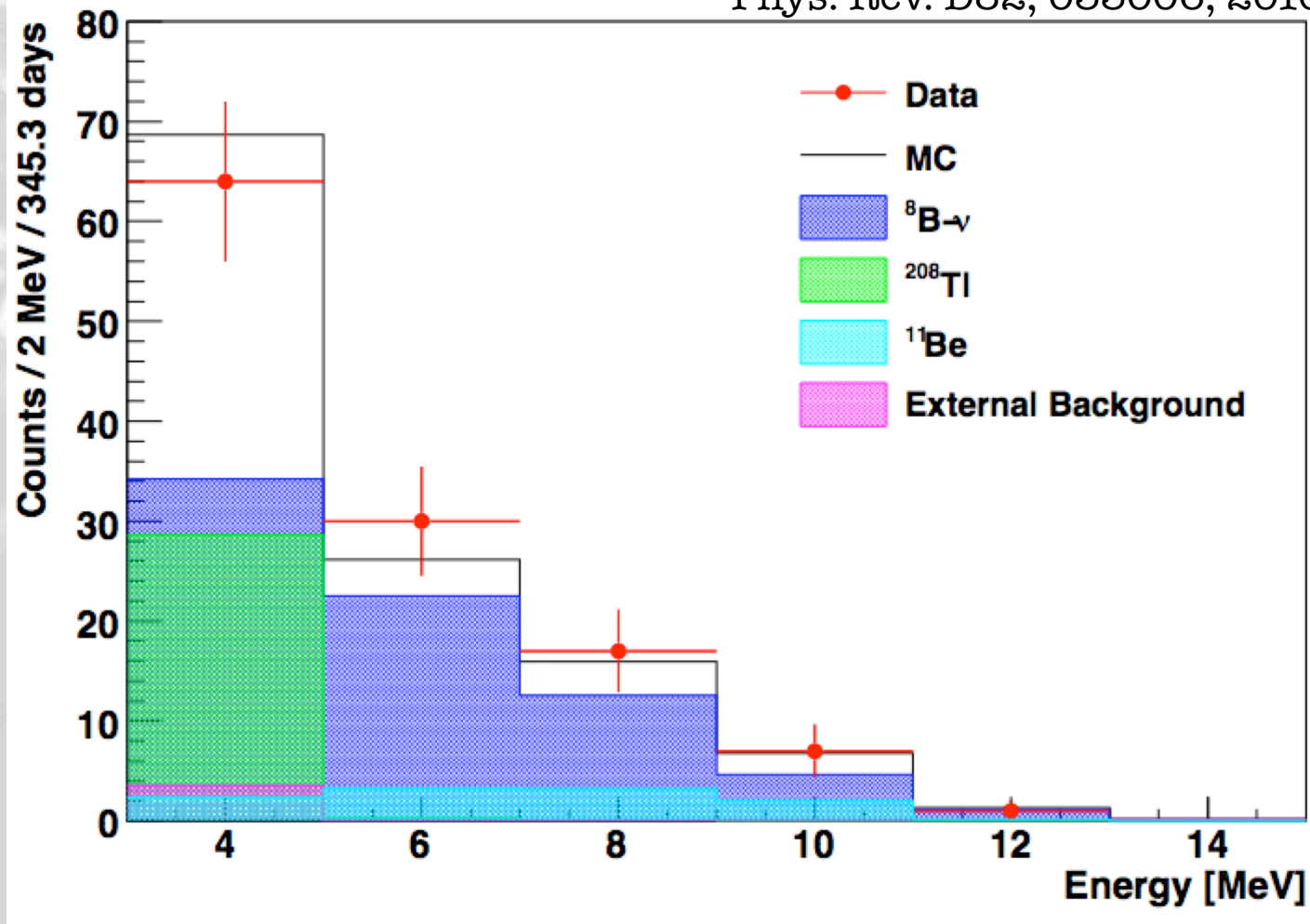


- After muons rejection
- After cosmogenic rejection
- After the radial cut

# RESULTS ON THE $^8\text{B}$ $\nu$ MEASUREMENT (1)

Comparison data/MonteCarlo for 345.3 d of livetime:

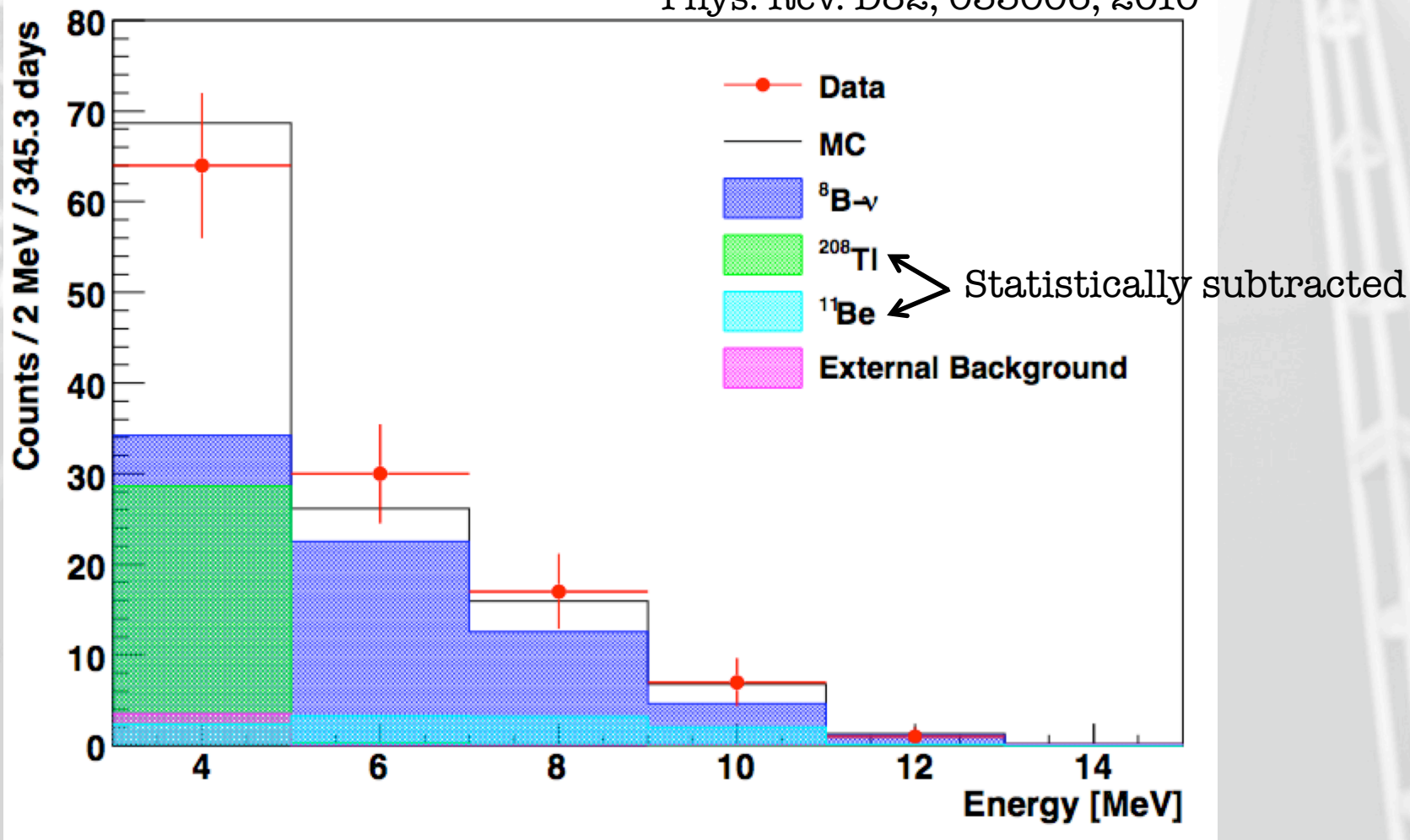
Phys. Rev. D82, 033006, 2010



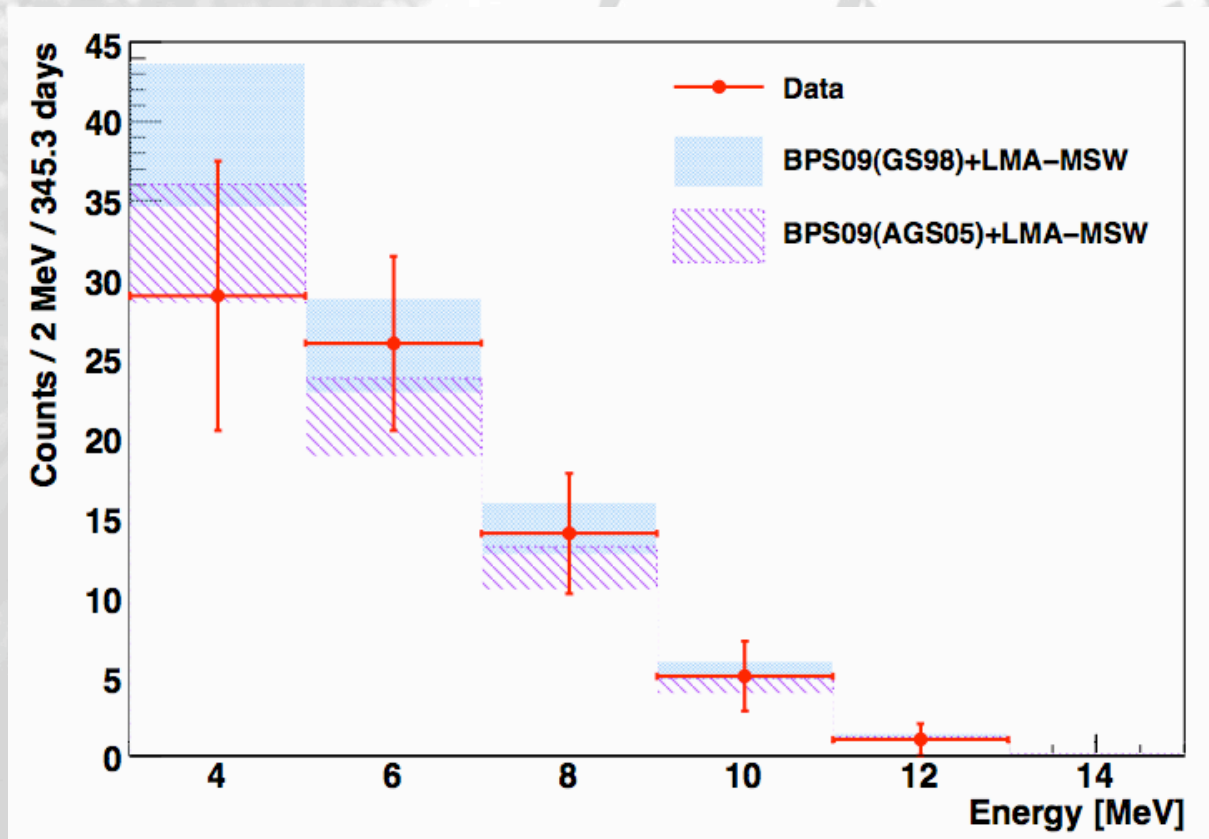
# RESULTS ON THE ${}^8\text{B}$ $\nu$ MEASUREMENT (1)

Comparison data/MonteCarlo for 345.3 d of livetime:

Phys. Rev. D82, 033006, 2010



# RESULTS ON THE $^8\text{B}$ $\nu$ MEASUREMENT (2)

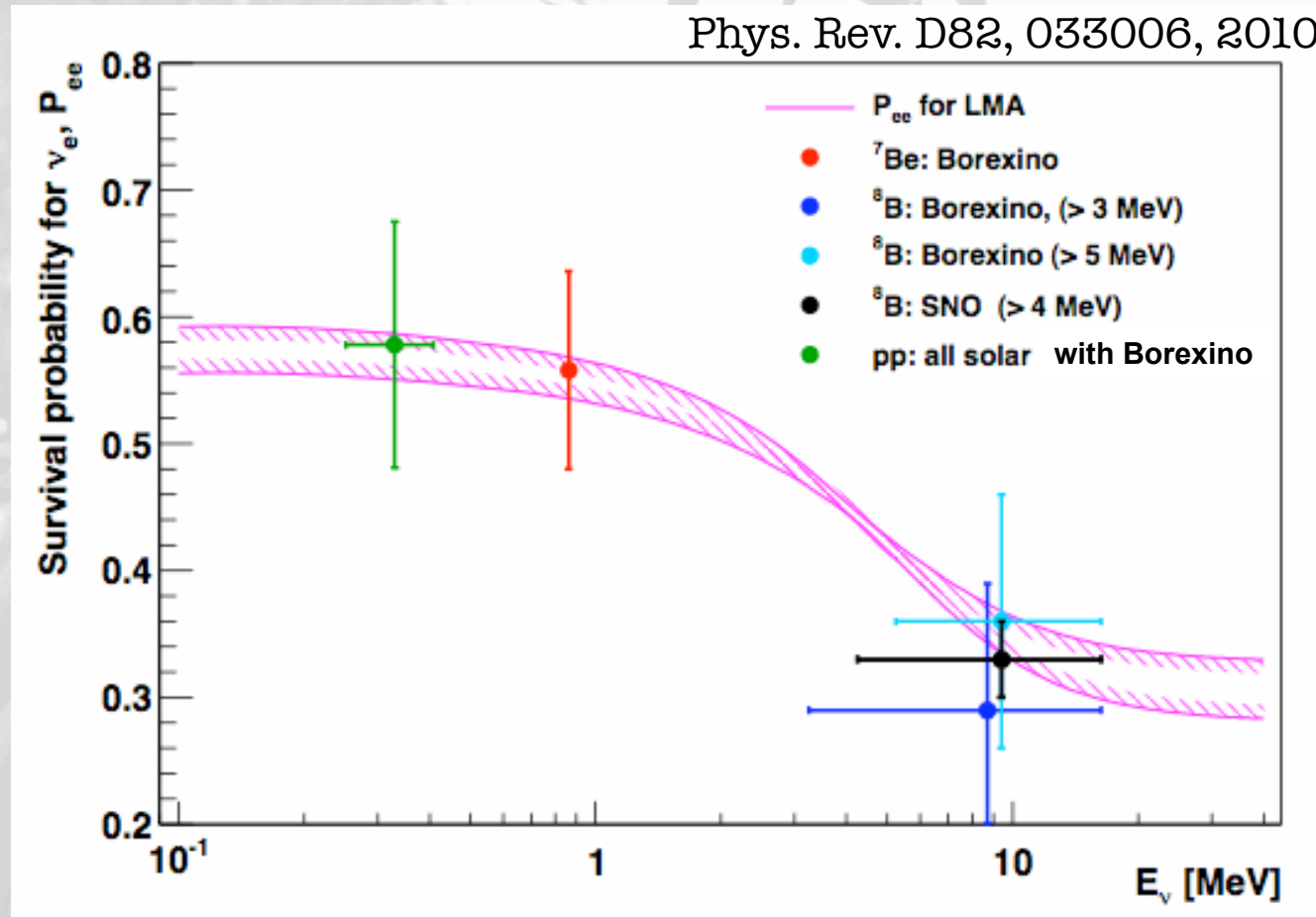


$$\mathbf{R}_{8\text{B}} (E>3\text{MeV}) = \mathbf{0.217} \pm \mathbf{0.038}_{\text{stat}} \pm \mathbf{0.008}_{\text{sys}} \text{ c/d/100 tons}$$

$$\mathbf{R}_{8\text{B}} (E>5\text{MeV}) = \mathbf{0.134} \pm \mathbf{0.022}_{\text{stat}} \pm \mathbf{0.008}_{\text{sys}} \text{ c/d/100 tons}$$

# BOREXINO AND THE MSW-LMA SOLUTION (2010)

Phys. Rev. D82, 033006, 2010



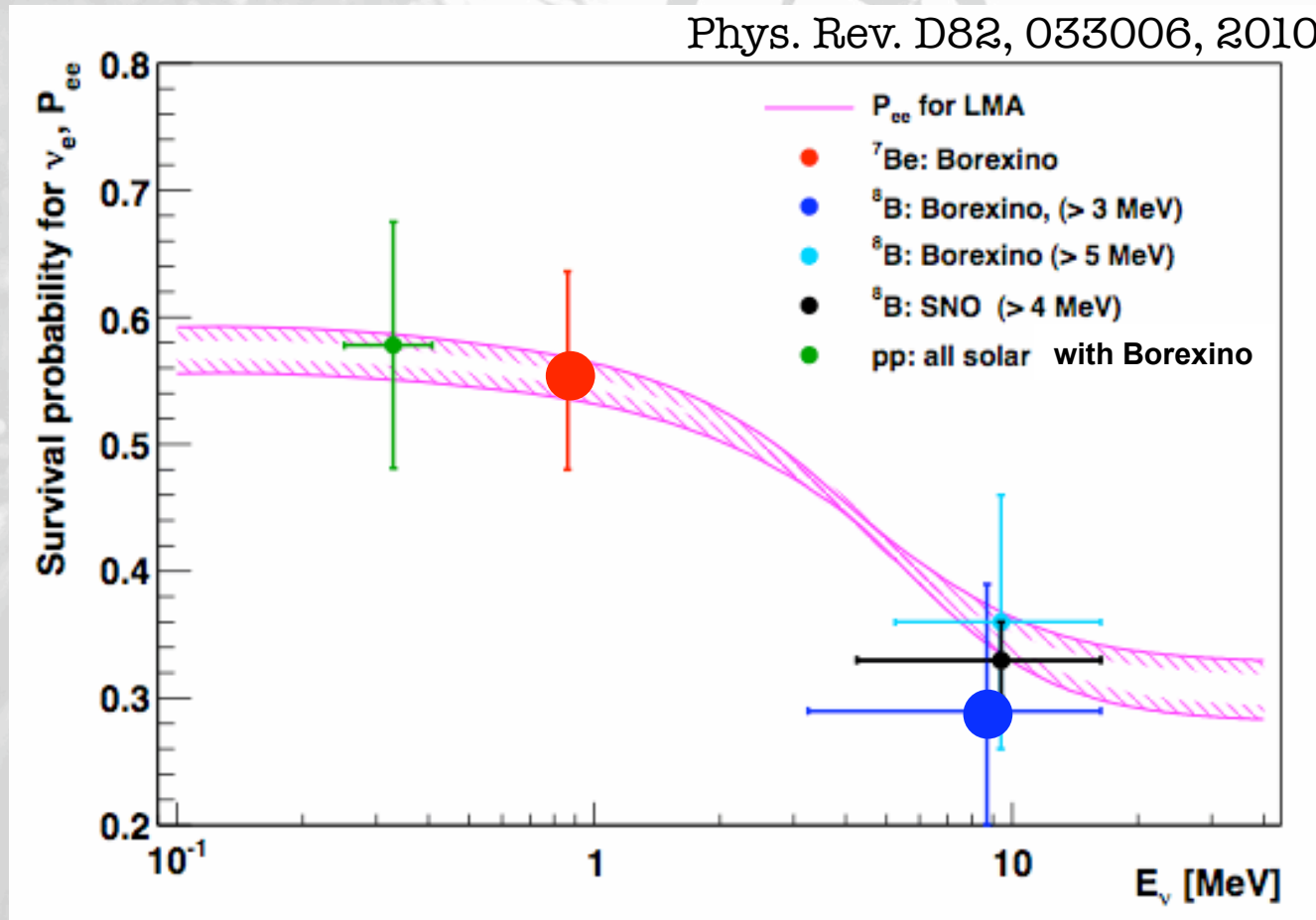
➤  ${}^7\text{Be}$  + constraint on pp: experimental evidence in the vacuum region!!!

➤  ${}^8\text{B}$ : 2 new points in the matter region!!!



# BOREXINO AND THE MSW-LMA SOLUTION (2010)

Phys. Rev. D82, 033006, 2010



$$P_{ee}({}^7\text{Be}) / P_{ee}({}^8\text{B}) = 1.93 \pm 0.75$$

Vacuum/matter ratio @  $1.9\sigma$  level !!

Using data from the same apparatus !!

# WHAT ELSE?? DAY-NIGHT ASYMMETRY !!

$$ADN = \frac{N - D}{(N + D)/2}$$

- **MSW mechanism:**  $\nu_{\mu}$  interaction in the Earth could lead to a  $\nu_e$  regeneration effect
- The size of the effect depends on:
  - detector **latitude**
  - neutrino **energy**
  - oscillation **parameters**
- Very **small effect expected with MSW-LMA:**
  - LOW solution predicts a large ADN effect
  - LMA and LOW predict similar  ${}^7\text{Be}$  absolute values but very different ADN

Observable	LMA ( $\pm 3 \sigma$ )	LOW ( $\pm 3 \sigma$ )
${}^7\text{Be}-\nu_e P_{ee}$	$0.64^{+0.09}_{-0.05}$	$0.58 \pm 0.05$
ADN(%) ${}^7\text{Be}$	$0.0^{+0.1}_{-0.0}$	$23^{+10}_{-13}$

# WHAT ELSE?? DAY-NIGHT ASYMMETRY !!

$\nu_e$  flux during night (average over 1 year)

$$ADN = \frac{N - D}{(N + D)/2}$$

$\nu_e$  flux during day (average over 1 year)

- **MSW mechanism:**  $\nu_\mu$  interaction in the Earth could lead to a  $\nu_e$  regeneration effect
- The size of the effect depends on:
  - detector **latitude**
  - neutrino **energy**
  - oscillation **parameters**
- Very **small effect expected with MSW-LMA:**
  - LOW solution predicts a large ADN effect
  - LMA and LOW predict similar  ${}^7\text{Be}$  absolute values but very different ADN

Observable	LMA ( $\pm 3 \sigma$ )	LOW ( $\pm 3 \sigma$ )
${}^7\text{Be}-\nu_e P_{ee}$	$0.64^{+0.09}_{-0.05}$	$0.58 \pm 0.05$
ADN(%) ${}^7\text{Be}$	$0.0^{+0.1}_{-0.0}$	$23^{+10}_{-13}$

## ADN STUDY IN BOREXINO

Although already excluded by reactor+radiochemical data,  
**Borexino alone** can add an independent confirmation.

The  ${}^7\text{Be}$  flux is obtained from the separate full fits of the day and night spectra

- ${}^7\text{Be}$  Day spectrum: 387.46 days
- ${}^7\text{Be}$  Night spectrum: 401.57 days

$$ADN = \frac{N - D}{(N + D)/2} = 0.007 \pm 0.073 \text{ (stat)}$$

**Borexino alone validates the MSW-LMA model!!!**

New analysis in progress with 2/3 times better sensitivity

# ADN STUDY IN BOREXINO

Although already excluded by reactor+radiochemical data,  
**Borexino alone** can add an independent confirmation.

The  ${}^7\text{Be}$  flux is obtained from the separate day and night spectra

- ${}^7\text{Be}$  Day spectrum: 387.46 days
- ${}^7\text{Be}$  Night spectrum: 401.57 days

$$ADN = 0.007 \pm 0.073 \text{ (stat)}$$

**Borexino falsifies the MSW-LMA model!!!**

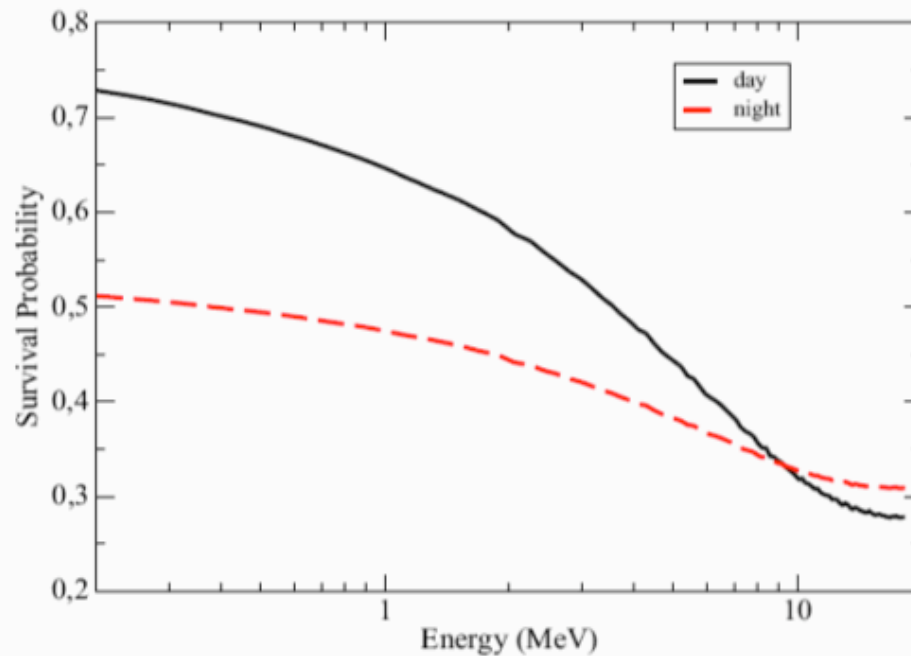
New analysis in progress with 2/3 times better sensitivity

**Preliminary**

# ADN MAY PROBE NEW PHYSICS

## Mass Varying Model

P. C. de Holanda, arXiv:0811:0567v1 (nov 2008)



Very large predicted effect:

**ADN = -23%**

(note negative sign!)

**Borexino alone excludes the MassVarying Model !!!**



# AND FOR THE FUTURE?!

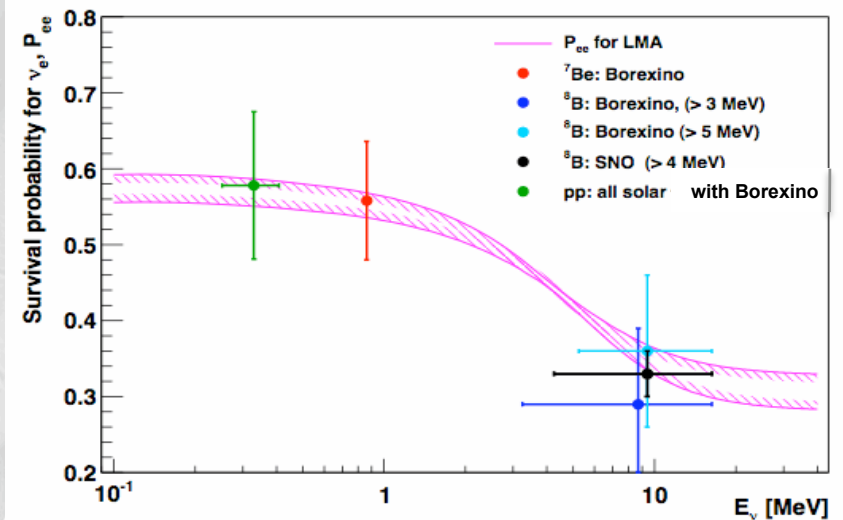
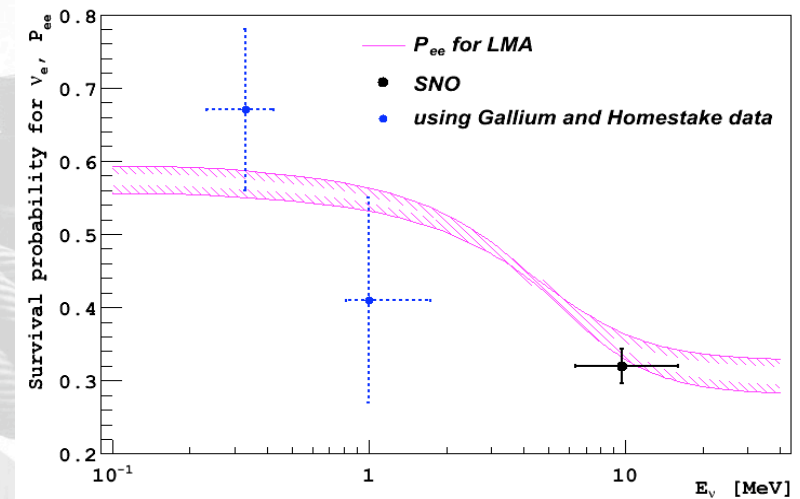
## $^8\text{B}$ and $^7\text{Be}$ NEUTRINOS!!

- Reduce the error on  $^7\text{Be}$  flux down to **5%**
  - ➔ better constraint also on **pp neutrinos**

# AND FOR THE FUTURE?!

## $^8\text{B}$ and $^7\text{Be}$ NEUTRINOS!!

- Reduce the error on  $^7\text{Be}$  flux down to **5%**
- ➔ better constraint also on **pp neutrinos**

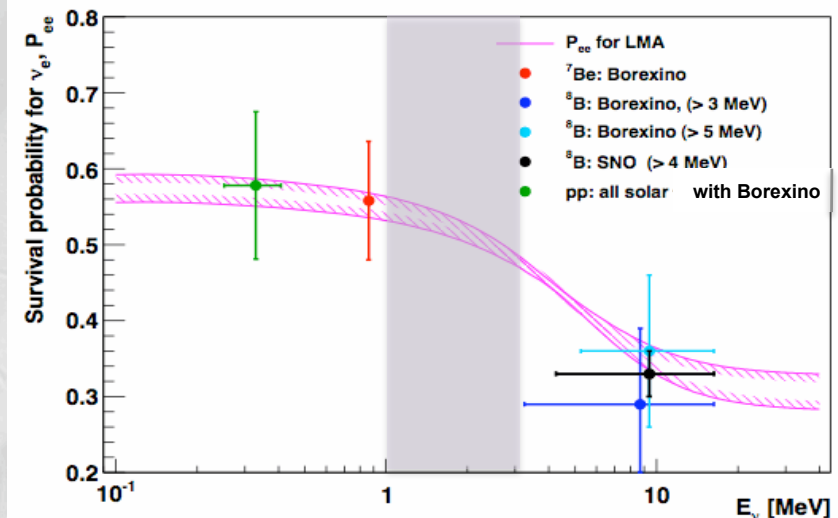
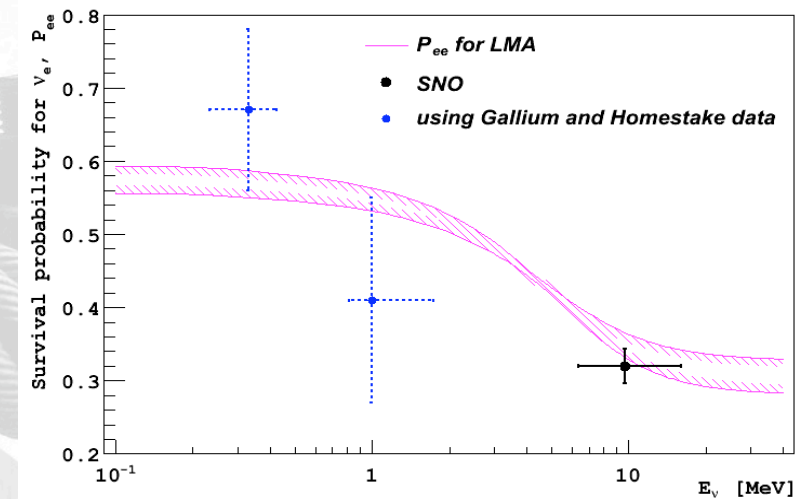




# AND FOR THE FUTURE?!

## $^8\text{B}$ and $^7\text{Be}$ NEUTRINOS!!

- Reduce the error on  $^7\text{Be}$  flux down to **5%**  
➔ better constraint also on **pp neutrinos**
- Reduce the energy threshold on  $^8\text{B}$  neutrino measurement down to **2MeV**  
➔ approaching the unexplored region between 1 and 3 MeV !!!



# AND FOR THE FUTURE?!

## $^7\text{Be}$ , $^8\text{B}$ and pep NEUTRINOS!!

- Reduce the error on  $^7\text{Be}$  flux down to **5%**
  - ➔ better constraint also on **pp neutrinos**
- Reduce the energy threshold on  $^8\text{B}$  neutrino measurements down to **2MeV**
  - ➔ approaching the unexplored region between 1 and 3 MeV !!!
- **Pep measurement:** fundamental test of  $P_{ee}$  in an unexplored energy region!!

