International School on Nuclear Physics, 32nd Course: *Particle and Nuclear Astrophysics* Erice, Sicily, 16–24 September 2010

# Neutrino Astronomy with KM3NeT

was: Mediterranean Neutrino Telescopes - Status and Future

Uli Katz ECAP / Univ. Erlangen 17.09.2010

#### ERLANGEN CENTRE FOR ASTROPARTICLE PHYSICS

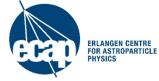
Friedrich-Alexander-Universität Erlangen-Nürnberg





- Introduction
- Technical solutions: Decisions and options
- Physics sensitivity
- Cost and implementation
- Summary

# **KM3NeT**



U. Katz: KM3NeT (Erice 2010)

#### **The Neutrino Telescope World Map**



AMANDA South Pole

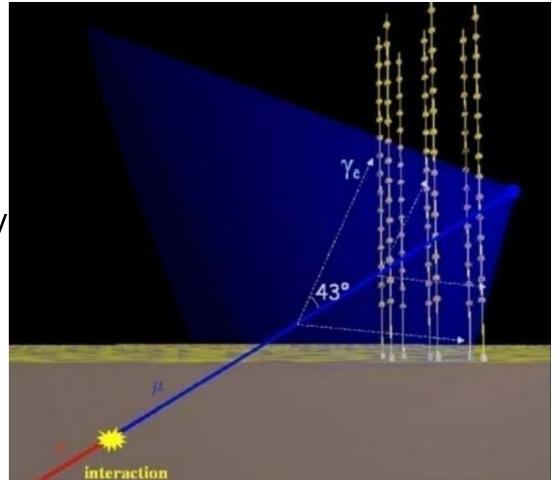




U. Katz: KM3NeT (Erice 2010)

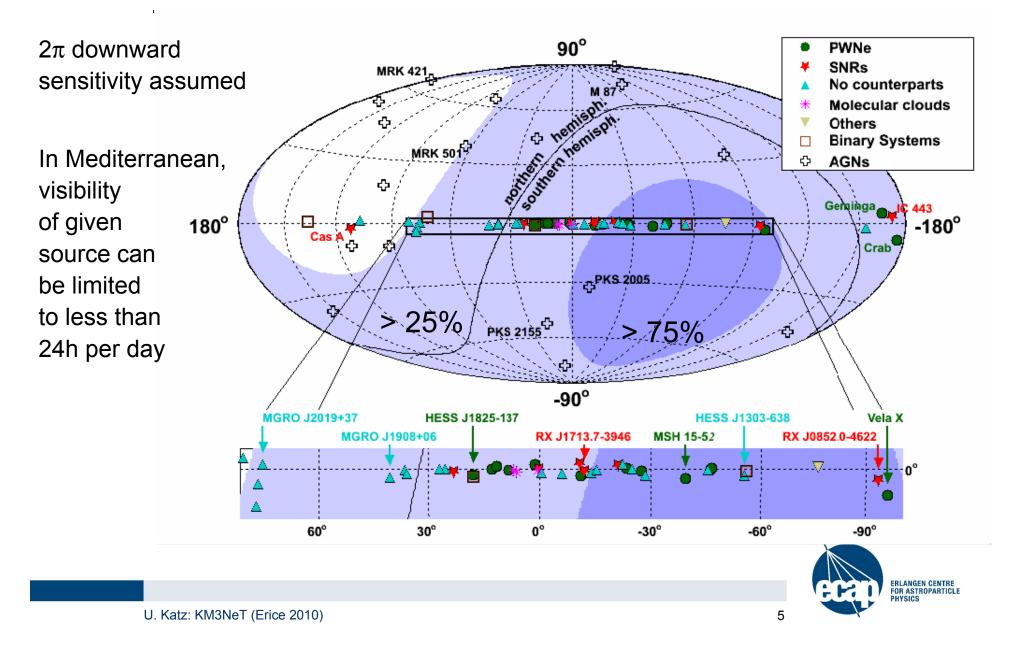
#### What is KM3NeT ?

- Future cubic-kilometre scale neutrino telescope in the Mediterranean Sea
- Exceeds Northernhemisphere telescopes by factor ~50 in sensitivity
- Exceeds IceCube sensitivity by substantial factor
- Provides node for earth and marine sciences





#### **South Pole and Mediterranean Fields of View**



## **The Objectives**

- <u>Central physics goals:</u>
  - Investigate neutrino "point sources" in energy regime 1-100 TeV
  - Complement IceCube field of view
  - Exceed IceCube sensitivity
  - Not in the central focus:
    - Dark Matter
    - Neutrino particle physics aspects
    - Exotics (Magnetic Monopoles, Lorentz invariance violation, ...)
- Implementation requirements:
  - Construction time ≤5 years
  - Operation over at least 10 years without "major maintenance"



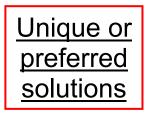
### **Technical Design**

<u>Objective</u>: Support 3D-array of photodetectors and connect them to shore (data, power, slow control)

- Optical Modules
- Front-end electronics
- Readout, data acquisition, data transport
- •••• Mechanical structures, backbone cable
- General deployment strategy
- Sea-bed network: cables, junction boxes
  - Calibration devices
    - Shore infrastructure
    - Assembly, transport, logistics
    - Risk analysis and quality control

#### Design rationale:

Cost-effective Reliable Producible Easy to deploy





#### **Further Challenges**

#### Site characteristics

<u>Objective</u>: Measure site characteristics (optical background, currents, sedimentation, ...)

#### Simulation

<u>Objective</u>: Determine detector sensitivity, optimise detector parameters;

#### Earth and marine science node <u>Objective</u>: Design interface to instrumentation for marine biology, geology/geophysics, oceanography, environmental studies, alerts, ...

#### Implementation

<u>Objective</u>: Take final decisions (technology and site), secure resources, set up proper management/governance, construct and operate KM3NeT;

OB ASTROPARTICLE

#### **The First-Generation Projects**

• ANTARES:

See presentation by Thomas Eberl, today 18:00-18:30

- NEMO and NESTOR Major contributions to R&D Site exploration
- All 3 have become part of KM3NeT



### **NEMO**

- Extensive site exploration • (Capo Passero near Catania, depth 3500 m);
- R&D towards km<sup>3</sup>: architecture, mechanical structures, readout, electronics, cables ...;
- Simulation.

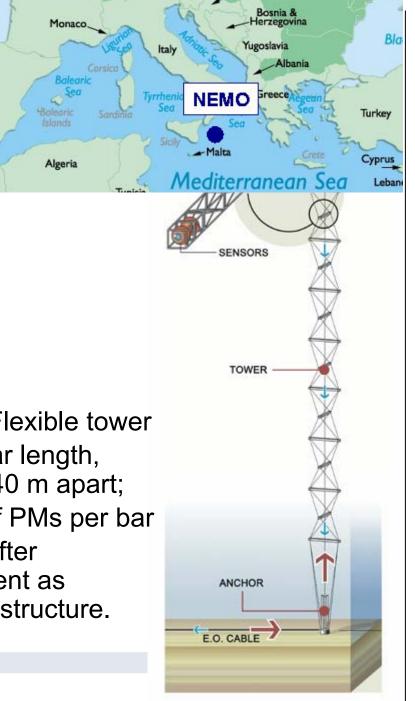
**Example:** Flexible tower

~10 m bar length, bars 30-40 m apart;

AUUIILIC Ocean

Spain

- 3 pairs of PMs per bar
- Unfurls after deployment as compact structure.



# NESTOR

- Tower based detector (titanium structures).
- Dry connections (recover – connect – redeploy).
- Up- and downward looking PMs (15").
- 4000-5200 m deep.
- Test floor (reduced size) deployed & operated in 2003.

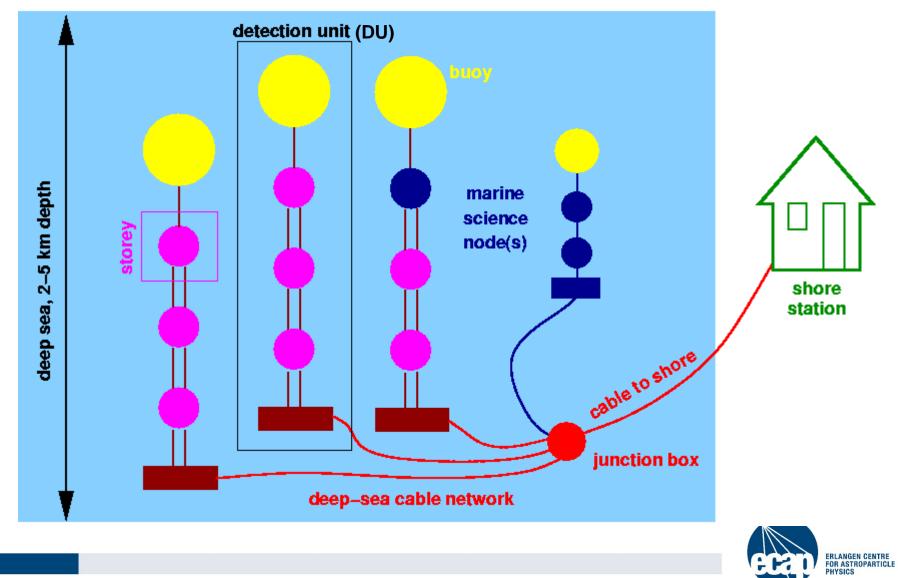


#### **NESTOR: the Delta-Berenike Platform**



U. Katz: KM3NeT (Erice 2010)

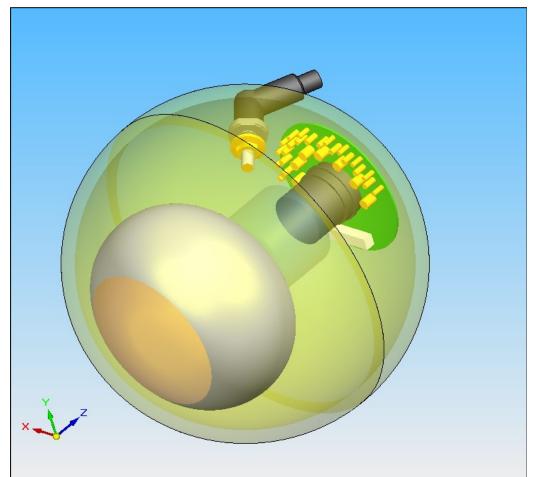
### The KM3NeT Research Infrastructure (RI)



### **OM "classical": One PMT, no Electronics**

#### Evolution from pilot projects:

- 8-inch PMT, increased quantum efficiency (instead of 10 inch)
- 13-inch glass sphere (instead of 17 inch)
- no valve (requires "vacuum" assembly)
- no mu-metal shielding

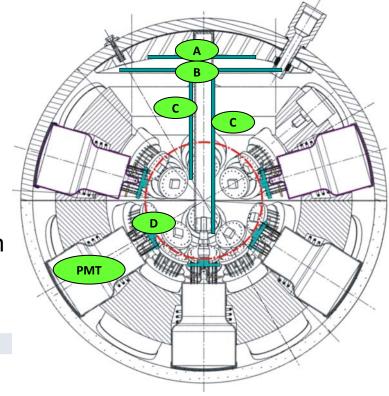




# **OM with many Small PMTs**

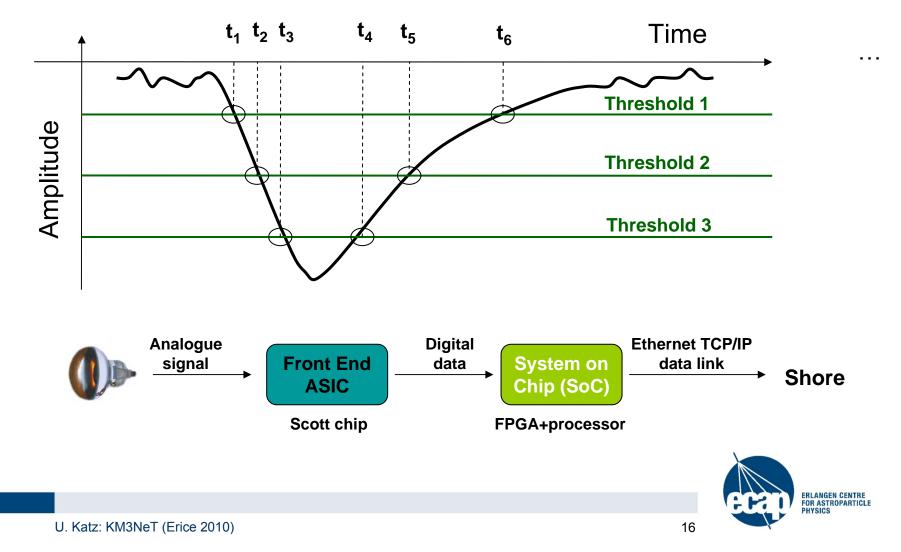
- 31 3-inch PMTs in 17-inch glass sphere (cathode area~ 3x10" PMTs)
  - 19 in lower, 12 in upper hemisphere
  - Suspended by compressible foam core
- 31 PMT bases (total ~140 mW) (D)
- Front-end electronics (B,C)
- Al cooling shield and stem (A)
- Single penetrator
- 2mm optical gel
- Advantages:
  - increased photocathode area
  - improved 1-vs-2 photo-electron separation
     → better sensitivity to coincidences
  - directionality





#### **Front-End Electronics: Time-over-Threshold**

From the analogue signal to time stamped digital data:



#### **Data Network**

All data to shore:

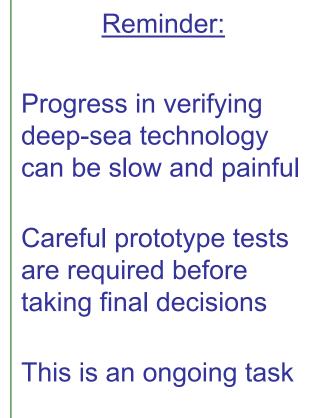
Full information on each hit satisfying local condition (threshold) sent to shore

- <u>Overall data rate</u> ~ 25 Gbyte/s
- <u>Data transport:</u> Optical point-to-point connection shore-OM Optical network using DWDM and multiplexing Served by lasers on shore Allows also for time calibration of transmission delays
- <u>Deep-sea components</u>: Fibres, modulators, mux/demux, optical amplifiers (all standard and passive)



# **DUs: Bars, Strings, Triangles**

- Flexible towers with horizontal bars
  - Simulation indicates that "local 3D arrangement" of OMs increases sensitivity significantly
  - Single- or multi-PMT OMs
- Slender strings with multi-PMT OMs
  - Reduced cost per DU, similar sensitivity per Euro
- Strings with triangular arrangements of PMTs
  - Evolution of ANTARES concept
  - Single- or multi-PMT OMs
  - "Conservative" fall-back solution





#### **The Flexible Tower with Horizontal Bars**

- 20 storeys
- Each storey supports 6 OMs in groups of 2
- Storeys interlinked by tensioning ropes, subsequent storeys orthogonal to each other
- Power and data cables separated from ropes; single backbone cable with breakouts to storeys
- Storey length = 6m
- Distance between storeys = 40 m
- Distance between DU base and first storey = 100m



40m

# **The Slender String**

- Mooring line:
  - Buoy (empty glass spheres, net buoyancy 2250N)
  - Anchor: concrete slab of 1m<sup>3</sup>
  - 2 Dyneema ropes (4 mm diameter)
  - 20 storeys (one OM each),
    30 m distance, 100m anchor-first storey
- Electro-optical backbone:
  - Flexible hose ~ 6mm diameter
  - Oil-filled

New concept, needs to be tested. Also for flexible tower if successful

One single pressure transition

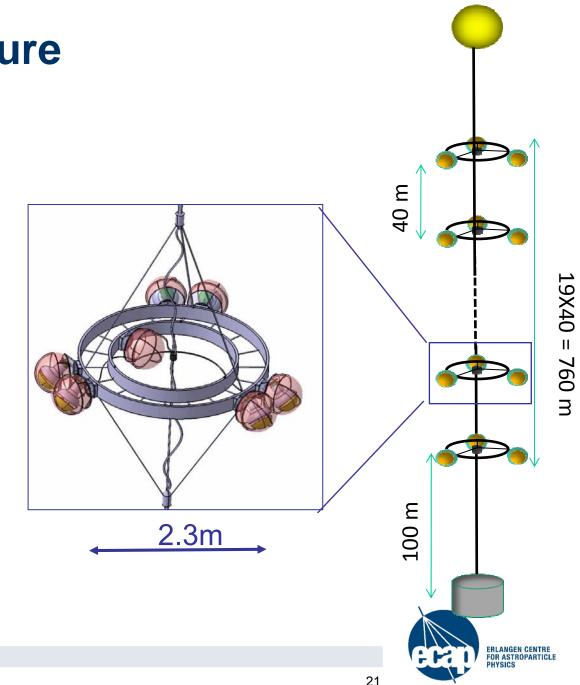
 Star network between master module and optical modules





# **Triangle Structure**

- Evolution from ANTARES concept
- 20 storeys/DU, spacing 40m
- Backbone: electrooptical-mechanical cable
- Reduced number of electro-optical penetrations
- Use ANTARES return
   of experience



## **Deployment Strategy**

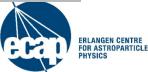
- All three mechanical solutions: Compact package – deployment – self-unfurling
  - Eases logistics (in particular in case of several assembly lines)
  - Speeds up and eases deployment; several DUs can be deployed in one operation
  - Self-unfurling concepts need to be thoroughly tested and verified
- Connection to seabed network by ROV
- Backup solution:

"Traditional" deployment from sea surface



#### **A Flexible Tower Packed for Deployment**

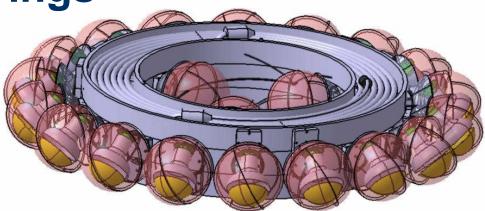




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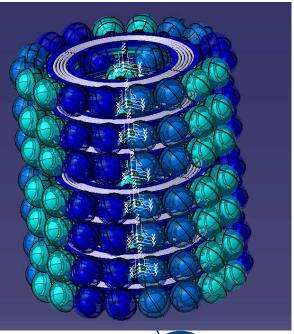
### **Compactifying Strings**

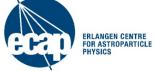
Slender string rolled up for self-unfurling:



3 triangles

DU





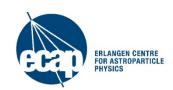
U. Katz: KM3NeT (Erice 2010)

# Hydrodynamic Stability

- DUs move under drag of sea current
  - Currents of up to 30cm/s observed
  - Mostly homogeneous over detector volume
  - Deviation from vertical at top:

Current	flexible tower	slender string	triangles
[cm/s]	d [m]	d [m]	d [m]
30	84.0	83.0	87.0

• Torsional stability also checked

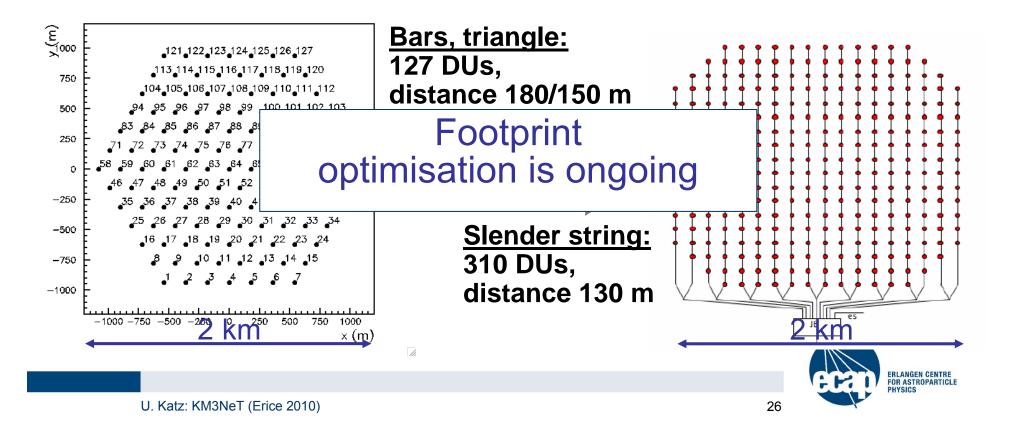


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### **Detector Building Blocks**

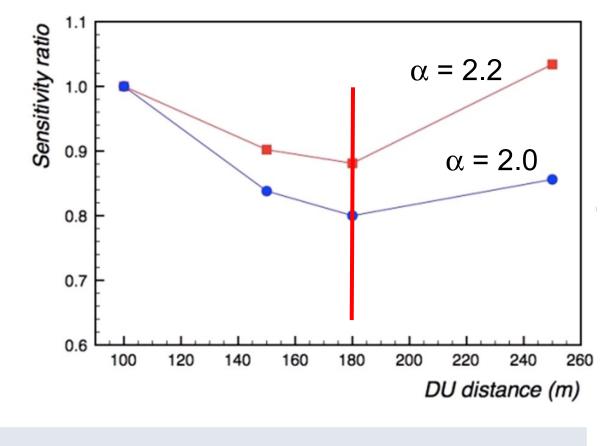
- Different DU designs
  - require different DU distance
  - differ in photocathode area/DU
  - are different in cost

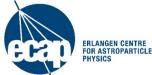




### **Optimisation Studies**

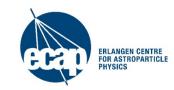
<u>Example</u>: Sensitivity dependence of point-source search on DU distance for flexible towers (for 2 different neutrino fluxes  $\sim E^{-\alpha}$ , no cut-off)



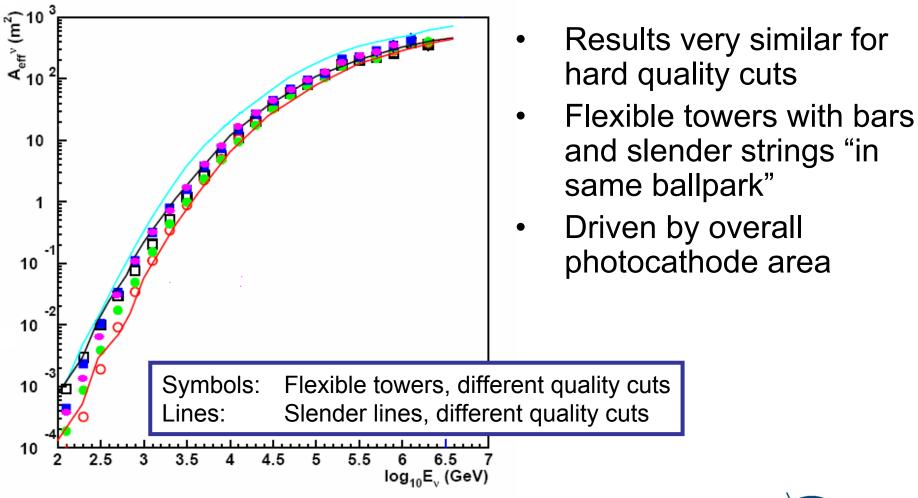


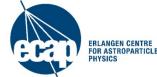
## **Angular Resolution**

median (degree) Investigate  $\bullet$ Ο. distribution of angle between incoming neutrino and reconstructed muon • Dominated by kinematics up to 10<sup>-2</sup>2 3 5 6 4 ~1TeV  $log_{10}E_{v}$  (GeV)



#### **Effective Areas (per Building Block)**





#### **Cost Estimates: Assumptions**

- Estimate of investment cost
  - no personnel costs included
  - no contingency, no spares
- Assumptions / procedure:
  - Quotations from suppliers are not official and subject to change
  - Common items are quoted with same price
  - Sea Sciences and Shore Station not estimated
  - Estimates worked out independently by expert groups and carefully cross-checked and harmonised thereafter

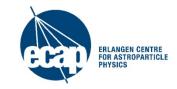


### **Cost Estimates: Results**

• Result of cost estimates (per building block):

Concept	DU Cost (M€)	No. of DUs	Total DU Cost (M€)	Seafloor Infrastr. (M€)	Deploy- ment (M€)	TOTAL COST (M€)
Flexible towers	0.54	127	68	8	11	87
Slender strings	0.25	310	76	13	14	103
Triangles	0.66	127	83	8	7	99

• Assembly man power (OMs, DU...) is roughly estimated to be 10% of the DU cost

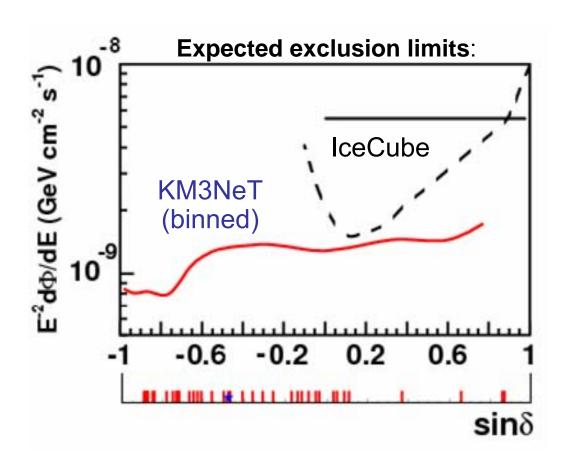


# **KM3NeT: Full Configuration**

- 2 "building blocks" needed to achieve objectives
- Increases sensitivity by a factor 2
- Overall investment ~220 M€
- Staged implementation possible
- Science potential from very early stage of construction on
- Operational costs 4-6 M€ per year (2-3% of capital investment), including electricity, maintenance, computing, data centre and management



# **Point Source Sensitivity (1 Year)**

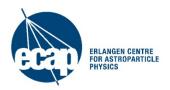


R. Abbasi et al. Astro-ph
(2009) scaled – unbinned method

Aharens et al. Astr. Phys. (2004) – binned method

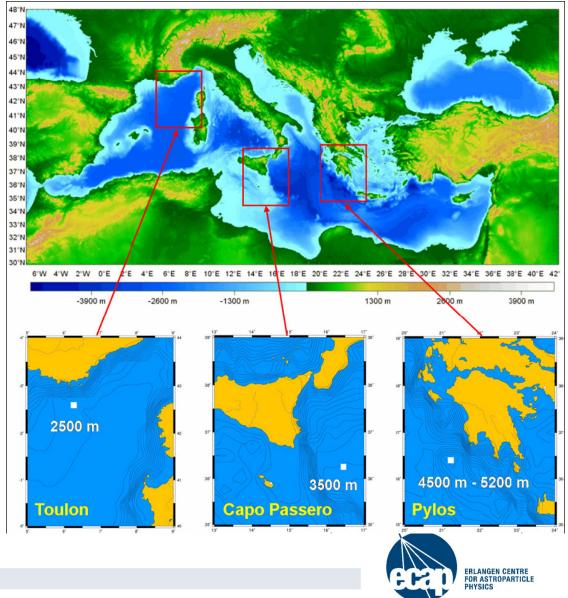
Observation of RXJ1713 with  $5\sigma$  within ~8 years

 Observed Galactic TeV-γ sources (SNR, unidentified, microquasars)
 F. Aharonian et al. Rep. Prog. Phys. (2008)
 Abdo et al., MILAGRO, Astrophys. J. 658 L33-L36 (2007)



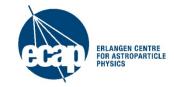
### **Candidate Sites**

- Locations of the three pilot projects:
  - ANTARES: Toulon
  - NEMO: Capo Passero
  - NESTOR: Pylos
- Long-term site characterisation measurements performed
- Site decision requires scientific, technological and political input



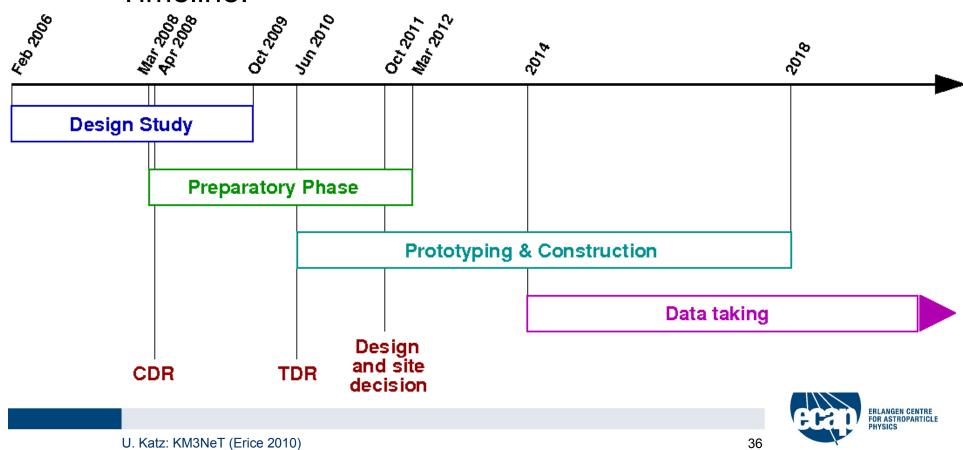
#### **Recent Developments**

- Convergence towards a bar structure with multi-PMT OMs: 6m bars with 1 OM at each end
- Prototyping of components under way
- Simulation and "footprint" studies ongoing
- Possible cooperation with IceCube being explored (towards a Global Neutrino Observatory)



### **Next Steps and Timeline**

- Next steps: Prototyping and design decisions
  - TDR public since June 2010
  - final decisions require site selection
  - expected to be achieved in 15 months
- Timeline:



### Conclusions

- A design for the KM3NeT neutrino telescope complementing the IceCube field in its of view and surpassing it in sensitivity by a substantial factor is presented.
- Readiness for construction expected in 15 months
- An overall budget of ~250 M€ will be required. Staged implementation, with increasing discovery potential, is technically possible.
- Within 15 months, remaining design decisions have to be taken and the site question clarified.
- Installation could start in 2013 and data taking soon after.

