p-process nucleosynthesis: activation experiments

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Outline



introduction

activation experiments

preliminary results

summary & outlook

Nucleosynthesis of heavy elements







Starting point: seed distribution of *s*- and *r*-process

- (γ,n) reactions shift distribution to proton-rich isotopes
- (γ , α) and (γ ,p) start to compete with (γ ,n) reaction rates
- position of branching nuclei determines final p-abundances



Reaction network of *p*-process



simulation of a type II SN explosion Bi (83) Pb (82) TI (81) ■M = 25 M_{solar} Hg (80) 126 Au (79) Pt (78) Ir (77) 121 Os (76) 118 120 shock front moving through O/Ne layer 117 Ta (73) 113 Hf (72) 110 112 integrated reaction flux Lu (71) Yb (70) Tm (69) during first second Er (68) 104 106 Ho (67) 103 Dy (66) 100 102 Tb (65) 99 Gd (64) 98 Eu (63) 97 Sm (62) 94 96 Pm (61) 93 Sn (50) Nd (60) ln (49) Pr (59) 72 Cd (48) 70 74 Ce (58) 86 88 90 Ag (47) 69 La (57) 85 66 68 Pd (46) Ba (56) Rh (45) 65 Cs (55) 64 Ru (44) 62 Xe (54) Tc (43) 6 l (53) Mo (42) Te (52) 82 Nb (41) Sb (51) Zr (40) 58 Sn (50) Y (39) 57 In (49) 56 Sr (38) 54 Cd (48) 72 70 48 50 52 54 Rb (37) 53 Kr (36) Br (35) Se (34) As (33) W. Rapp et al 2006 ApJ 653 474 Ge (32) Ga (31) 33 29 31 35

Reaction network of *p*-process





Reaction network of *p*-process



- complex reaction network involving more than ten thousand reactions
- calculations need lot of physics input (reaction rates)
- rates cannot be measured for all reactions
 - -> calculation within the statistical Hauser-Feshbach model
- input parameters for statistical model codes:
 - ➤ optical potentials
 - $\succ \gamma$ -width
 - level densities
 - ➤ masses
- HF predictions have to be tested carefully by measurement
 - \rightarrow optimization of input parameters

Measurements @ Darmstadt & Notre Dame





Calculations in the statistical HF model



Compound nucleus theory

entrance & exit channel are independent

$$\sigma_{\rm HF} \propto \sum_{n} \left(2J_n + 1\right) \frac{\langle \Gamma_{J_n}^{\rm form} \rangle \langle \Gamma_{J_n}^{\rm dec} \rangle}{\langle \Gamma_{J_n}^{\rm dec, tot} \rangle}$$

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Consequence for (α,n) & (p,n) reactions

- cp-width << n-width @ stellar temperatures (Coulomb supression)</p>
- above n-separation energy: $\langle \Gamma_{J_n}^{\text{dec,neutron}} \rangle \approx \langle \Gamma_{J_n}^{\text{dec,tot}} \rangle$
- smallest width determines cross section

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smallest width determines cross section

(α,n) & (p,n) reactions are not sensitive to neutron-width (sub-Coulomb)
good possibility to test α- & p-nucleus optical potentials



compound nucleus 1 compound nucleus, 3 reactions same exit channel $S_n = 8,47 MeV$ ¹⁶⁹Yb 168, 169, 0.13 32.0 3.04 d 167 Tm 168. 169 I m 100 9.25 d 93.1 d 166 167 168 ¹⁷⁰Yb 33.503 22.869 26.978















> equal systematic uncertainties influence all cross sections

γ-spectra



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Evaluation of data



CP-induced reactions (monoenergetic beam)

peak area directly proportional to cross section

$$Y_{\mathrm{peak}} \propto \sigma\left(E\right)$$

energy resolved cross section

Photo-induced reactions (bremsstrahlung)

peak area proportional to energy integrated cross section

$$Y_{\mathrm{peak}} \propto \int_{0}^{\infty} \sigma\left(E\right) n_{\gamma}\left(E\right) dE$$

energy integrated cross section

Preliminary results ¹⁷⁰Yb(γ,n)





Preliminary results ¹⁶⁹Tm(p,n)





Preliminary results ¹⁶⁶Er(α ,n)





Summary



Probing statistical Hauser-Feshbach model

Three activation measurements

- production of one compound nucleus ¹⁷⁰Yb
- ¹⁷⁰Yb(γ,n) @ S-DALINAC, Darmstadt, Germany
- In 166 Er(α,n) & 169 Tm(p,n) @ FN-TANDEM, University of Notre Dame, USA

Preliminary results compared to theory

- good description of $^{169}Tm(p,n)$ and $^{170}Yb(\gamma,n)$
- deviation by a factor of 6 for ${}^{166}Er(\alpha,n)$

Outlook



Further evaluation of data

correction for summing effects

Detailed comparison with HF predictions

- calculate XS with different HF codes (e.g. TALYS, NON-SMOKER)
- systematic variation of input parameters (optical potential)

Collect more experimental data

- cp-induced reactions in the rare earth region (x,n)
- ¹⁶⁵Ho(α,n) & ¹⁷⁵Lu(p,n)
- compare to recent results
- extend existing systematic studies

Thank you ...





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... for your attention!



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