# Nuclear Medium Effects in Lepton-Nucleus Deep Inelastic Scattering

#### Farhana Zaidi

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## **Farhana Zaidi**<sup>1</sup> Md. Sajjad Athar<sup>1</sup>, Huma Haider<sup>1</sup>, S. K. Singh<sup>1</sup> and I. Ruiz Simo<sup>2</sup>

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 Lepton scattering experiments have been performed for a wide range of energy.

- Lepton induced processes may be subdivided as
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  - Inelastic scattering
  - Obep Inelastic scattering(DIS)
- Early experiments at SLAC exhibited Bjorken scaling phenomenon corresponding to DIS.
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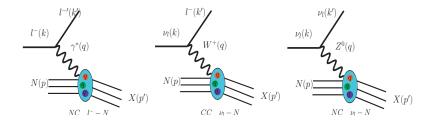
OY process

## Basic reaction for deep inelastic scattering process is

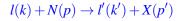
 $l(k) + N(p) \rightarrow l'(k') + X(p')$ 

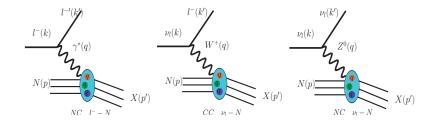
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The general expression of two body scattering cross section

$$d\sigma = \frac{1}{2E_l 2E_N} (2\pi)^4 \,\delta^4(k+p-k'-\sum_{i=1}^n p'_i) \,\frac{d^3k'}{(2\pi)^3 E_{l'}} \prod_{i=1}^n \frac{d^3p'_i}{(2\pi)^3 E_X} \,\sum \bar{\sum} |\mathcal{M}|^2$$

Square of matrix element is

$$|\mathcal{M}|^2 \propto L_{\mu
u}W^{\mu
u}$$

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The leptonic tensor  $L_{\mu\nu}$ 

$$L_{\mu\nu} = 2(k_{\mu} k_{\nu}' + k_{\mu}' k_{\nu} - k \cdot k' g_{\mu\nu})$$

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## In general the hadronic tensor is defined as

$$W^{\mu\nu} = -g_{\mu\nu} W_1(\nu, Q^2) + \frac{p_{\mu}p_{\nu}}{M^2} W_2(\nu, Q^2) - i\varepsilon_{\mu\nu\lambda\sigma} \frac{p^{\lambda}q^{\sigma}}{2M^2} W_3(\nu, Q^2) + \frac{q_{\mu}q_{\nu}}{M^2} W_4(\nu, Q^2) + \frac{(p_{\mu}q_{\nu} + p_{\nu}q_{\mu})}{2M^2} W_5(\nu, Q^2) + \frac{i(p_{\mu}q_{\nu} - p_{\nu}q_{\mu})}{2M^2} W_6(\nu, Q^2)$$

By contraction of hadronic tensor with  $L_{\mu\nu}$ 

$$W_3(\mathbf{v}, \mathbf{Q}^2) \rightarrow 0$$
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Conservation of current leads to

$$q_{\mu}W^{\mu\nu} = 0$$
  

$$W_4(\mathbf{v}, Q^2) = \frac{-2p \cdot q}{q^2} W_2(\mathbf{v}, Q^2)$$
  

$$W_5(\mathbf{v}, Q^2) = \frac{M^2}{q^2} W_1(\mathbf{v}, Q^2) + \left(\frac{p \cdot q}{q^2}\right)^2 W_2(\mathbf{v}, Q^2)$$

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Therefore, we are left with only  $W_1(v, Q^2)$  and  $W_2(v, Q^2)$  and the hadronic tensor is

$$W^{\mu\nu} = \left(-g^{\mu\nu} + \frac{q^{\mu}q^{\nu}}{q^2}\right)W_1(\nu, Q^2) + \left(p^{\mu} - \frac{p \cdot q}{q^2}q^{\mu}\right)\left(p^{\nu} - \frac{p \cdot q}{q^2}q^{\nu}\right)\frac{W_2(\nu, Q^2)}{M^2}$$

$$\frac{d\sigma}{dQ^2d\nu} = \frac{\pi\alpha^2}{4E_l^3 E_l' \sin^4\left(\frac{\theta}{2}\right)} \left\{ 2\sin^2\left(\frac{\theta}{2}\right) W_1(\nu, Q^2) + \cos^2\left(\frac{\theta}{2}\right) W_2(\nu, Q^2) \right\}$$

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$$\frac{d\sigma}{dxdy} = 2ME_l^2 y \frac{\pi\alpha^2}{4E_l^3 E_l' \sin^4(\frac{\theta}{2})} \left[ W_2(\mathbf{v}, Q^2) \frac{E_l}{E_l'} \left( 1 - y - \frac{Mxy}{2E_l} \right) + 2W_1(\mathbf{v}, Q^2) \left( xy \frac{M}{2E_l'} \right) \right]$$

The nucleon structure functions  $W_i(\mathbf{v}, Q^2)$  are redefined as

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Differential scattering cross section may also be written as

$$\frac{d^2\sigma}{dx\,dy} = \frac{8ME_l\pi\alpha^2}{Q^4} \left( y^2 x F_1(x) + \left[ (1 - y - \frac{xyM}{2E_l}) \right] F_2(x) \right)$$

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$$F_{2}^{ep} = \frac{4x}{9} (u_{v} + u_{s} + \bar{u}_{s} + c + \bar{c} + ....) + \frac{x}{9} (d_{v} + d_{s} + \bar{d}_{s} + s + \bar{s} + ....)$$

$$F_{2}^{en} = \frac{x}{9} (u_{v} + u_{s} + \bar{u}_{s} + s + \bar{s} + ....) + \frac{4x}{9} (d_{v} + d_{s} + \bar{d}_{s} + c + \bar{c} + ....)$$
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### Parton distribution functions have been parametrized by many groups

- MRST/MSTW
- GJR/GRV
- ALEKHIN
- CTEQ
- **()** .....

We have used CTEQ6.6 PDFs in our numerical calculations.

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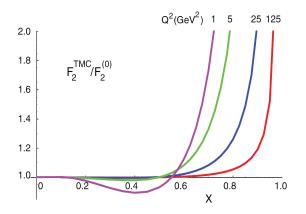
$$\xi = \frac{2x}{1 + \sqrt{1 + 4\mu x^2}}, \qquad \mu = \frac{M^2}{Q^2}$$

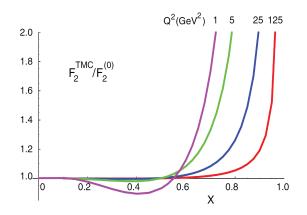
This effect is known as Target mass correction
 TMC is effective at low Q<sup>2</sup> and high x

- In this region PDFs are not very well determined
- Hence to precisely determine the PDFs TMC should be taken into account
- Structure function depend on the dimensionless variable x
- Therefore, nucleon structure function with TMC is given by

$$F_2^{TMC}(x,Q^2) \approx \frac{x^2}{\xi^2 \gamma^3} F_2(\xi) \left(1 + \frac{6\mu x\xi}{\gamma} (1-\xi)^2\right)$$
$$\gamma = \sqrt{1 + \frac{4M^2 x^2}{Q^2}}$$

Schienbein et al. JPG 35 (2008) 053101





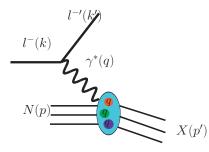
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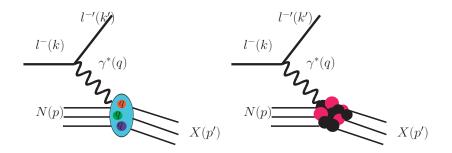
"Structure functions in DIS and their scale evolution are closely related to the origins of quantum chromodynamics (QCD)."

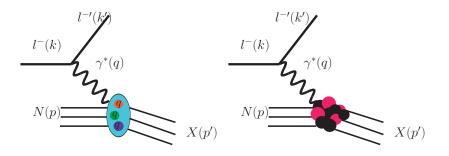
The expression for structure function  $F_2$  can be expressed as a function of the PDFs by

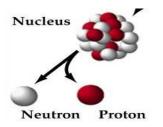
$$x^{-1}F_2 = \sum_{f=q,g} C_2 \otimes f$$

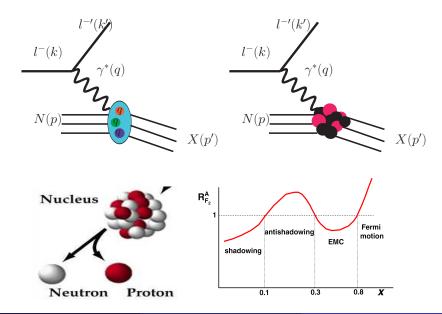
where  $C_2$  is the coefficient function for the quarks and gluons,  $\otimes$  symbols is for the Mellin convolution which turns into a simple multiplication in N-space and *f* represents the quark and gluon distributions. Vermaseren and van Neerven et al.NPB 724(2005)3 van Neerven and Vogt NPB 568(2000)263.

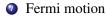












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- Shadowing and antishadowing
   Kulagin and Petti PRD76,094033, 2007

## Medium effects in lepton-A scattering

• Kinematic effect which arises as the struck nucleon is not at rest but is moving with a Fermi momentum in the rest frame of the nucleus.

• Dynamic effect which arises due to the strong interaction of the initial nucleon in the nuclear medium.

*In a nuclear medium for em interaction the expression for the cross section is written as:* 

$$\frac{d^2 \sigma^A}{d\Omega' dE'} = \frac{\alpha^2}{q^4} \frac{|\vec{k}'|}{|\vec{k}|} L^{\mu\nu} W^A_{\mu\nu},$$

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#### Nuclear hadronic tensor:

$$W_{\mu\nu}^{A} = \left(\frac{q_{\mu}q_{\nu}}{q^{2}} - g_{\mu\nu}\right) W_{1}^{A}(\nu, Q^{2}) + \frac{W_{2}^{A}(\nu, Q^{2})}{M_{A}^{2}} \left(p_{\mu} - \frac{p.q}{q^{2}} q_{\mu}\right) \left(p_{\nu} - \frac{p.q}{q^{2}} q_{\nu}\right)$$

 $W_i^A(\mathbf{v}, Q^2)$  are redefined as:

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# • We use a relativistic nucleon spectral function to describe the momentum distribution of nucleons in nuclei.

- The spectral function has been calculated using Lehmann's representation for the relativistic nucleon propagator.
- Nuclear many body theory is used to calculate it for an interacting Fermi sea in nuclear matter.
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The probability per unit time for the incoming lepton to collide with nucleons when traveling through nuclear matter:

$$\Gamma = -\frac{2m_l}{|\vec{k}|} \mathrm{Im}\Sigma$$

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The cross section  $\sigma$  for lepton-scattering from an element of volume  $d^3r$  and surface dS in the nucleus:

$$\sigma = \Gamma dt dS = \Gamma \frac{dt}{dl} dS dl = \Gamma \frac{1}{v} dV$$

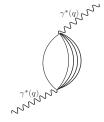
$$= \Gamma \frac{E_l(\vec{k})}{|\vec{k}|} dV = -\frac{2m_l}{|\vec{k}|} \operatorname{Im} \Sigma d^3 r$$

$$\stackrel{l^{-}(k)}{\underset{l^{-}(k)}{}} \bigvee_{N(p)} \bigvee_{X(p')} X(p')$$

d

*Lepton self energy*  $\Sigma(k)$  *for em interaction is written as:* 

$$-i\Sigma(k) = \int \frac{d^4q}{(2\pi)^4} \bar{u}_l(\vec{k}) ie\gamma^{\mu} i \frac{k'+m}{k'^2 - m^2 + i\varepsilon} ie\gamma^{\nu} u_l(\vec{k}) \frac{-ig_{\mu\rho}}{q^2} (-i) \Pi^{\rho\sigma}(q) \frac{-ig_{\sigma\nu}}{q^2}$$



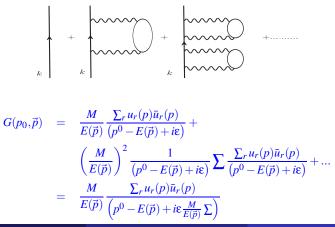
photon self-energy  $\Pi^{\mu\nu}(q)$  in the nuclear medium:

$$\begin{split} \Pi^{\mu\nu}(q) &= e^2 \int \frac{d^4p}{(2\pi)^4} G(p) \sum_X \sum_{s_p, s_l} \prod_{i=1}^N \int \frac{d^4p'_i}{(2\pi)^4} \prod_l G_l(p'_l) \prod_j D_j(p'_j) \\ &< X |J^{\mu}| H > < X |J^{\nu}| H >^* (2\pi)^4 \, \delta^4(q+p-\sum_{i=1}^N p'_i) \end{split}$$

Relativistic Dirac propagator  $G^0(p_0, \vec{p})$  for a free nucleon:

$$G^{0}(p_{0},\vec{p}) = \frac{M}{E(\vec{p})} \left\{ \frac{\sum_{r} u_{r}(p)\bar{u}_{r}(p)}{p^{0} - E(\vec{p}) + i\varepsilon} + \frac{\sum_{r} v_{r}(-p)\bar{v}_{r}(-p)}{p^{0} + E(\vec{p}) - i\varepsilon} \right\}$$

The nucleon propagator in the interacting Fermi sea is obtained by making a perturbative expansion of  $G(p^0, p)$  in terms of  $G^0(p^0, p)$  by retaining the positive energy contributions only:



This allows us to write the relativistic nucleon propagator in a nuclear medium in terms of the spectral functions of holes and particles as:

$$G(p^{0},\vec{p}) = \frac{M}{E(\vec{p})} \sum_{r} u_{r}(\vec{p}) \bar{u}_{r}(\vec{p}) \left[ \int_{-\infty}^{\mu} d\omega \frac{S_{h}(\omega,\vec{p})}{p^{0} - \omega - i\varepsilon} + \int_{\mu}^{\infty} d\omega \frac{S_{p}(\omega,\vec{p})}{p^{0} - \omega + i\varepsilon} \right]$$

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for  $p^{0} \leq \mu$ 

$$S_{h}(p^{0},\vec{p}) = \frac{1}{\pi} \frac{\frac{M}{E(\vec{p})} Im\Sigma(p^{0},\vec{p})}{(p^{0} - E(\vec{p}) - \frac{M}{E(\vec{p})} Re\Sigma(p^{0},\vec{p}))^{2} + (\frac{M}{E(\vec{p})} Im\Sigma(p^{0},\vec{p}))^{2}}$$

for  $p^0 > \mu$ 

$$S_{p}(p^{0},\vec{p}) = -\frac{1}{\pi} \frac{\frac{M}{E(\vec{p})} Im\Sigma(p^{0},\vec{p})}{(p^{0} - E(\vec{p}) - \frac{M}{E(\vec{p})} Re\Sigma(p^{0},\vec{p}))^{2} + (\frac{M}{E(\vec{p})} Im\Sigma(p^{0},\vec{p}))^{2}}$$

P. Fernandez de Cordoba and E. Oset, Phys. Rev. C 46 (1992) 1697E. Marco, E. Oset and P. Fernandez de Cordoba, Nucl. Phys. A 611 (1996) 484

Farhana Zaidi (AMU, India)

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## Local Density Approximation

In the local density approximation reaction takes place at a point *r*, lying inside a volume  $d^3r$  with local density  $\rho_p(r)$  and  $\rho_n(r)$  corresponding to the proton and neutron densities

$$\rho_p(r) = \frac{Z}{A}\rho(r)$$
  

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This leads to the spectral functions for the protons and neutrons to be the function of local Fermi momentum given by

$$2\int \frac{d^3p}{(2\pi)^3} \int_{-\infty}^{\mu} S_h(\omega, \vec{p}, p_{F_{p,n}}(\vec{r})) d\omega = \rho_{p,n}(\vec{r})$$
$$4\int d^3r \int \frac{d^3p}{(2\pi)^3} \int_{-\infty}^{\mu} S_h(\omega, \vec{p}, \rho(r)) d\omega = A$$

#### Nuclear hadronic tensor:

In the LDA, the nuclear hadronic tensor can be written as a convolution of nucleonic hadronic tensor with the hole spectral function

$$W^{A}_{\alpha\beta} = 4 \int d^{3}r \int \frac{d^{3}p}{(2\pi)^{3}} \int_{-\infty}^{\mu} dp^{0} \frac{M}{E(\vec{p})} S_{h}(p^{0}, \vec{p}, \rho(r)) W^{N}_{\alpha\beta}(p, q)$$

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#### Taking the xx component

$$W_{xx}^{N} = \left(\frac{q_{x}q_{x}}{q^{2}} - g_{xx}\right) W_{1}^{N} + \frac{1}{M^{2}} \left(p_{x} - \frac{p \cdot q}{q^{2}} q_{x}\right) \left(p_{x} - \frac{p \cdot q}{q^{2}} q_{x}\right) W_{2}^{N}$$

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Chosing  $\vec{q}$  along the z-axis

$$W_{xx}^{N}(\mathbf{v}_{N},Q^{2}) = W_{1}^{N}(\mathbf{v}_{N},Q^{2}) + \frac{1}{M^{2}}p_{x}^{2}W_{2}^{N}(\mathbf{v}_{N},Q^{2})$$

Similarly taking xx component of nuclear hadronic tensor

$$W_{xx}^{A}(\mathbf{v}_{A}, Q^{2}) = W_{1}^{A}(\mathbf{v}_{A}, Q^{2}) = \frac{F_{1}^{A}(x_{A})}{AM}$$

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# $F_1(x) = M W_1(v, Q^2), F_2(x) = v W_2(v, Q^2)$

$$\frac{F_{1}^{A}(x_{A})}{AM} = 4 \int d^{3}r \int \frac{d^{3}p}{(2\pi)^{3}} \frac{M}{E(\vec{p})} \int_{-\infty}^{\mu} dp^{0} S_{h}(p^{0}, \vec{p}, \rho(\vec{r})) \times \left[\frac{F_{1}^{N}(x_{N})}{M} + \frac{1}{M^{2}} p_{x}^{2} \frac{F_{2}^{N}(x_{N})}{\nu}\right]$$

By using Callan-Gross relation  $2xF_1(x) = F_2(x)$ 

$$F_{2}^{A}(x_{A}) = 4 \int d^{3}r \int \frac{d^{3}p}{(2\pi)^{3}} \frac{M}{E(\vec{p})} \int_{-\infty}^{\mu} dp^{0} S_{h}(p^{0}, \vec{p}, \rho(\vec{r})) \frac{x}{x_{N}}$$
$$\times \left(1 + \frac{2x_{N}p_{x}^{2}}{M\nu_{N}}\right) F_{2}^{N}(x_{N})$$

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$$\times \left(1 + \frac{2x_{N}p_{x}^{2}}{Mv_{N}}\right) F_{2}^{N}(x_{N})$$

where

$$x_A = \frac{x}{A},$$
  

$$x_N = \frac{Q^2}{2(p^0q^0 - p_zq_z)}$$

## Effect of nuclear medium on Callan Gross relation

#### To obtain $F_2(x)$ independently

$$W_{zz}^{N} = \left(\frac{q_{z}^{2}}{q^{2}} - g_{zz}\right) W_{1}^{N} + \frac{1}{M^{2}} \left(p_{z} - \frac{p \cdot q}{q^{2}} q_{z}\right)^{2} W_{2}^{N}$$
$$= \frac{q_{0}^{2}}{q^{2}} W_{1}^{N} + \frac{1}{M^{2}} \left(\frac{(p_{z}q^{2} - p \cdot q q_{z})^{2}}{q^{4}}\right) W_{2}^{N}$$

$$W_{zz}^{A}(\mathbf{v}_{A},Q^{2}) = \left(\frac{q_{z}^{2}}{q^{2}} - g_{zz}\right) W_{1}^{A} + \frac{1}{M_{A}^{2}} \left(-\frac{p_{A} \cdot q}{q^{2}} q_{z}\right)^{2} W_{2}^{A}$$

$$F_2(x) = \mathbf{v} W_2(\mathbf{v}, Q^2)$$

#### Finally, we obtain

$$F_{2}^{A}(x_{A}) = 2\sum_{p,n} \int d^{3}r \int \frac{d^{3}p}{(2\pi)^{3}} \frac{M}{E(\vec{p})} \int_{-\infty}^{\mu} dp^{0} S_{h}^{p,n}(p^{0},\vec{p},\rho_{p,n}(\vec{r})) F_{2}^{N}(x_{N})$$

$$\times \left[ \frac{Q^{2}}{q_{z}^{2}} \left( \frac{p^{2} - p_{z}^{2}}{2M^{2}} \right) + \frac{(p.q)^{2}}{M^{2}\nu^{2}} \left( \frac{p_{z}}{p.qq_{z}} \frac{Q^{2}}{p} + 1 \right)^{2} \frac{q_{0}M}{p_{0}q_{0} - p_{z}q_{z}} \right]$$

This expression of  $F_2^A(x_A)$  is obtained without applying Callan-Gross relation.

# $\pi$ and $\rho$ mesons contribution to the nuclear structure function

For mesons cloud contribution

$$2\pi \frac{M}{E(\vec{p})} S_h(p_0,\vec{p}) W_N^{\alpha\beta}(p,q) \to 2ImD(p)\theta(p_0) W_\pi^{\alpha\beta}(p,q)$$

Pion propagator in the nuclear medium

$$D(p) = [p_0^2 - \vec{p}^2 - m_{\pi}^2 - \Pi_{\pi}(p_0, \vec{p})]^{-1}$$

with

$$\Pi_{\pi} = \frac{f^2 / m_{\pi}^2 F^2(p) \vec{p}^2 \Pi^*}{1 - f^2 / m_{\pi}^2 V_L' \Pi^*}$$

 $\pi NN$  form factor

$$F(p) = (\Lambda^2 - m_\pi^2)/(\Lambda^2 + \vec{p}^2)$$

Similar to nucleonic case

$$W_{A,\pi}^{\mu\nu} = 3 \int d^3r \int \frac{d^4p}{(2\pi)^4} \,\theta(p_0)(-2) \, ImD(p) \, 2m_\pi W_\pi^{\mu\nu}(p,q)$$

Factor  $3 \Rightarrow$  Three Charged states of pion

For pion excess in nuclear medium

$$ImD(p) \rightarrow \delta ImD(p) \equiv ImD(p) - \rho \frac{\partial ImD(p)}{\partial \rho} \Big|_{\rho=0}$$

which leads to

$$F_{1,\pi}^{A}(x_{\pi}) = -6AM \int d^{3}r \int \frac{d^{4}p}{(2\pi)^{4}} \Theta(p_{0}) \, \delta ImD(p) \, 2m_{\pi} \times \left[ \frac{F_{1\pi}(x_{\pi})}{m_{\pi}} + \frac{|\vec{p}|^{2} - p_{z}^{2}}{2(p_{0} q_{0} - p_{z} q_{z})} \frac{F_{2\pi}(x_{\pi})}{m_{\pi}} \right]$$

#### Contribution from rho meson

Propagator for rho meson

$$D_{\rho}(p) = [p_0^2 - \vec{p}^2 - m_{\rho}^2 - \Pi_{\rho}^*(p_0, \vec{p})]^{-1}$$

with irreducible  $\rho$  self-energy

$$\Pi_{\rho}^{*} = \frac{f^{2}/m_{\rho}^{2}C_{\rho}F_{\rho}^{2}(p)\vec{p}^{2}\Pi^{*}}{1 - f^{2}/m_{\rho}^{2}V_{T}'\Pi^{*}}$$

 $\rho NN$  form factor

$$F_{\rho}(p) = (\Lambda_{\rho}^2 - m_{\rho}^2)/(\Lambda_{\rho}^2 + \vec{p}^2)$$

Finally,

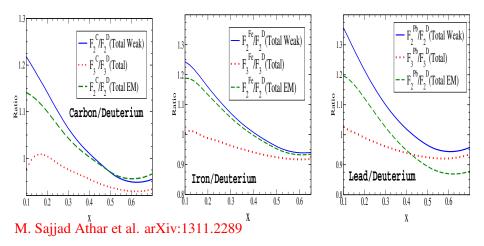
$$F_{1,\rho}^{A}(x_{\rho}) = -12AM \int d^{3}r \int \frac{d^{4}p}{(2\pi)^{4}} \Theta(p_{0}) \, \delta ImD_{\rho}(p) \, 2m_{\rho} \times \\ \left[ \frac{F_{1\rho}(x_{\rho})}{m_{\rho}} + \frac{|\vec{p}|^{2} - p_{z}^{2}}{2(p_{0} q_{0} - p_{z} q_{z})} \frac{F_{2\rho}(x_{\rho})}{m_{\rho}} \right]$$

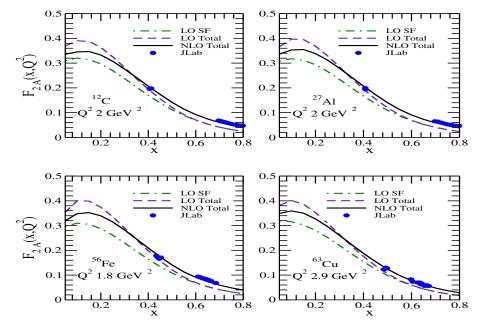
Structure functions for  $\pi$  and  $\rho$  mesons without using Callan-Gross relation:

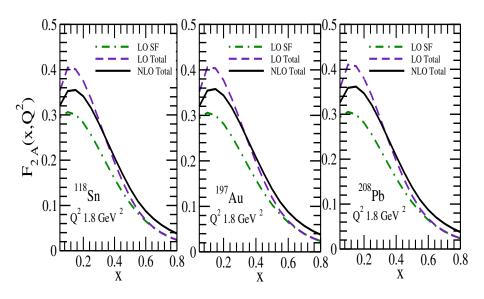
$$F_{2,\pi}^{A}(x_{\pi}) = -6 \int d^{3}r \int \frac{d^{4}p}{(2\pi)^{4}} \theta(p_{0}) \, \delta ImD(p) \, 2m_{\pi} \, \frac{m_{\pi}}{p_{0} - p_{z} \, \gamma} \times \\ \left[ \frac{Q^{2}}{q_{z}^{2}} \left( \frac{|\vec{p}|^{2} - p_{z}^{2}}{2m_{\pi}^{2}} \right) + \frac{(p_{0} - p_{z} \, \gamma)^{2}}{m_{\pi}^{2}} \left( \frac{p_{z} \, Q^{2}}{(p_{0} - p_{z} \, \gamma)q_{0}q_{z}} + 1 \right)^{2} \right] F_{2\pi}(x_{\pi})$$

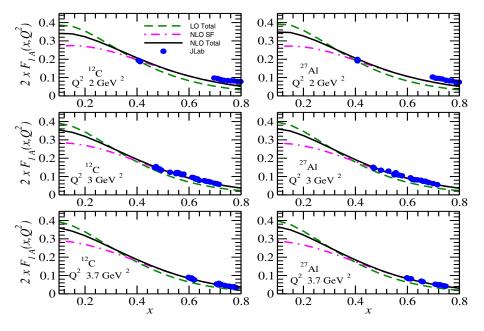
$$F_{2,\rho}^{A}(x_{\rho}) = -12 \int d^{3}r \int \frac{d^{4}p}{(2\pi)^{4}} \theta(p_{0}) \, \delta Im D_{\rho}(p) \, 2m_{\rho} \, \frac{m_{\rho}}{p_{0} - p_{z} \, \gamma} \times \\ \left[ \frac{Q^{2}}{q_{z}^{2}} \left( \frac{|\vec{p}|^{2} - p_{z}^{2}}{2m_{\rho}^{2}} \right) + \frac{(p_{0} - p_{z} \, \gamma)^{2}}{m_{\rho}^{2}} \left( \frac{p_{z} \, Q^{2}}{(p_{0} - p_{z} \, \gamma)q_{0}q_{z}} + 1 \right)^{2} \right] F_{2\rho}(x_{\rho})$$

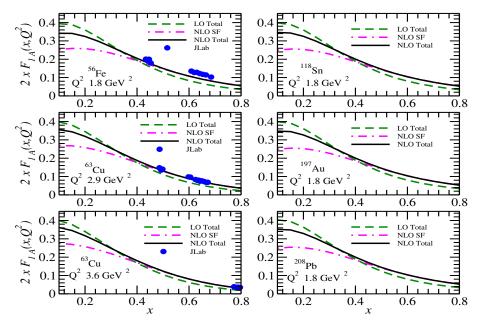
### Ratio of Structure functions in Weak and E.M. cases



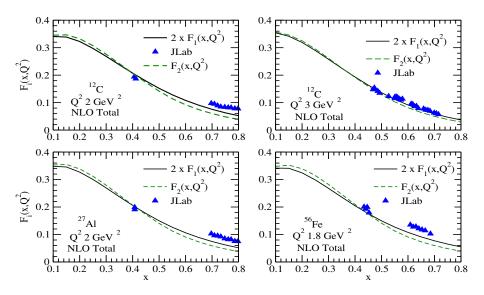


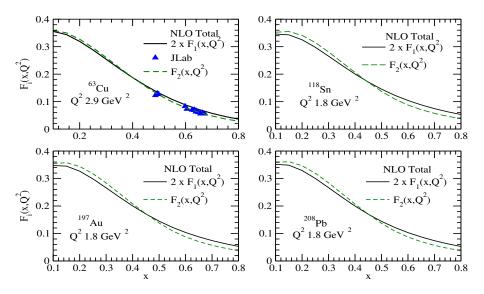


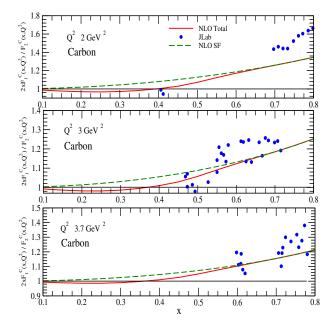




NPA 943 58 (2015), V. Mamyan arXiv:1202.1457

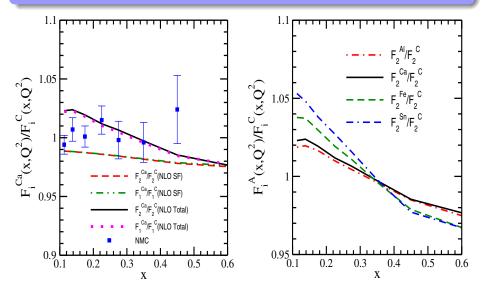






NPA 943 58 (2015), V. Mamyan arXiv:1202.1457

Nuclear dependence in  $\frac{F_i^A(x,Q^2)}{F_i^C(x,Q^2)}$ 



NPA 943 58 (2015), Nucl. Phys. B 441 3 (1995)

# *Percantage difference for* $F_2$ *in different nuclei*

Nucleus	$Q^2(GeV^2)$	$\left \frac{LO_{SF}-LO_{T}}{LO_{T}}\right \%$	$\left \frac{LO_T - NLO_T}{NLO_T}\right \%$	$ \frac{NLO_T - NLO_{WOS}}{NLO_{WOS}} \%$
<sup>12</sup> C(x=0.2)	2	18	9	≈2
(x=0.4)		12	6	1
(x=0.8)		0.16	38	0.25
<sup>27</sup> Al(x=0.2)	2	20	≈9	2
(x=0.4)		14	6	1
(x=0.8)		0.2	38	0.3
<sup>56</sup> Fe(x=0.2)	1.8	23	10	2
(x=0.4)		17	5	1.3
(x=0.8)		0.24	37	0.37
<sup>63</sup> Cu(x=0.2)	2.9	20	7.7	1.9
(x=0.4)		12	6.5	1
(x=0.8)		0.18	38.7	0.3
<sup>118</sup> Sn(x=0.2)	1.8	24.6	9.8	2.3
(x=0.4)		18.5	5.3	1.3
(x=0.8)		0.26	37	0.4
<sup>197</sup> Au(x=0.2)	1.8	25	9.7	2.6
(x=0.4)		19.8	5	1.5
(x=0.8)		0.3	37	0.46
<sup>208</sup> Pb(x=0.2)	1.8	26	9.6	2.4
(x=0.4)		19	5	1.4
(x=0.8)		0.28	37	0.43

# *Percantage difference for* $|\frac{2xF_1-F_2}{2xF_1}|$ *in different nuclei*

Nucleus	$Q^2(GeV^2)$	$\left \frac{2xF_{1}-F_{2}}{2xF_{1}}\right \%$
<sup>12</sup> C(x=0.2)	2	≈ 3
(x=0.7)		$\approx 21$
(x=0.2)	3	≈ 2
(x=0.7)		pprox 16
(x=0.2)	3.7	1.4
(x=0.7)		$\approx 13$
<sup>56</sup> Fe(x=0.2)	1.8	$\approx 5$
(x=0.7)		$\approx 23$
(x=0.2)	5	1.3
(x=0.7)		11
(x=0.2)	10	$\approx 0.6$
(x=0.7)		$\approx 6$
(x=0.2)	20	≈ 0.3
(x=0.7)		≈ 3
<sup>63</sup> Cu(x=0.2)	2.9	2.5
(x=0.7)		$\approx 17$
(x=0.2)	3.6	2
(x=0.7)		$\approx 14$

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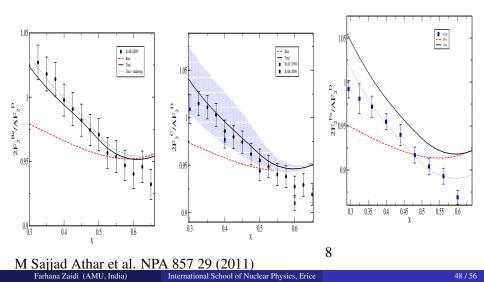
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- We compare our results with the JLab and NMC data and found them in good agreement.
- The present work will make useful predictions for the future experiments in the low *x* and moderate  $Q^2$ .



# **BACK UP SLIDES**

# Electromagnetic Nuclear Structure Function $\frac{2F_2^A}{AF_2^D}(A = Be, C, Fe)$ vs x



# Theoretical Study

Along with the experimental efforts theoretical groups also performed calculations in order to understand the nuclear medium effects for the e.m. DIS process.

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#### Aligarh group:

NPA 943 58 (2015)

NPA 940 138 (2015)

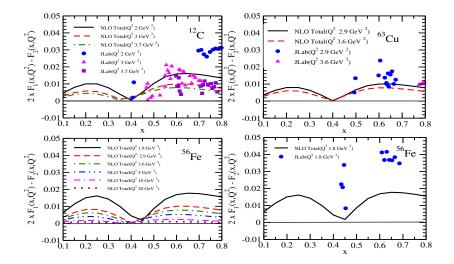
PRC 87 035502 (2013)

PRC 85 055201 (2012)

PRC 84 054610 (2011)

NPA 857 29 (2011)

PLB 668 133 (2008)



"Significant at low-x and low- $Q^2$ "

The shadowing suppression at small x occurs due to coherent multiple scattering inside the nucleus of a quark-anti quark pair coming from the virtual boson with destructive interference of the amplitudes. The shadowing effect is important at low x and low  $Q^2$ .

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For the shadowing and antishadowing effects, Glauber-Gribov multiple scattering model has been used following the work of Kulagin and Petti. PRD76(2007)094033.

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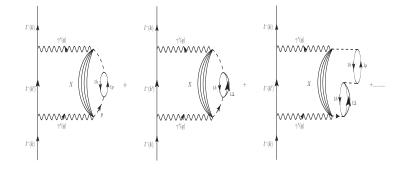
"Significant at low-x and low- $Q^2$ "

The shadowing suppression at small x occurs due to coherent multiple scattering inside the nucleus of a quark-anti quark pair coming from the virtual boson with destructive interference of the amplitudes. The shadowing effect is important at low x and low  $Q^2$ .

The anti-shadowing effect is due to constructive interference of the multiple scattering amplitudes. This effect is also important at low x but greater than x region of shadowing.

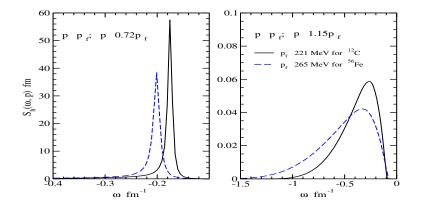
For the shadowing and antishadowing effects, Glauber-Gribov multiple scattering model has been used following the work of Kulagin and Petti. PRD76(2007)094033.

# $\pi$ and $\rho$ mesons contribution to the nuclear structure function



- More effective for heavier nuclei
- 2 Contributes to the intermediate *x* region
- Implemented following the many body field theoretical approach
- Using parametrization by Gluck et al.

# Nature of Spectral Function



- Behave like a  $\delta$  function for  $p < p_F$ ,
- Long range for  $p > p_F$ .

Hole and particle spectral function fulfills

$$\int_{-\infty}^{\mu} dp_0 \, S_h(p_0, \vec{p}) + \int_{\mu}^{\infty} dp_0 \, S_p(p_0, \vec{p}) = 1$$

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When interactions are not present

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When interactions are not present

$$egin{array}{rcl} E^N(p)&=&0\ G(p)& o&G^0(p) \end{array}$$

Therefore,

$$S_h(p_0, \vec{p}) = S_p(p_0, \vec{p}) = \delta(p_0 - E(\vec{p}))$$

$$\int_{-\infty}^{\mu} dp_0 \, S_h(p_0, \vec{p}) = \int_{-\infty}^{\mu} dp_0 \, \delta(p_0 - E(\vec{p})) = \begin{cases} 1 & \text{if } \mu > E(\vec{p}) \\ 0 & \text{if } \mu < E(\vec{p}) \end{cases}$$

$$\int_{\mu}^{\infty} dp_0 \, S_p(p_0, \vec{p}) = \int_{\mu}^{\infty} dp_0 \, \delta(p_0 - E(\vec{p})) = \begin{cases} 1 & \text{if } \mu < E(\vec{p}) \\ 0 & \text{if } \mu > E(\vec{p}) \end{cases}$$

# Chemical potential is defined as

 $\mu = M + \varepsilon_F$ and  $p_0 = M + \omega$ 

$$\begin{split} \int_{-\infty}^{\mu} dp_0 \, S_h(p_0, \vec{p}) &= \int_{-\infty}^{\mu-M} d\omega \, \delta(\omega + M - E(\vec{p})) \\ &= \begin{cases} 1 & ; \, \mu - M > E(\vec{p}) - M \Rightarrow \varepsilon_F > \varepsilon(\vec{p}) \\ 0 & ; \, \mu - M < E(\vec{p}) - M \Rightarrow \varepsilon_F < \varepsilon(\vec{p}) \end{cases} \\ \int_{\mu}^{\infty} dp_0 \, S_p(p_0, \vec{p}) &= \int_{\mu-M}^{\infty} d\omega \, \delta(\omega + M - E(\vec{p})) \\ &= \begin{cases} 1 & ; \, \mu - M < E(\vec{p}) - M \Rightarrow \varepsilon_F < \varepsilon(\vec{p}) \\ 0 & ; \, \mu - M > E(\vec{p}) - M \Rightarrow \varepsilon_F > \varepsilon(\vec{p}) \end{cases} \end{split}$$

# Hole spectral function

With Fermi energy

$$\varepsilon_F = \frac{p_F^2}{2M}$$

we have

$$\int_{-\infty}^{\mu} dp_0 \, S_h(p_0, \vec{p}) = \theta(p_F - |\vec{p}|) \equiv n_0(\vec{p})$$
$$\int_{\mu}^{\infty} dp_0 \, S_p(p_0, \vec{p}) = \theta(|\vec{p}| - p_F) \equiv 1 - n_0(\vec{p})$$

Hole spectral function is

- **()**  $\mathcal{P}$  of removing a nucleon from correlated ground state
- ②  $\mathcal{P}$  of finding the nucleons with an energy  $p_0 < E < p_0 + dp_0$