

Project 8: A frequency based approach to measure the neutrino mass

Martin Fertl

International School of Nuclear Physics

39th Course

Neutrinos in Cosmology, in Astro-, Particle- and Nuclear Physics

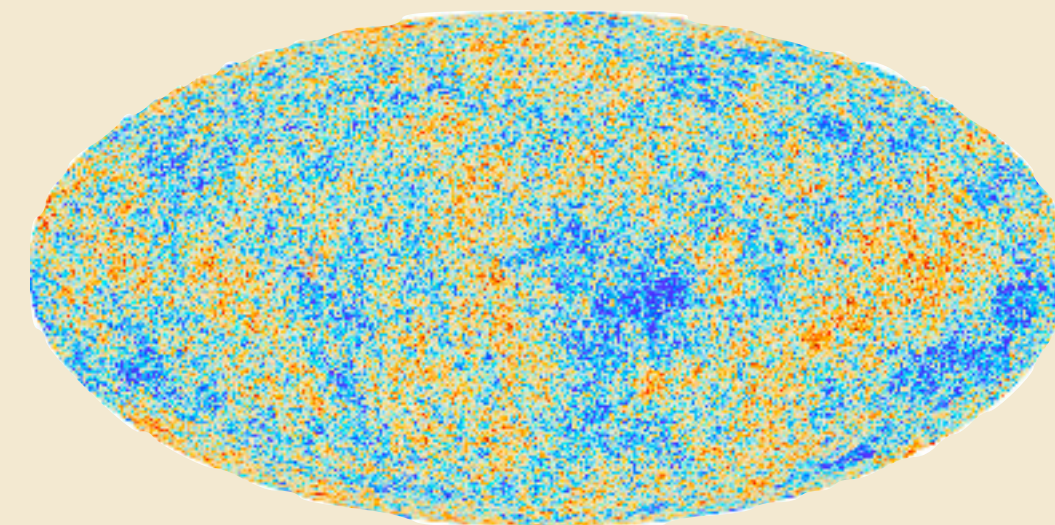
Erice-Sicily

September 16-24, 2017



Where to look for neutrino masses

M.C. Gonzalez-Garcia, K. Valerius and others



Three different
combination of
eigen masses

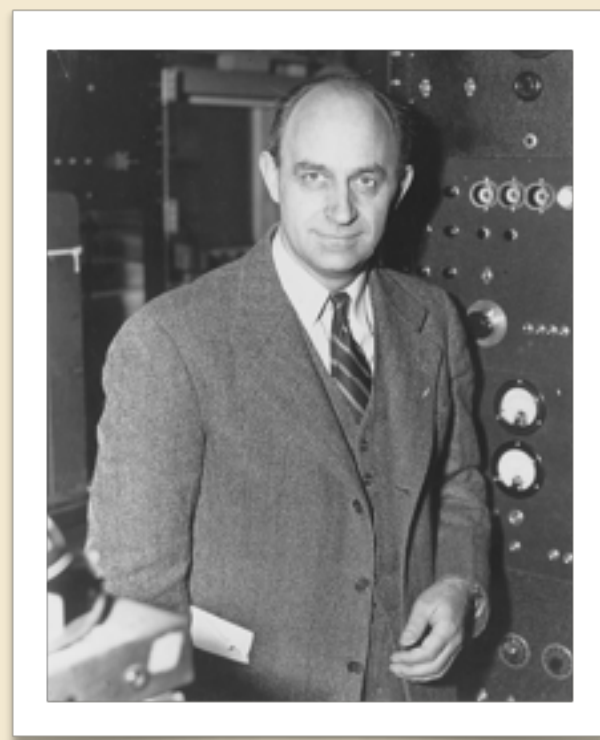
<i>Tool</i>	<i>Cosmology CMB + BAO + LSS</i>
<i>Observable</i>	$\sum m_\nu = \sum_{i=1}^3 m_i$
<i>Present best limit</i>	<i>0.15 - 1 eV</i>
<i>Potential reach</i>	<i>20-50 meV</i>
<i>Model dependence</i>	<i>Multi-parameter cosmological model</i>

Anti-electron neutrino mass and tritium beta decay spectrum

Finite neutrino mass modifies the decay electron spectrum!

Idealized situation:

1. Super-allowed β^- -decay of *isolated* atom
2. Single neutrino mass state



PHYSIQUE NUCLÉAIRE. — Possibilité d'émission de particules neutres de masse intrinsèque nulle dans les radioactivités β . Note de M. FRANCIS PERRIN, présentée par M. Jean Perrin.

SÉANCE DU 18 DÉCEMBRE 1933.

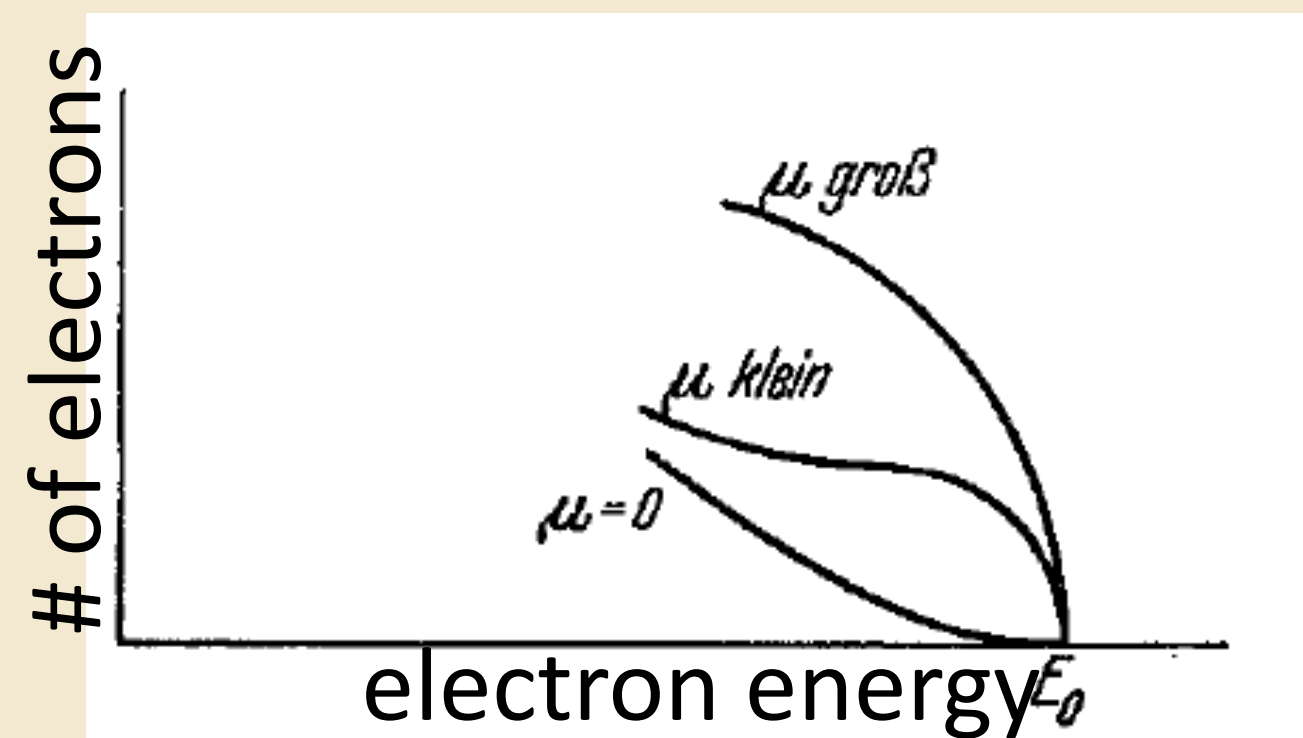


Versuch einer Theorie der β -Strahlen. I¹⁾.

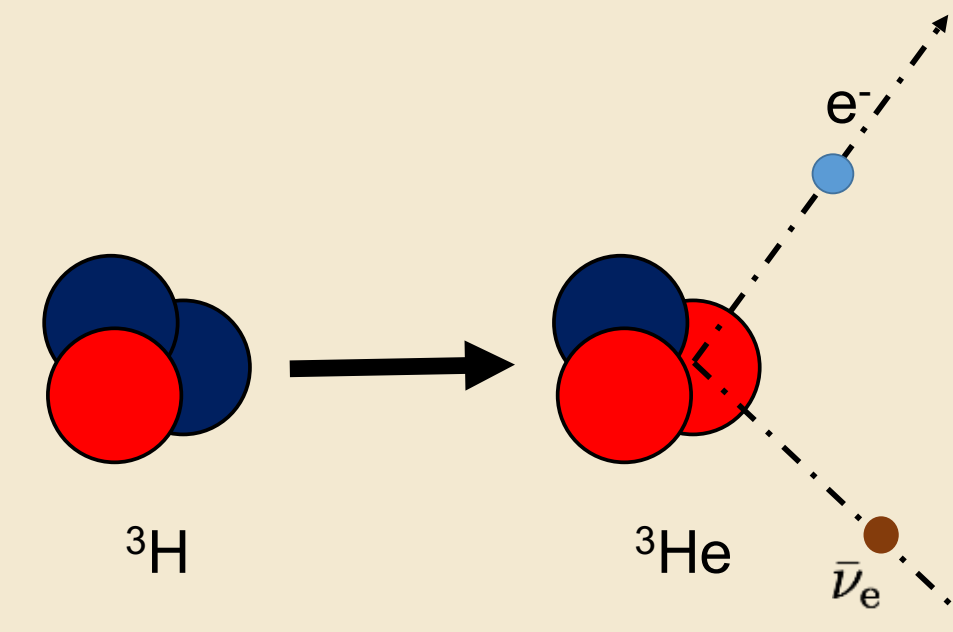
Von E. Fermi in Rom.

Mit 3 Abbildungen. (Eingegangen am 16. Januar 1934.)

Zeitschrift für Physik, Vol. 88, p. 161



Any new final state can modify the decay spectrum!
Sterile neutrinos!?



nuclear matrix element for T

$$\frac{dN}{dE_e} = \frac{G_F^2 m_e^5 \cos^2 \theta_C}{2\pi^3 \hbar^7} |M_{\text{nuc}}|^2 p \cdot (E + m_e) \cdot (E_0 - E) \cdot \sqrt{(E_0 - E)^2 - m^2(\nu_e)}$$

constants

purely kinematic variables

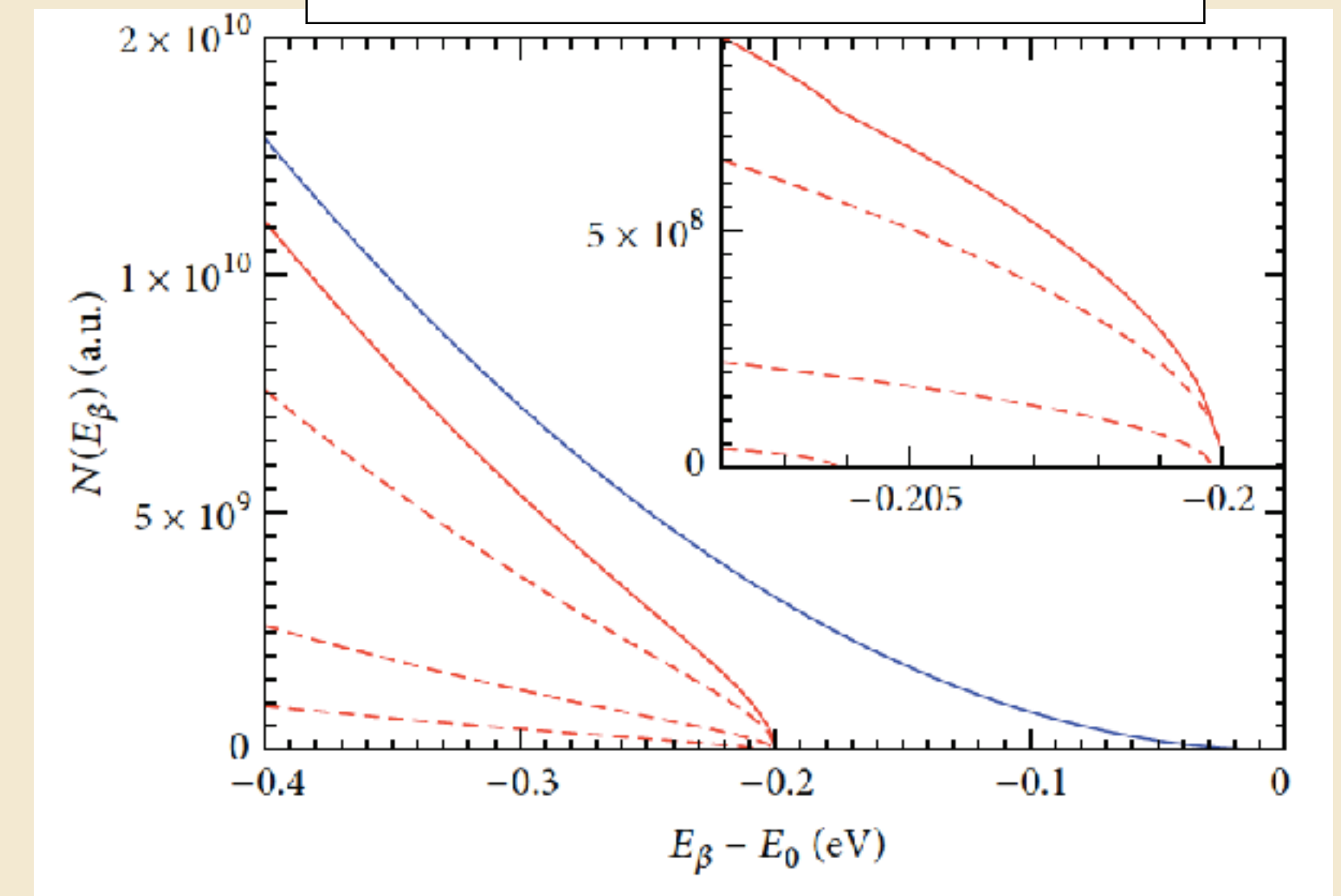
β^- -decay electron spectrum

With neutrino mixing and nuclear recoil for T_{nuc} :

$$\frac{dN}{dE_e} = \frac{G_F^2 m_e^5 \cos^2 \theta_C}{2\pi^3 \hbar^7} |M_{\text{nuc}}|^2 F(Z, E_e) p_e E_e \sum_i |U_{ei}|^2 (E_{\text{max}} - E_e) \times \sqrt{(E_{\text{max}} - E_e)^2 - m_{\nu i}^2} \cdot \Theta(E_{\text{max}} - E_e - m_{\nu i})$$

$$m(\nu_e) = \sqrt{\sum_i |U_{e,i}|^2 m_i^2} \quad BR \approx \left(\frac{\delta E}{E_0}\right)^3$$

For T: $m_\beta=0.0$ eV $m_\beta=0.2$ eV



Nucciotti, Advances in High Energy Physics, Vol. 2016, 9153024

Tritium

$Q(T_A) = 18.59201(7)$ keV
 Super allowed transition
 $T_{1/2} = 12.32$ y
 $BR(1\text{eV}) = 2 \times 10^{-13}$

^{187}Re

$Q = 2.4667(15)$ keV
 Unique first forbidden decay
 $T_{1/2} = 4 \times 10^{10}$ y
 $BR(1\text{eV}) = 7 \times 10^{-11}$

Experimentally disfavored

^{115}In to $^{115}\text{Sn}^*$

$Q = 0.173(12)$ keV
 Only 1.2×10^{-6} of the
 498 keV decay to
 ground state

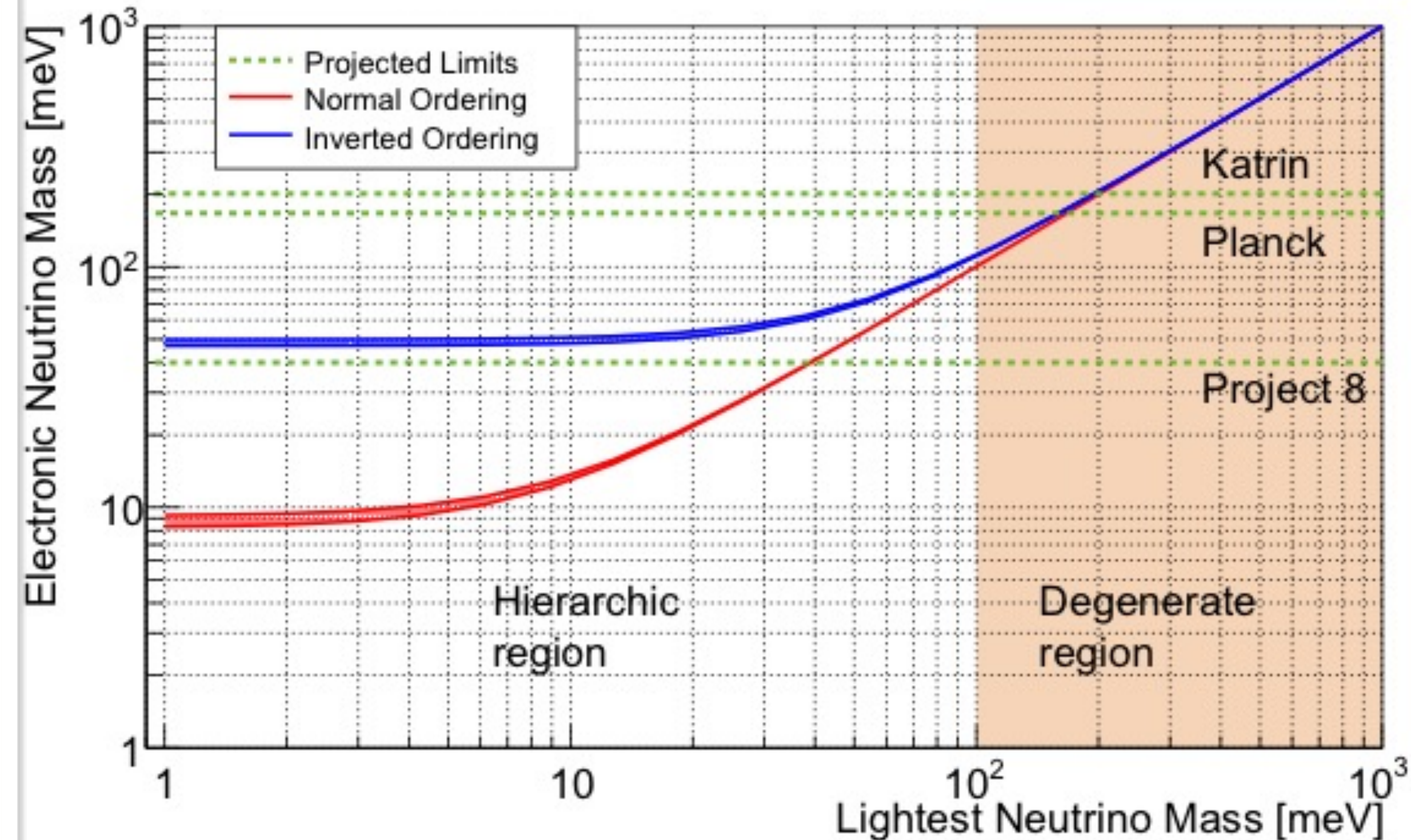
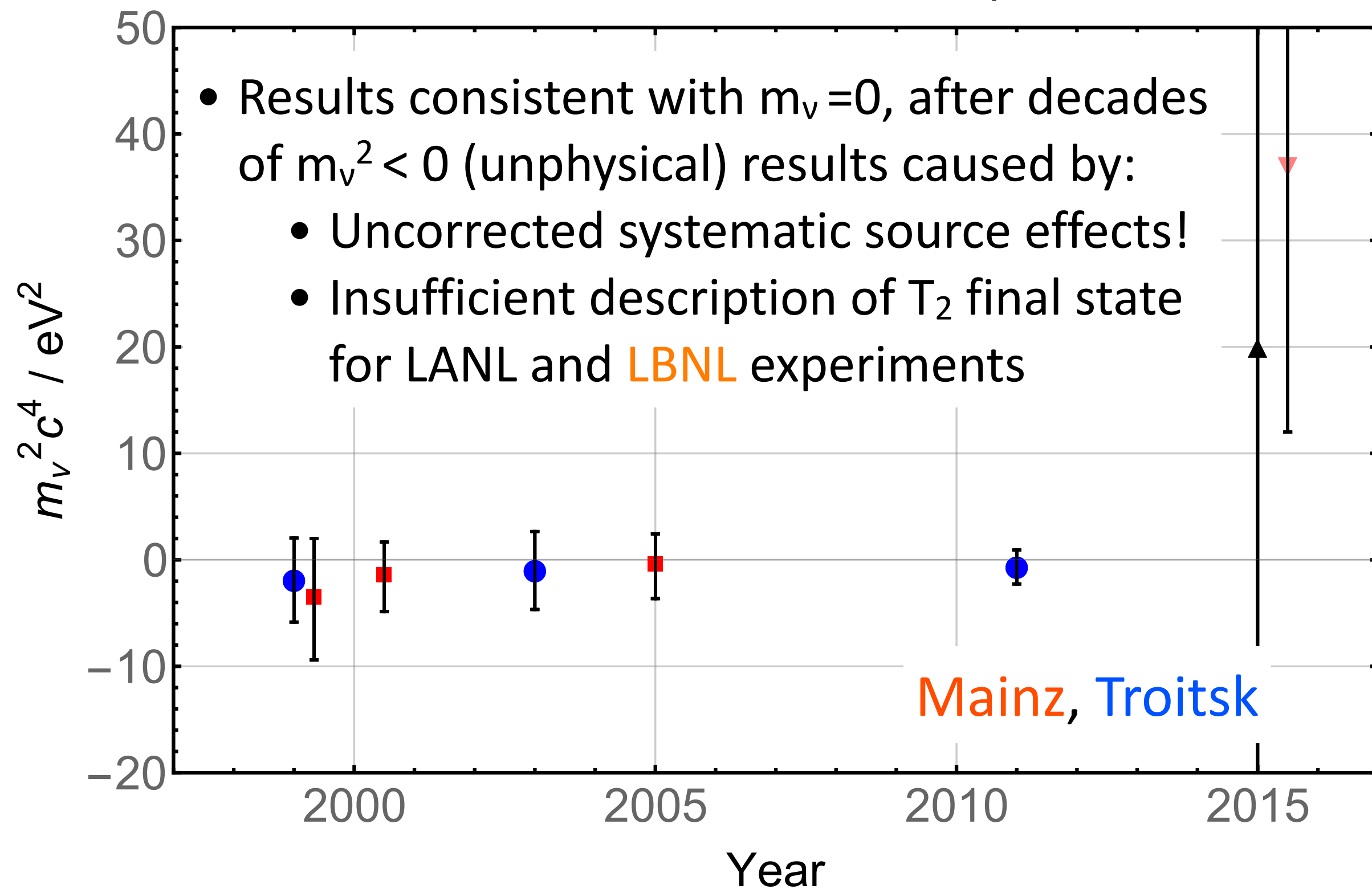
No experimental scheme devised!

For ^{163}Ho EC experiments:
 L. Gastaldo,
 A. Faessler,
 T. Kieck

Limits on anti-electron neutrino mass from tritium decay

Tritium has been the “workhorse” for decades, but ... :

The last two decades of tritium experiments ...



KATRIN: The largest MAC-E filter ever built



Magnetic Adiabatic Collimation with Electrostatic filter

A. Picard et al., Nucl. Instr. Meth. B 63 (1992)

$$\mu = \frac{E_{\perp}}{B} = \text{const.} \quad \frac{\Delta E}{E} = \frac{B_{\min}}{B_{\max}}$$

Retarding electric field:
Count decay e^{-} above threshold!

Experimental challenges:

- Highest source luminosity
- Highest energy resolution
- Control of inelastic scattering rate
- Safe tritium handling

First light in October 2016!

Tritium operation starting June 2018!

70 m total length!

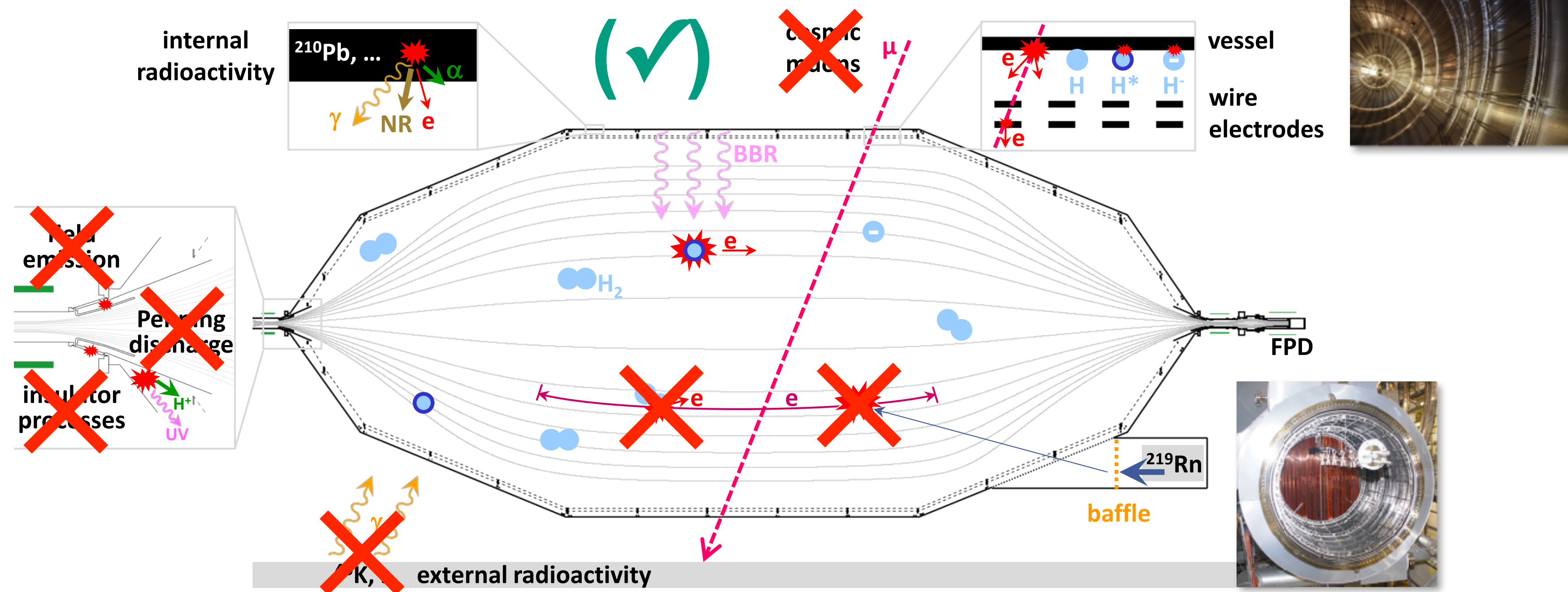


<http://www.katrin.kit.edu>

Karlsruhe Tritium Neutrino Experiment

KATRIN talks: K. Valerius, G. Drexlin, H. Seitz-Moskaliuk, S. Mirz, ...

Backgrounds in KATRIN



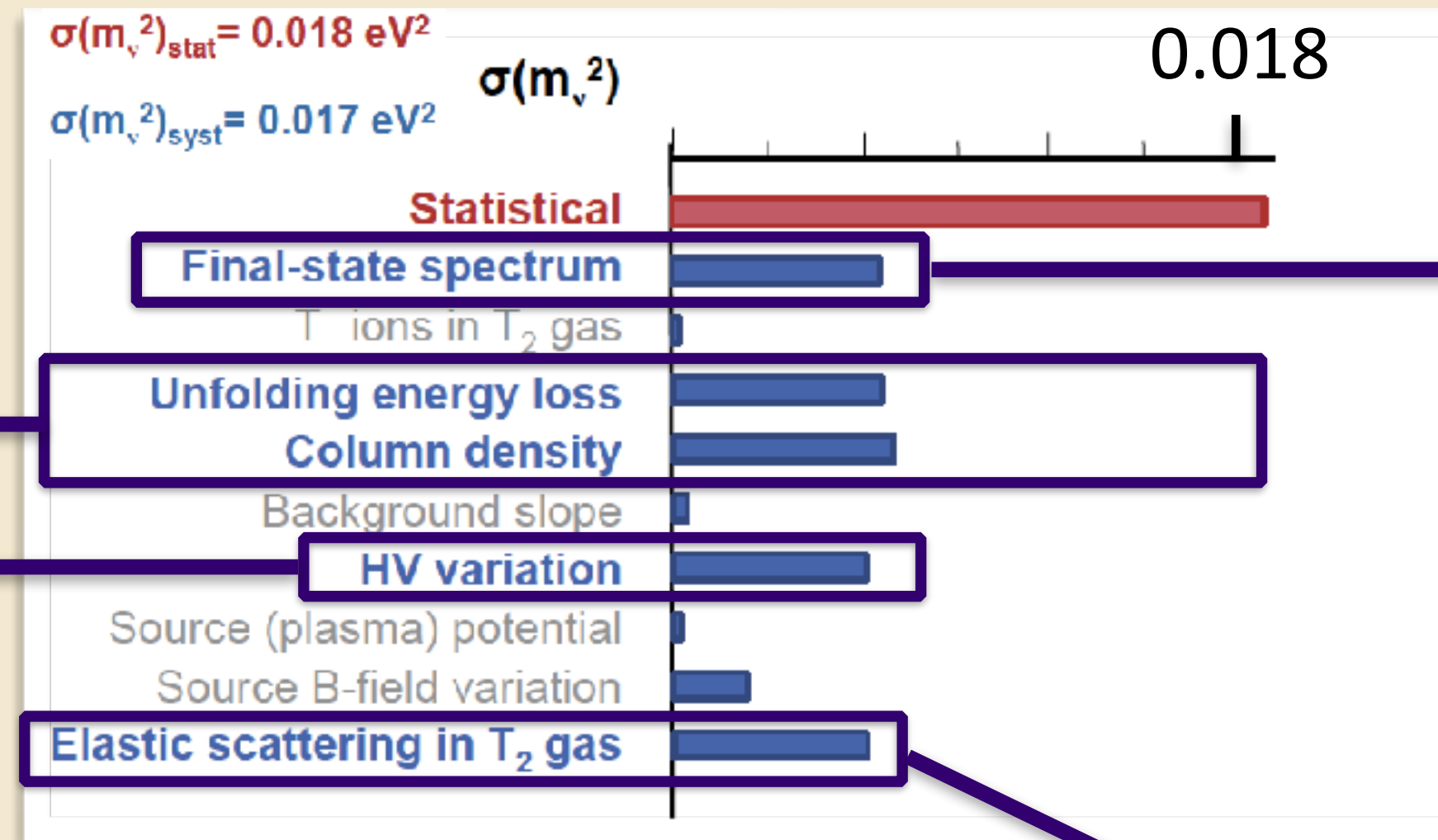
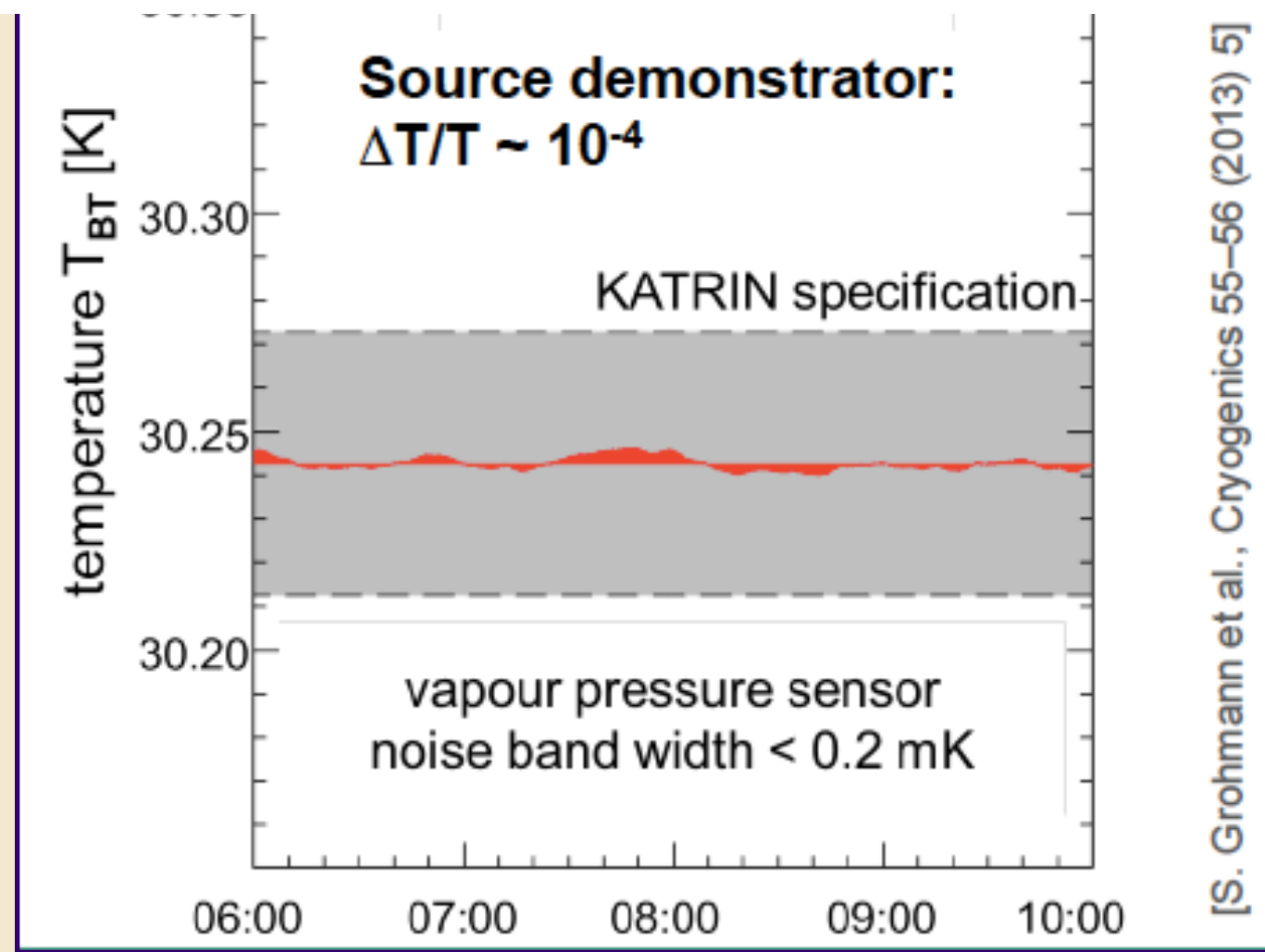
- **8 sources of background investigated and understood**
- **7 out of 8 avoided or actively eliminated by**
 - fine-shaping of special electrodes
 - inner electrode (wire grids on neg. potential)
 - symmetric magnetic fields
 - cold traps (LN₂-cooled baffles to remove ²¹⁹Rn)

- **1 out of 8 remaining:**
²¹⁰Pb on spectrometer walls (thermal ionisation of neutral H* atoms)
- Countermeasures:
 - extensive bake-out (done)
 - irradiation by strong UV source (ongoing investigation)



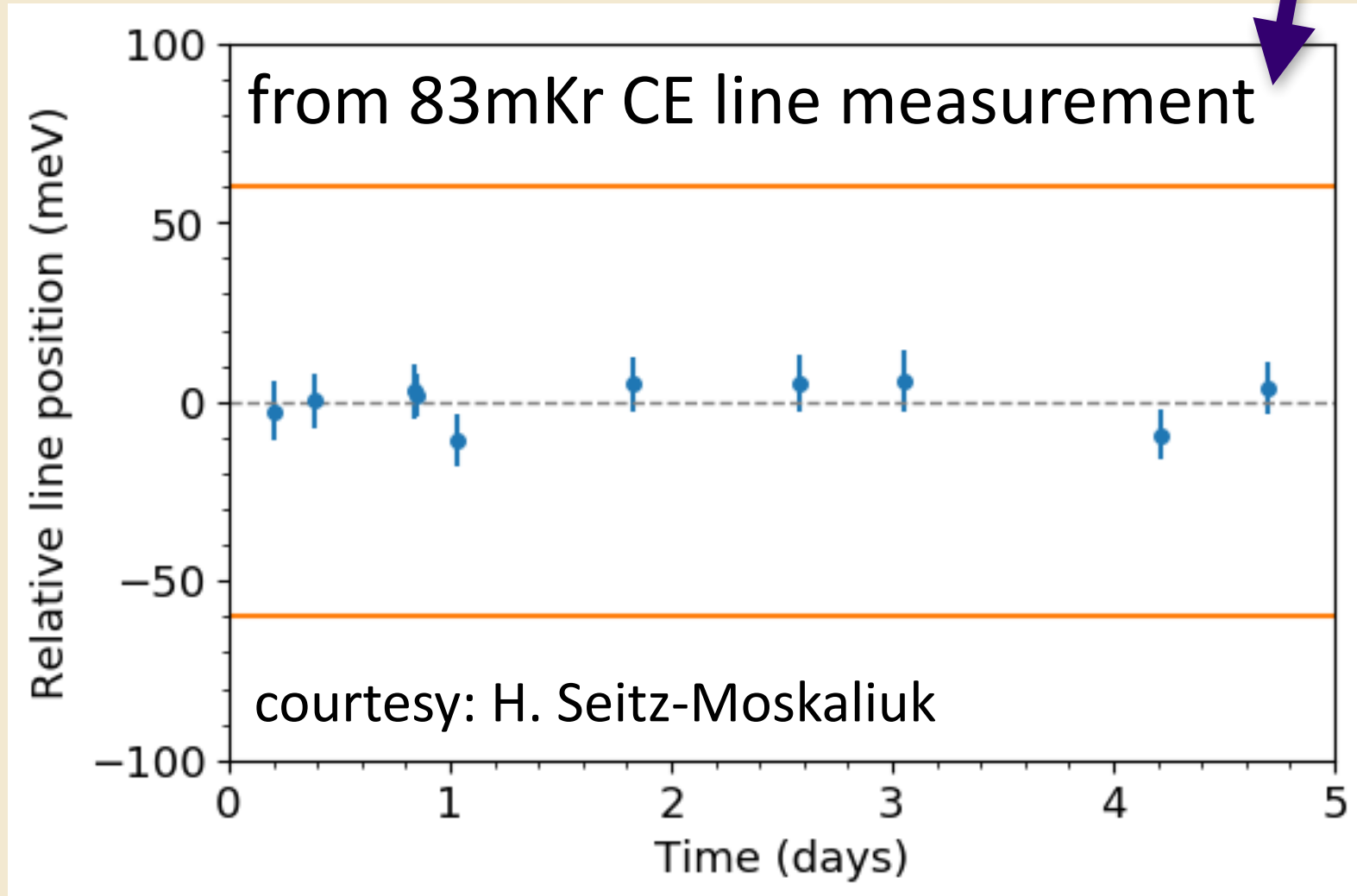
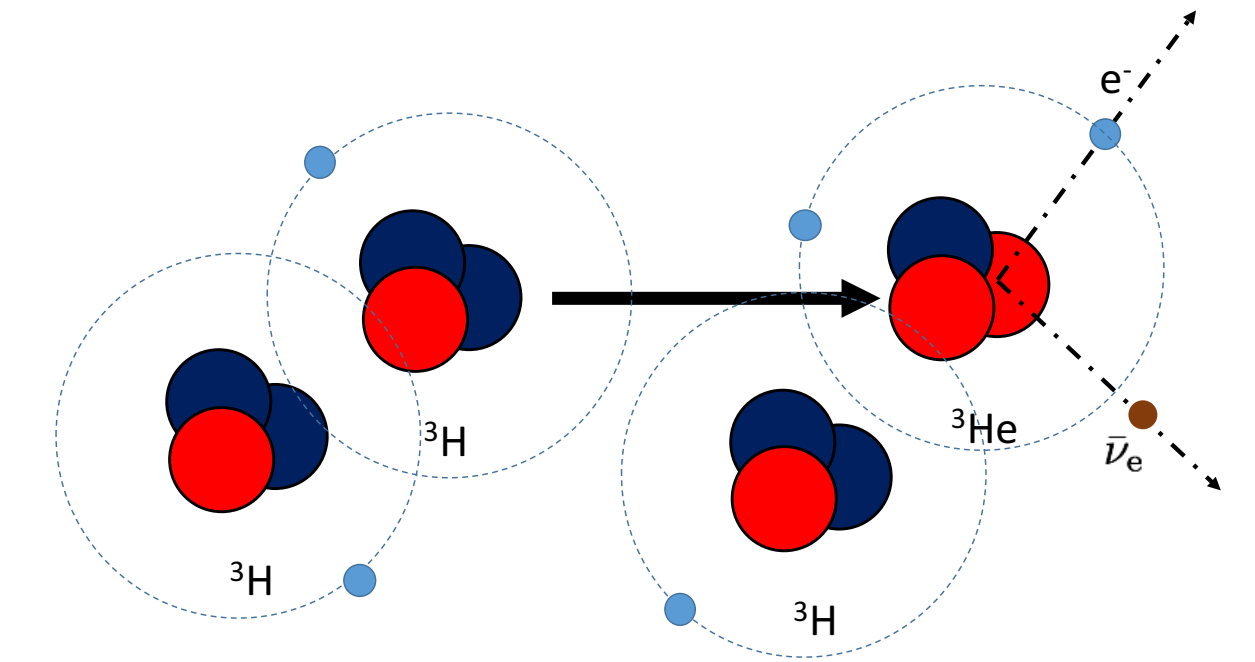
Current systematic limitations of a MAC-E filter

Excellent temperature stability



Kathrin Valerius, HEPHY Colloquium, Vienna, Jan 2016

Irreducible excitation of ro-vibrational initial and final states of T₂ molecule.

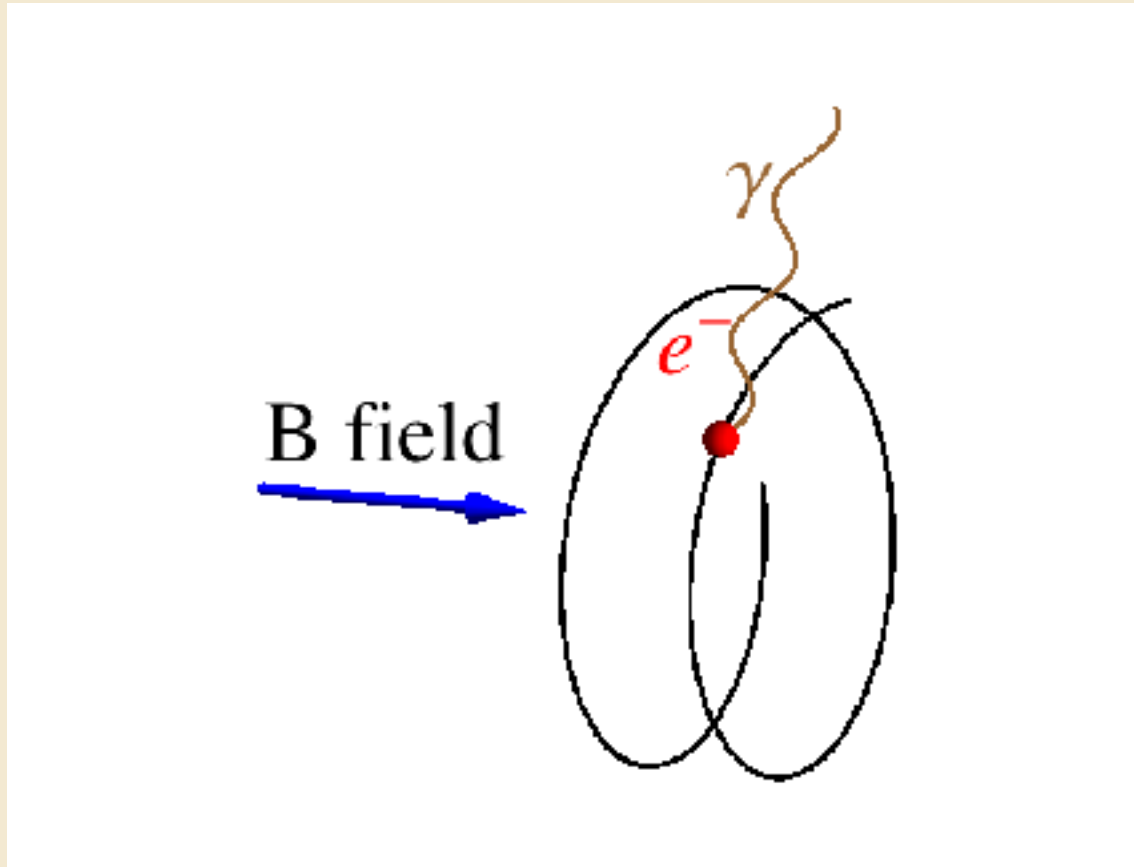


Big challenge for new techniques:

- Development of orthogonal set of systematic effects!
- Different classes of bkgds!
- Statistics!

1. e⁻ energy loss
2. e⁻ can be scattered into angular acceptance cone of the MAC-E filter

Project 8: Cyclotron radiation emission spectroscopy of T₂



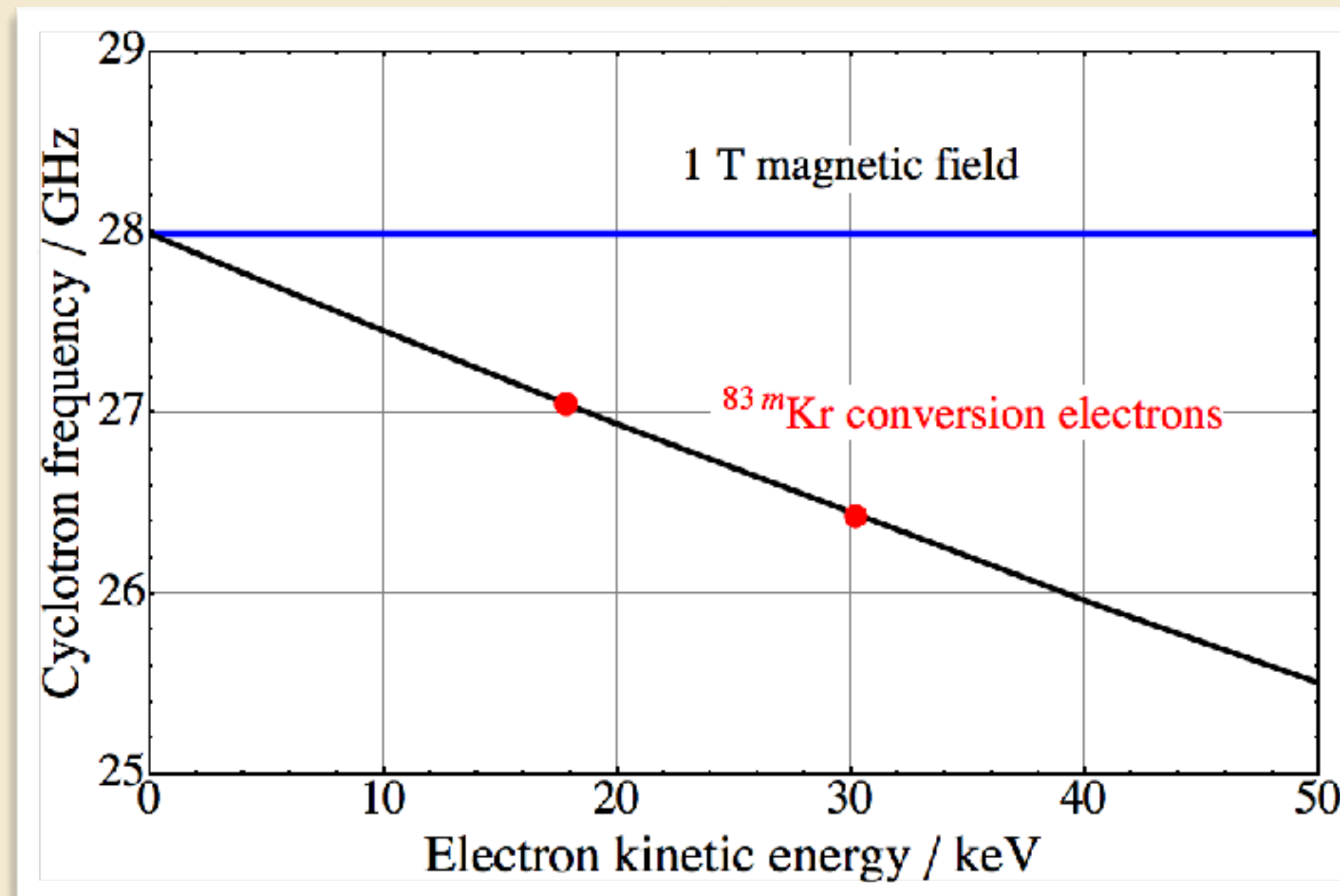
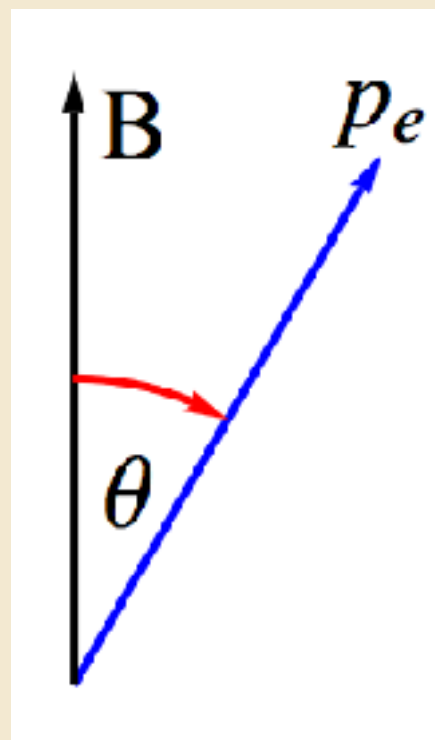
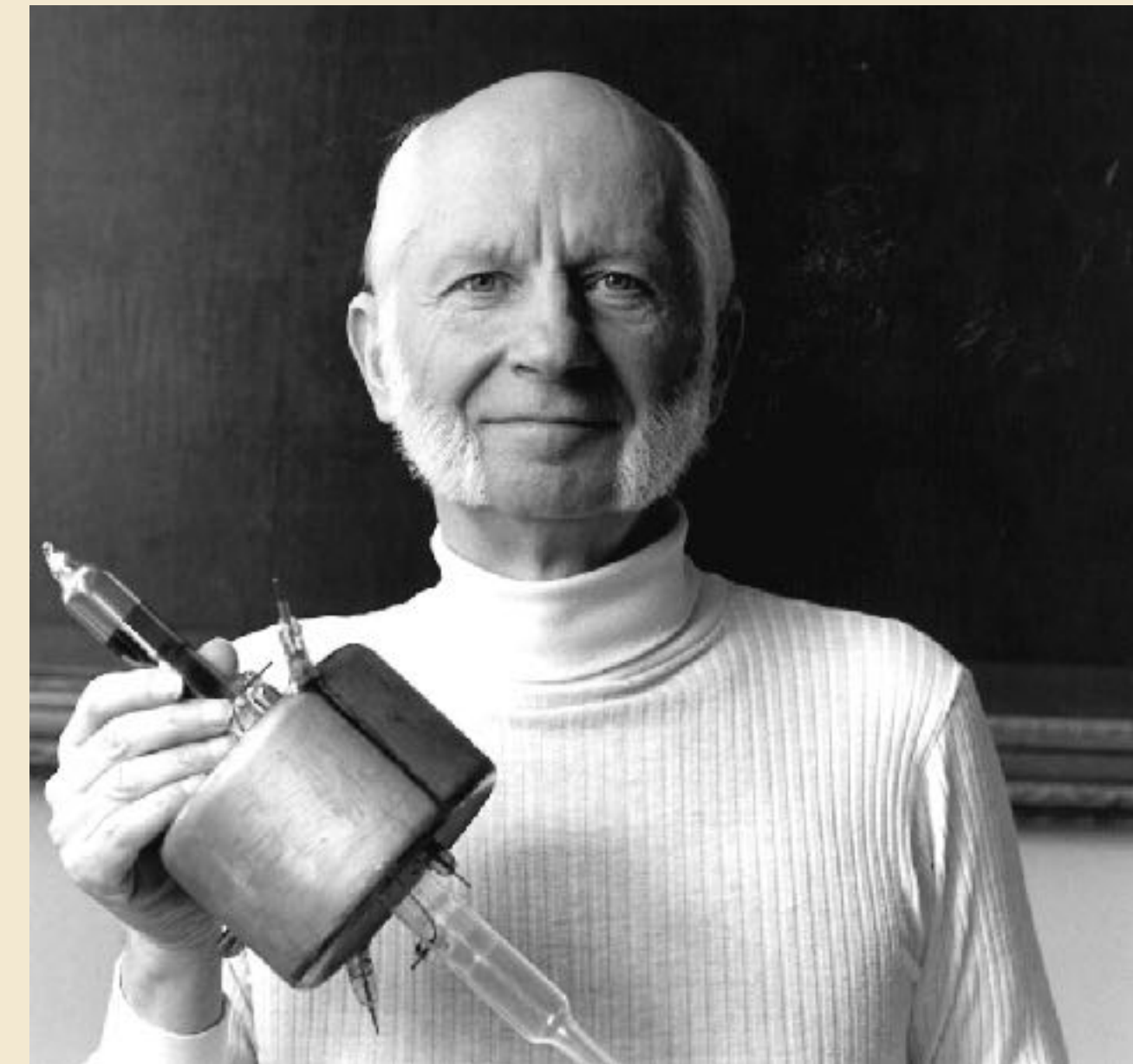
Novel approach: J. Formaggio and B. Monreal, Phys. Rev D 80:051301 (2009)

- Cyclotron radiation from single electrons
- Source transparent to microwave radiation
- No e⁻ transport from source to detector
- Highly precise frequency measurement

$$f_c = \frac{f_{c,0}}{\gamma} = \frac{1}{2\pi} \frac{eB}{m_e + E_{\text{kin}}/c^2}$$

Hans Dehmelt

9.9.1922- 7.3.2017

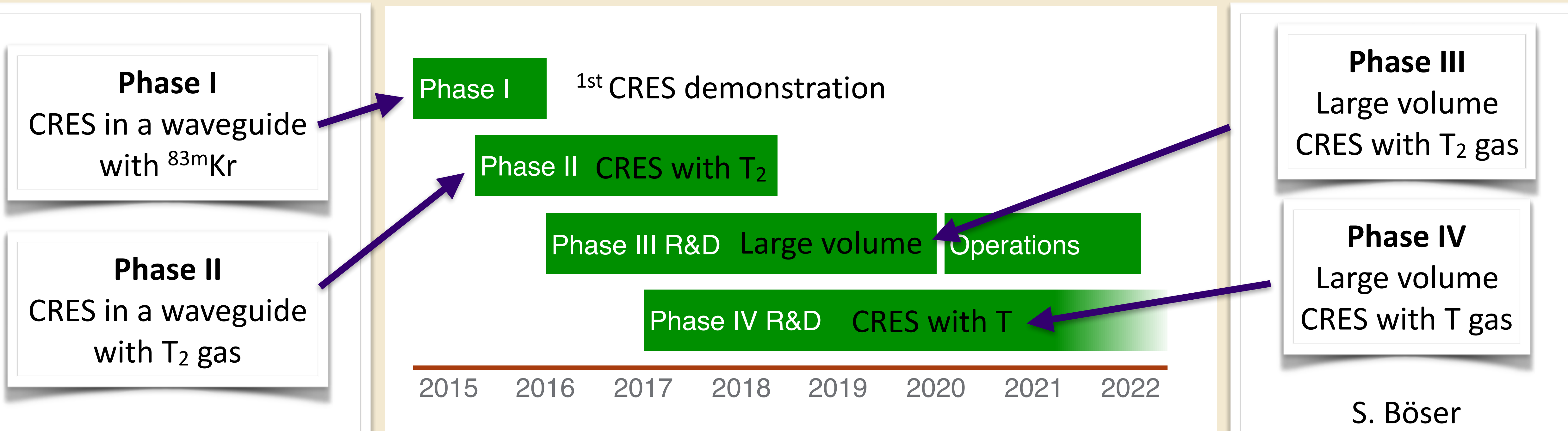


$(m_e c^2) \sin^2 \theta$

... but readily detectable with
... of the art detectors

Project 8: A four phases approach

Every aspect of the new technology has to be demonstrated before we can design an atomic tritium experiment!



Phase I
CRES in a waveguide
with ^{83m}Kr

Phase II
CRES in a waveguide
with T₂ gas

C. Claessens
Signal reconstruction
Thursday, 18:55 - 19:15

Phase III
Large volume
CRES with T₂ gas

Phase IV
Large volume
CRES with T gas

S. Böser
Thursday, 16:35 - 17:00

Frequency and energy resolution of CRES

Energy resolution vs. frequency resolution

$$\frac{\Delta E_{\text{kin}}}{E_{\text{kin}}} = \left(1 + \frac{m_e c^2}{E_{\text{kin}}}\right) \frac{\Delta \nu_c}{\nu_c} \approx 28 \text{ for 18.6 keV electron}$$

$$\Delta E_{\text{kin}} \approx 0.2 \text{ eV} \rightarrow \frac{\Delta \nu_c}{\nu_c} \approx 4 \times 10^{-7}$$

$$\nu_c \approx 27 \text{ GHz} \rightarrow \Delta \nu_c \approx 11 \text{ kHz}$$

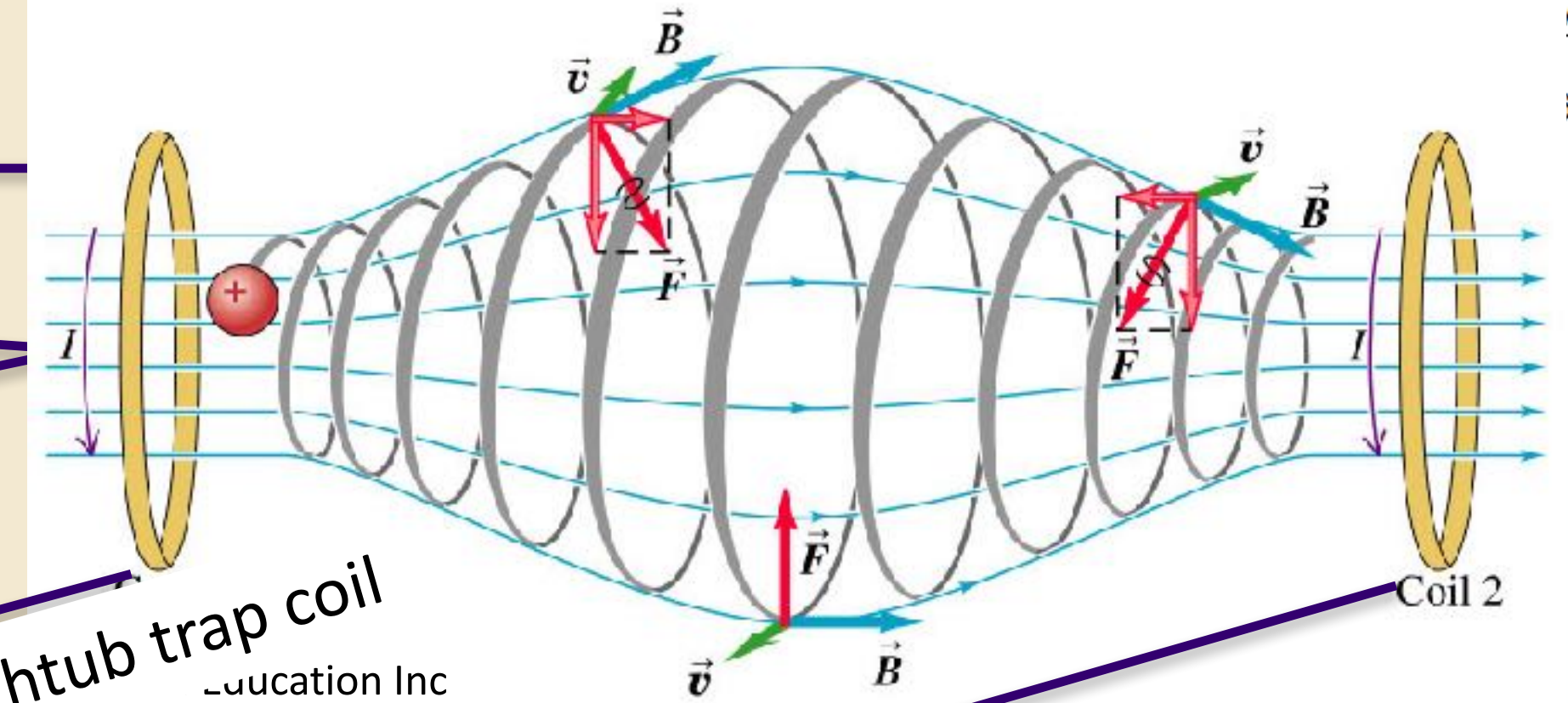
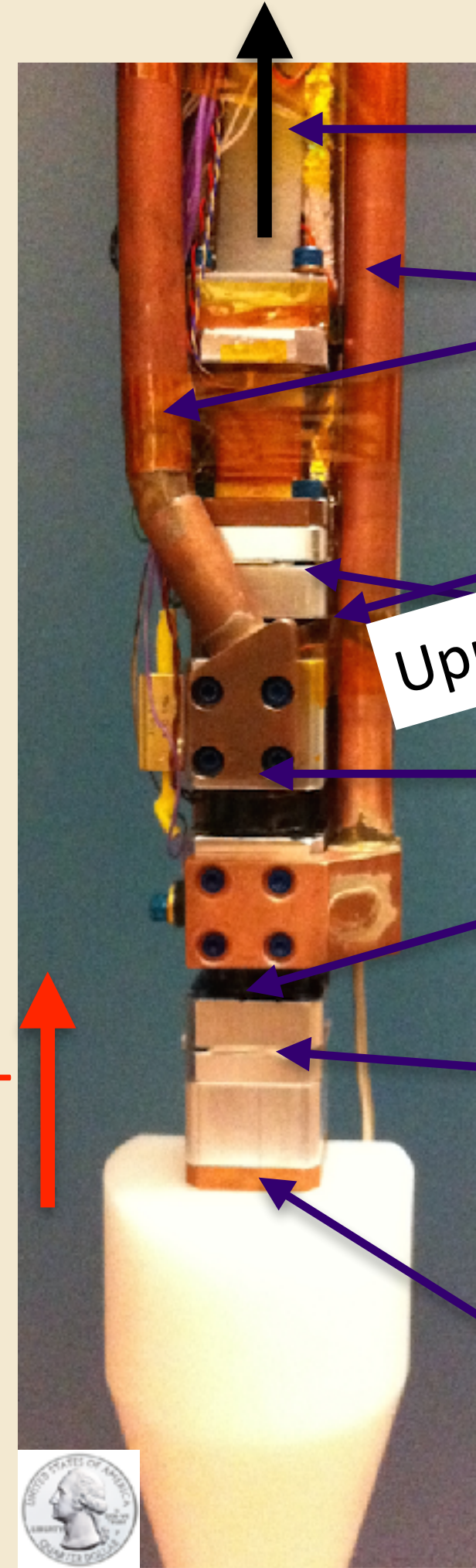
Frequency resolution vs. observation time

$$\Delta \nu_c \times t_{\text{obs}} \geq \frac{1}{2\pi} \rightarrow t_{\text{obs}} \geq 14 \mu\text{s}$$

→ Need for a magnetic (no work) trap!

To cryogenic amplifier

Compare with MAC-E filter in KATRIN!



Upper bathtub trap coil

Indium gasket

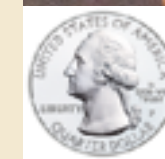
Gas volume

Lower bathtub trap coil

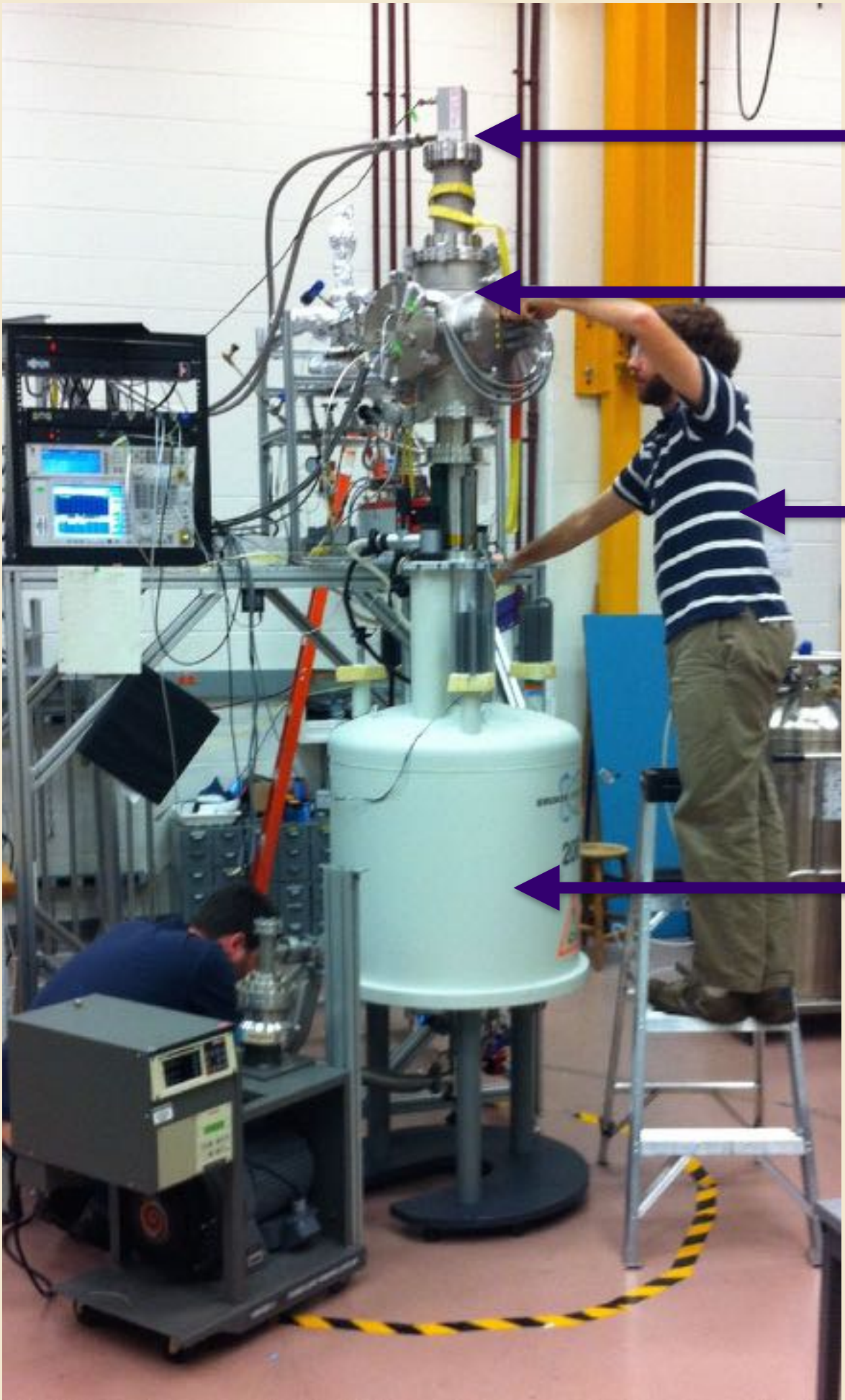
Lower Kapton window
Indium gasket

WR42 waveguide short

1T ↑



Phase I: A table top demonstrator experiment



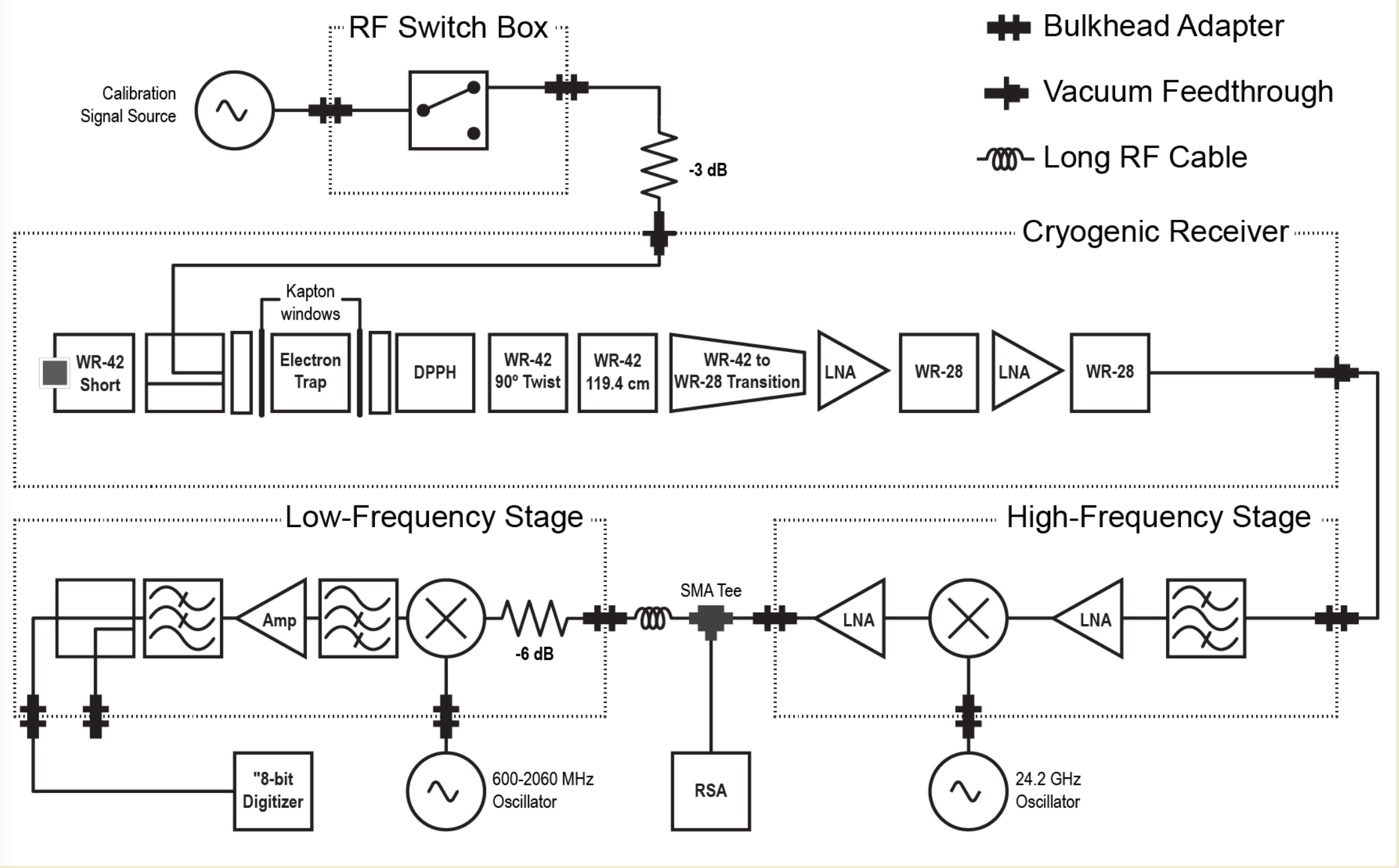
30 K cold head

Low noise cryogenic amplifiers

Ben LaRoque

NMR magnet

Double stage superheterodyne receiver chain:
25.2 GHz → 1 GHz



The first light in the detector from a single electron

Digitize output of a super-heterodyne mixing stage



Fourier transform short time slices

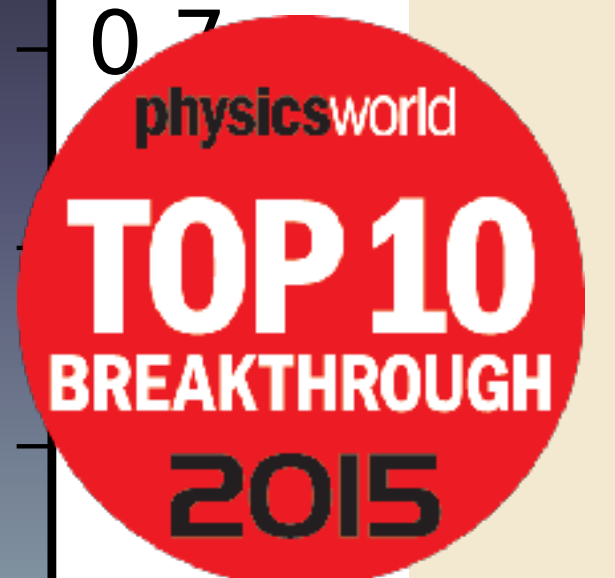
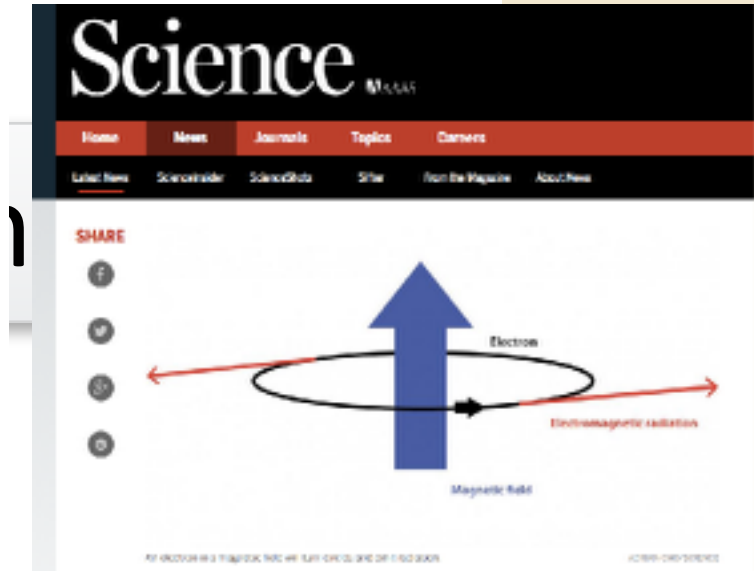
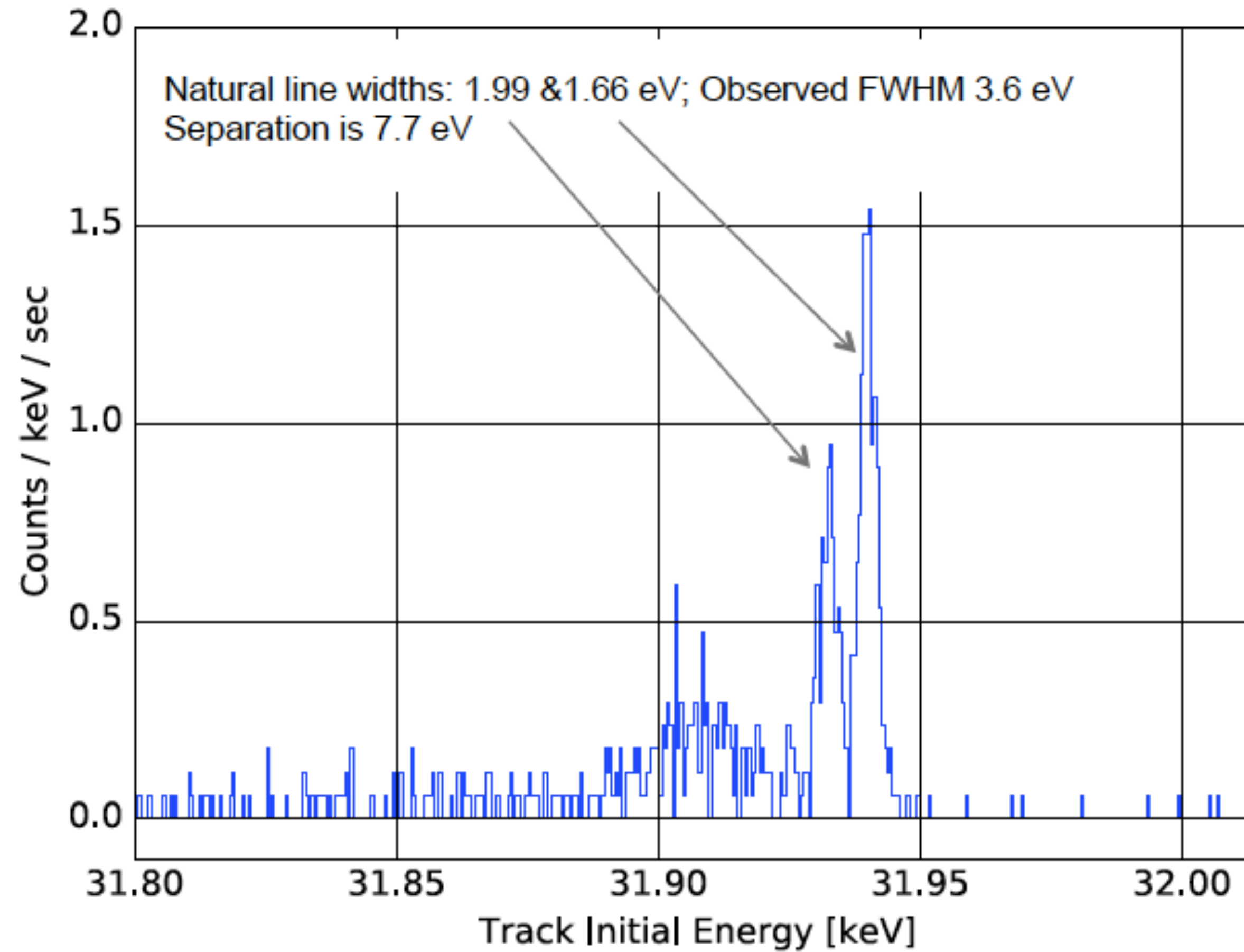


Plot power vs. frequency as function of time

Asht

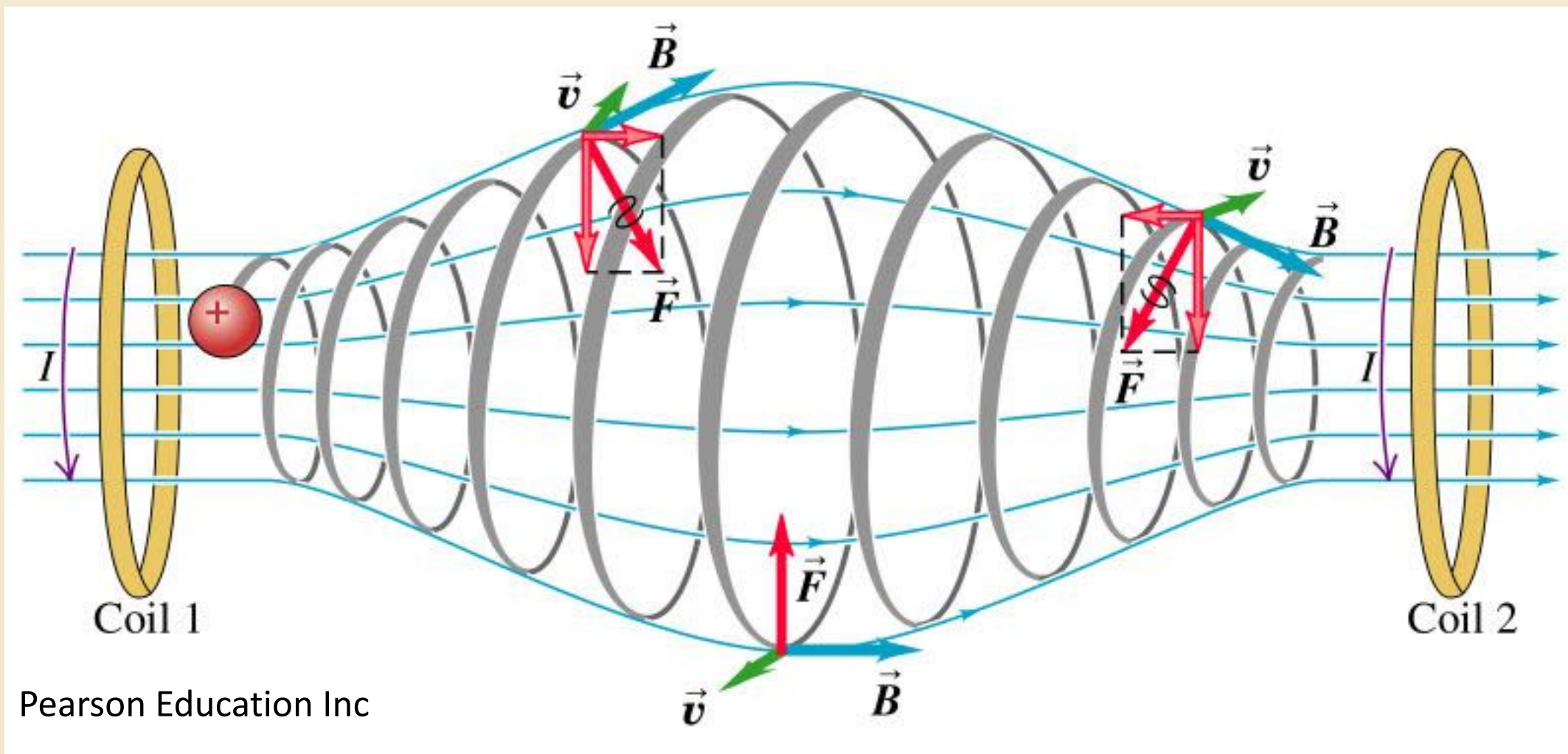
M. Fertl

Region of interest near the 32 keV lines
(bins are 0.5 eV wide)

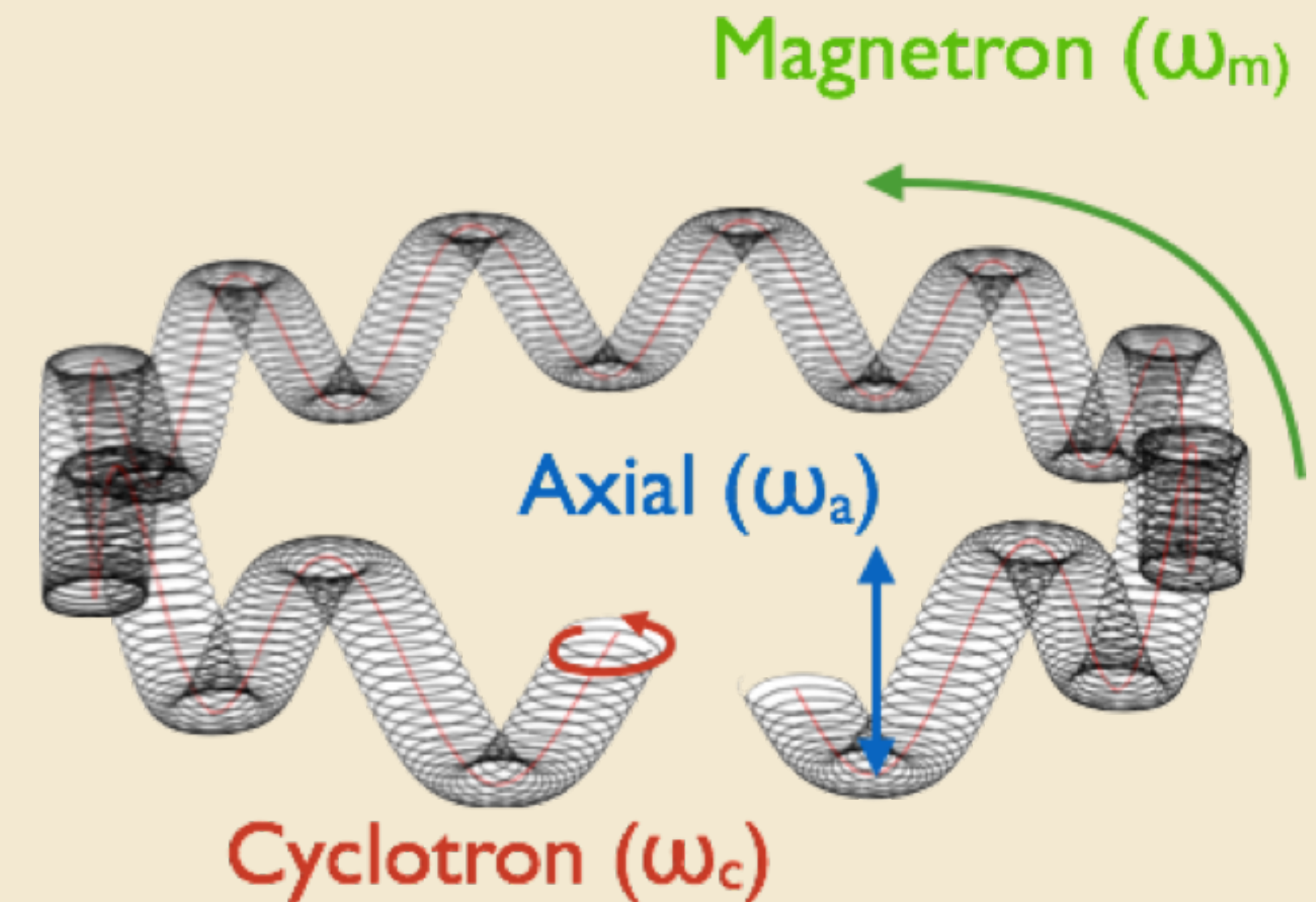


Rich frequency spectrum of raw CRES signal

e^- confinement in magnetic bottle:
→ Non-trivial e^- motion



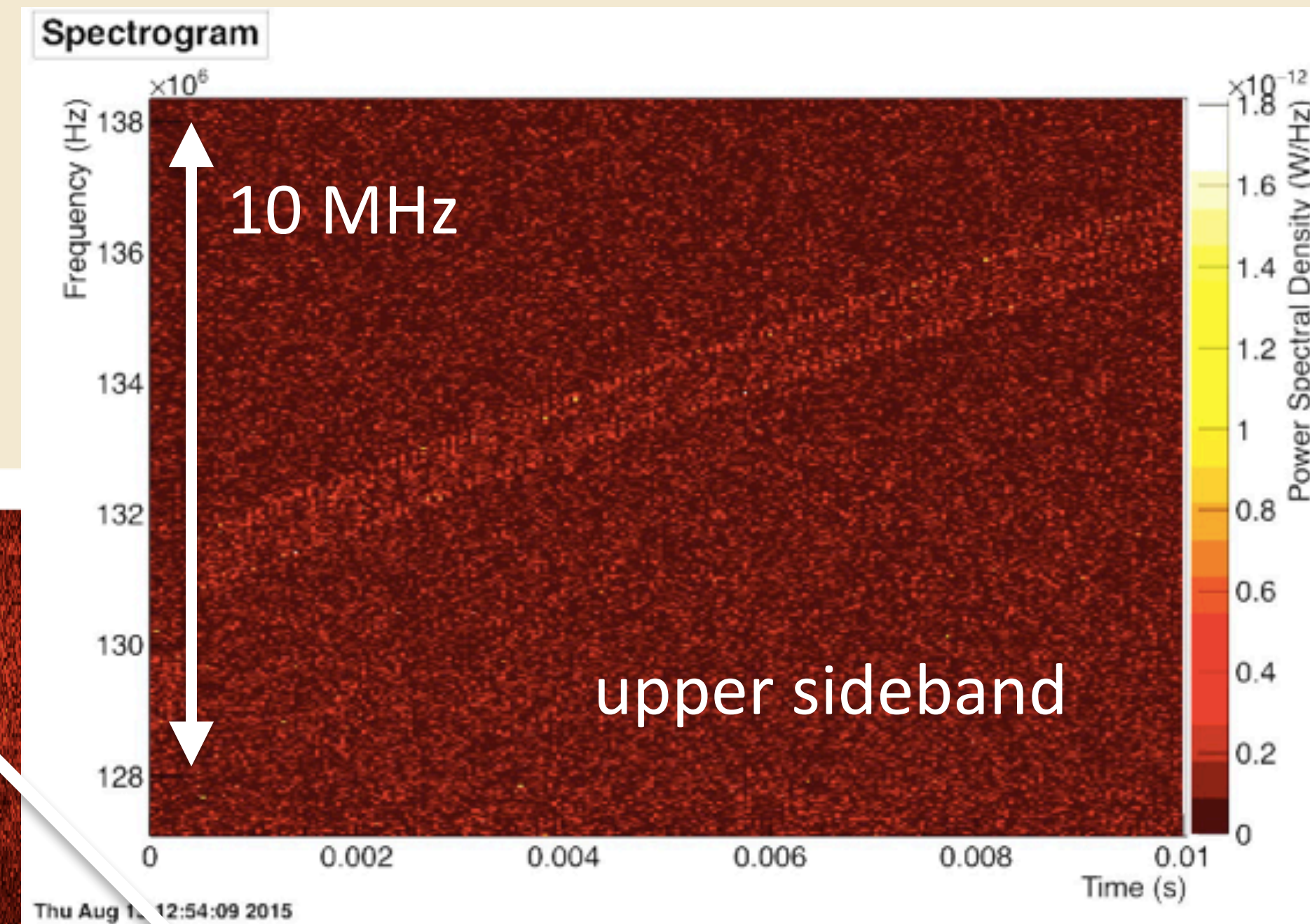
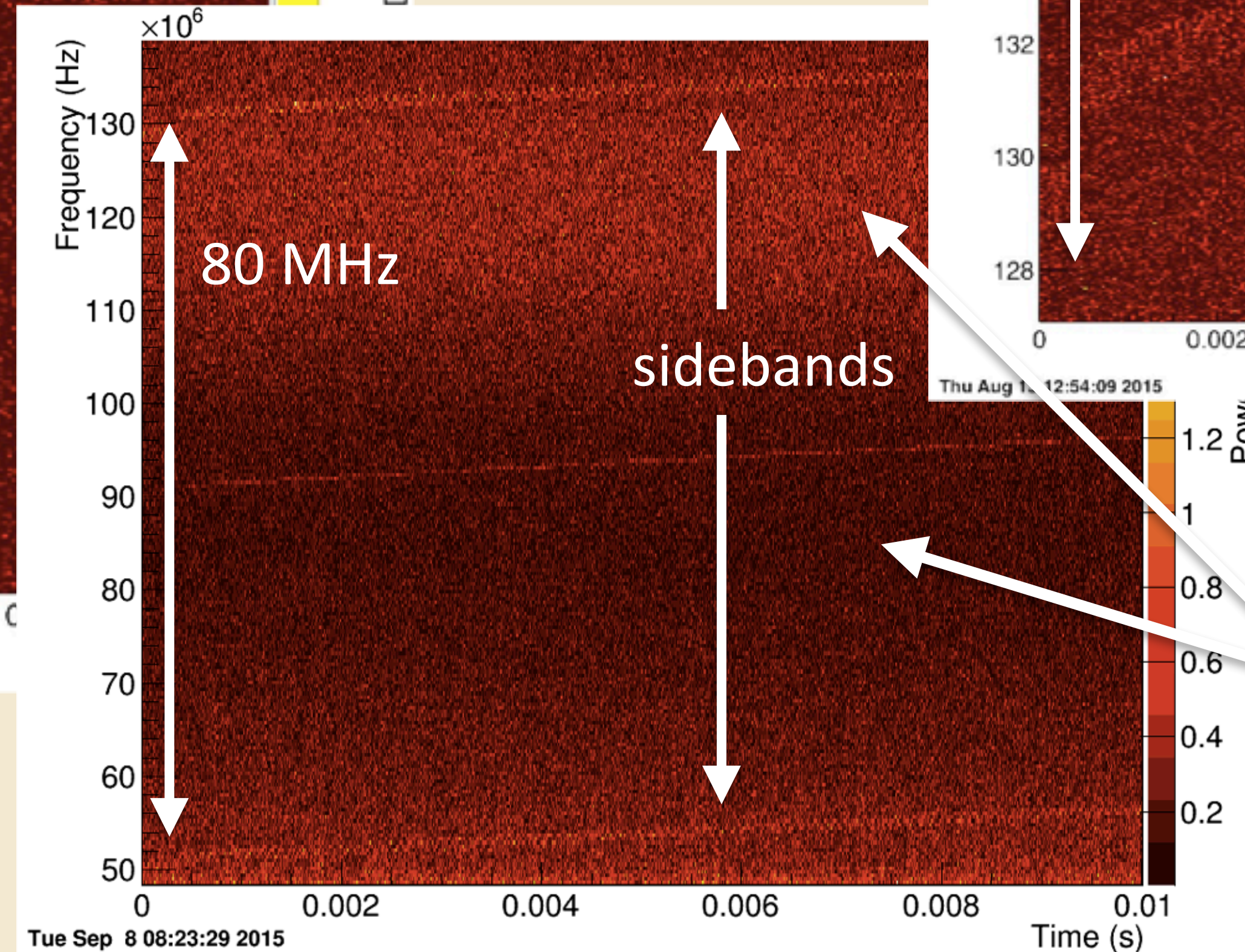
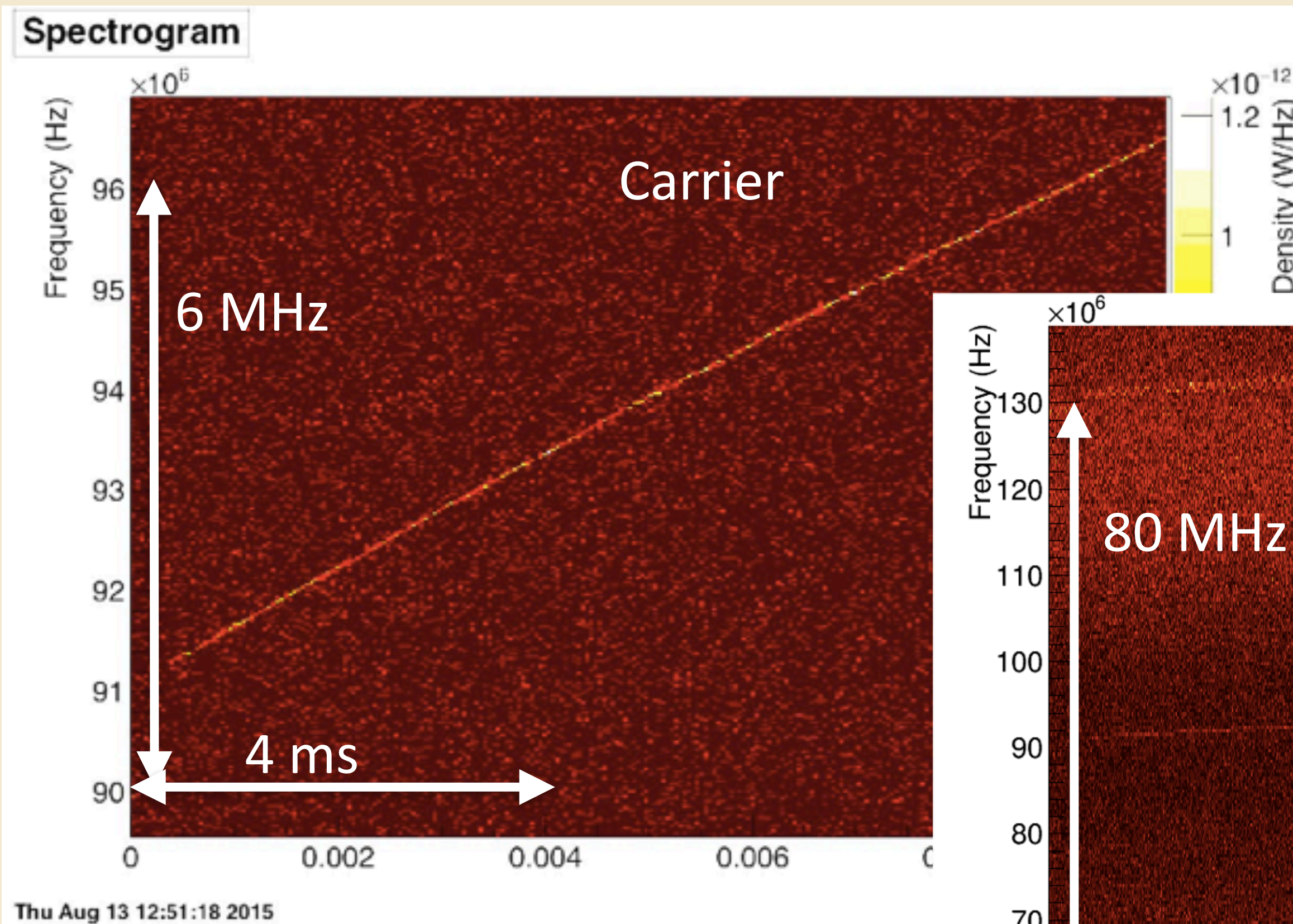
Similarity to Penning trap with
three separated eigenmotions



http://wswww.physik.uni-mainz.de/werth/g_fak/penning.htm

Signal model predicts frequency side bands due to B-modulation and Doppler-effect!

Rich frequency spectrum of CRES signals

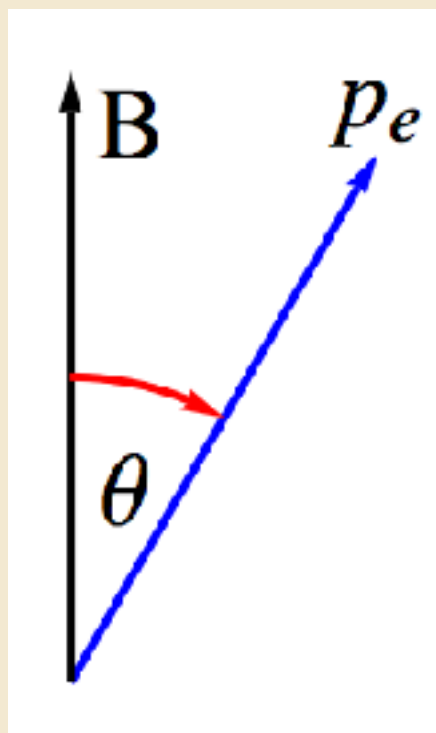


Phase I:
Non-uniform microwave
background!

Pitch angle correction

In a harmonic magnetic trap ($B \sim z^2$) the measured microwave frequency depends on the electron pitch angle:

magnetic field at trap bottom



pitch angle $\cong 90^\circ$

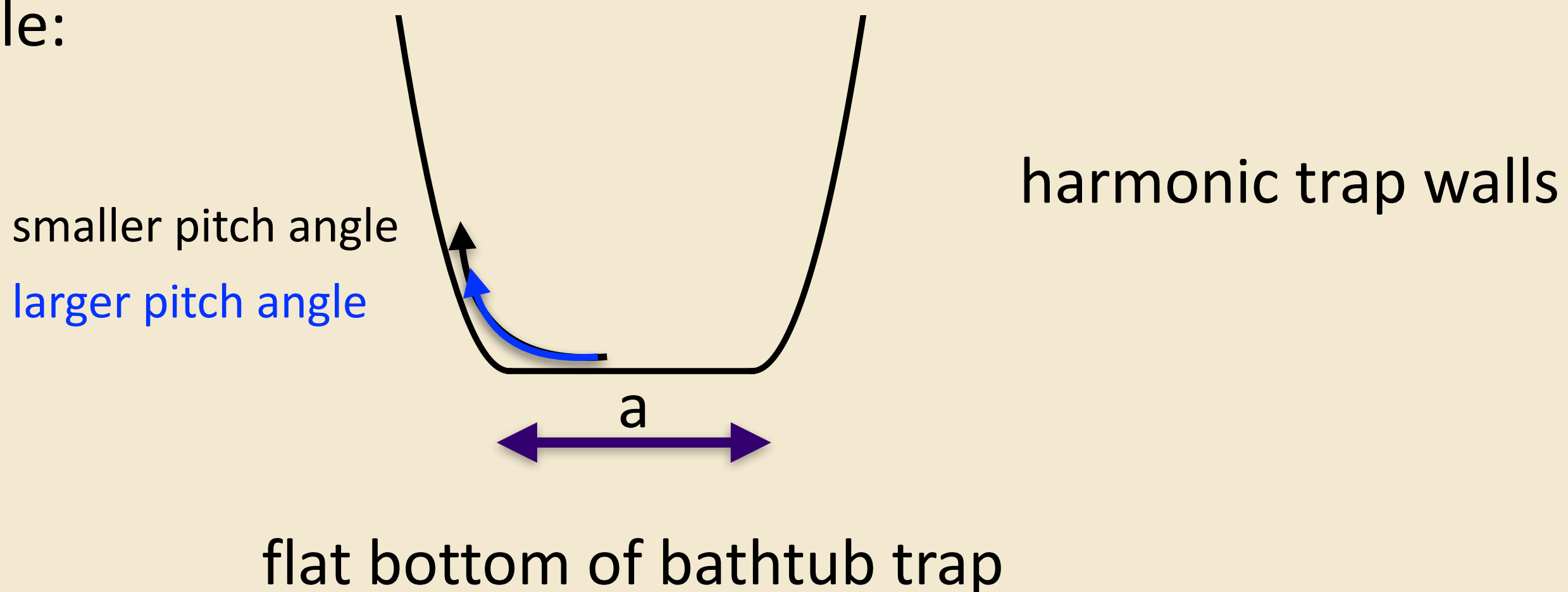
$$\omega_\gamma = \frac{eB}{m + E_{\text{kin}}} \left(1 + \frac{\cot^2 \theta}{2} \right)$$

measured microwave frequency

kinetic energy

More details: see talk C. Claessens

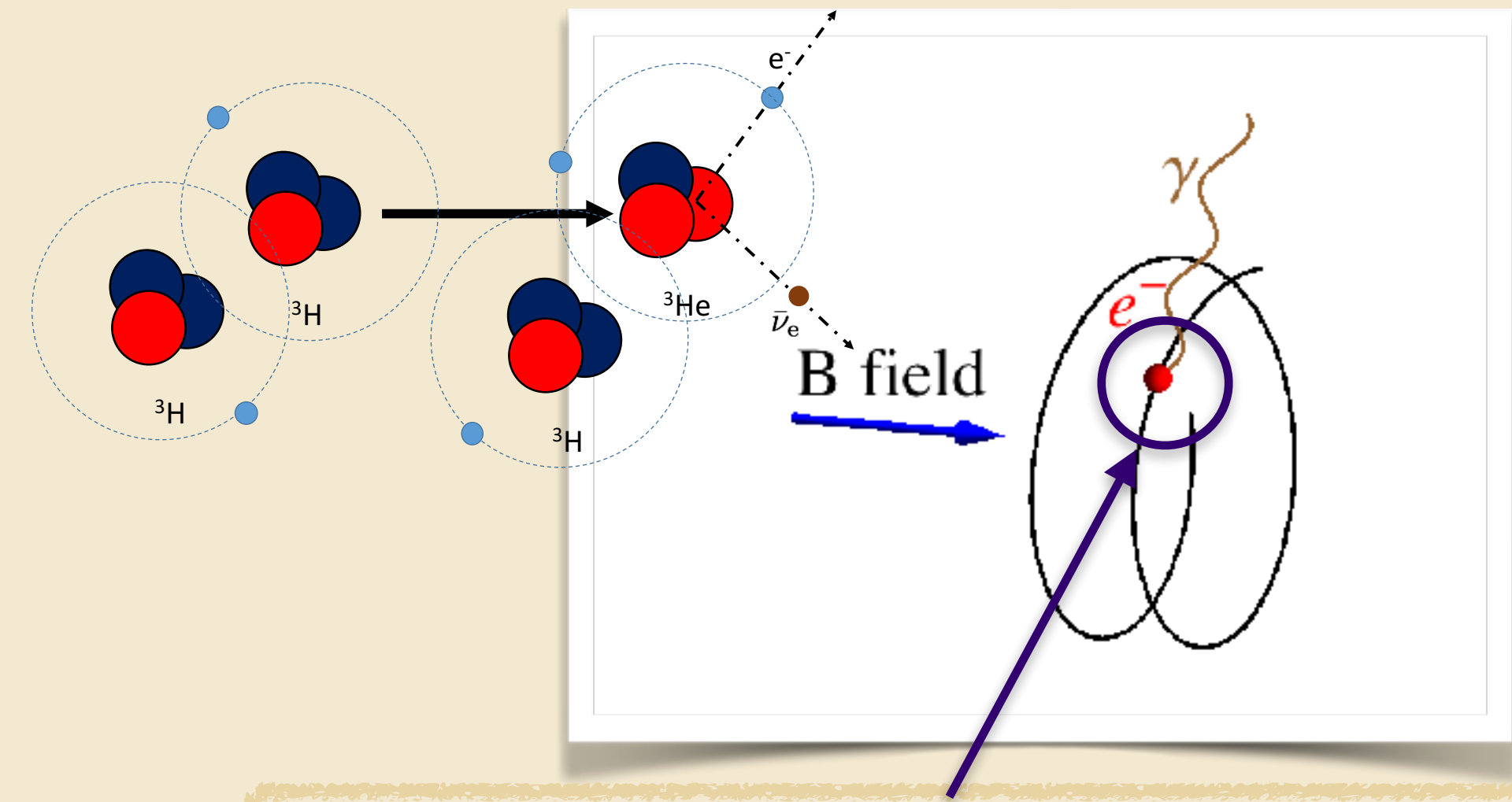
Approximate model of the bathtub trap



Approximate expression for axial frequency

$$\omega_a \propto v \left(\frac{a}{\sin \theta} + \frac{4 \sin \theta}{m \cos^2 \theta} \right)^{-1}$$

Project 8 Phase II: A combined T₂ and ^{83m}Kr source



$$f_c = \frac{f_{c,0}}{\gamma} = \frac{1}{2\pi m_e + \frac{eB}{c^2} + E_{\text{kin}}/c^2}$$

Source of T₂ decay electrons:

- Safe handling of T₂ gas
- Temperature range: ≅ 80-300 K
- Fine T₂ pressure regulation
- 10⁻⁹-10⁻⁵ torr operation pressure
- Gas composition measurements

Source of ^{83m}Kr conversion electrons:

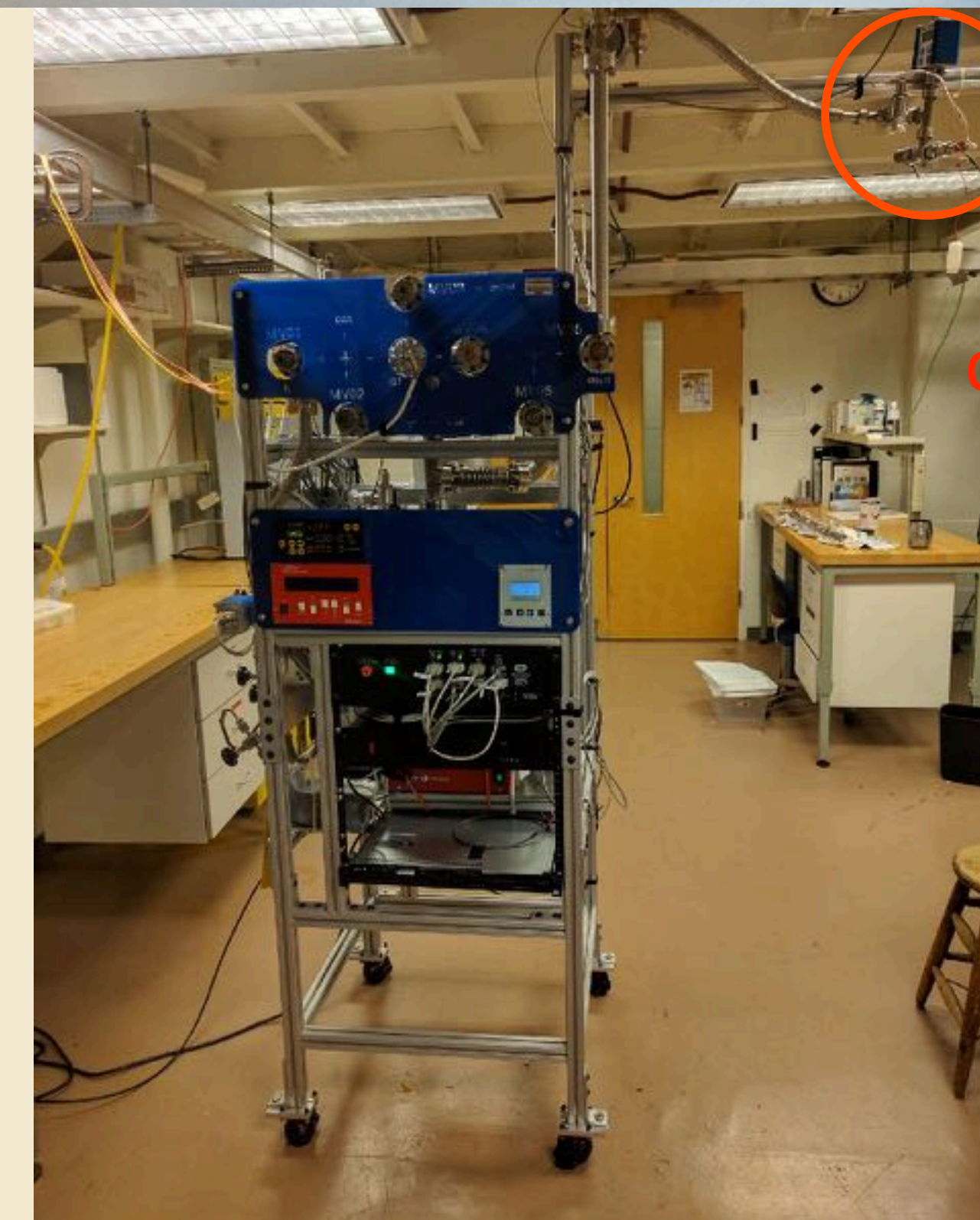
- ^{83m}Kr CE energies well known
- Matching CE energy range 9 - 32 keV
- B field long term stability
- ^{83m}Kr goes where T₂ goes (co-magnetometer)

Tritium storage in metal getter

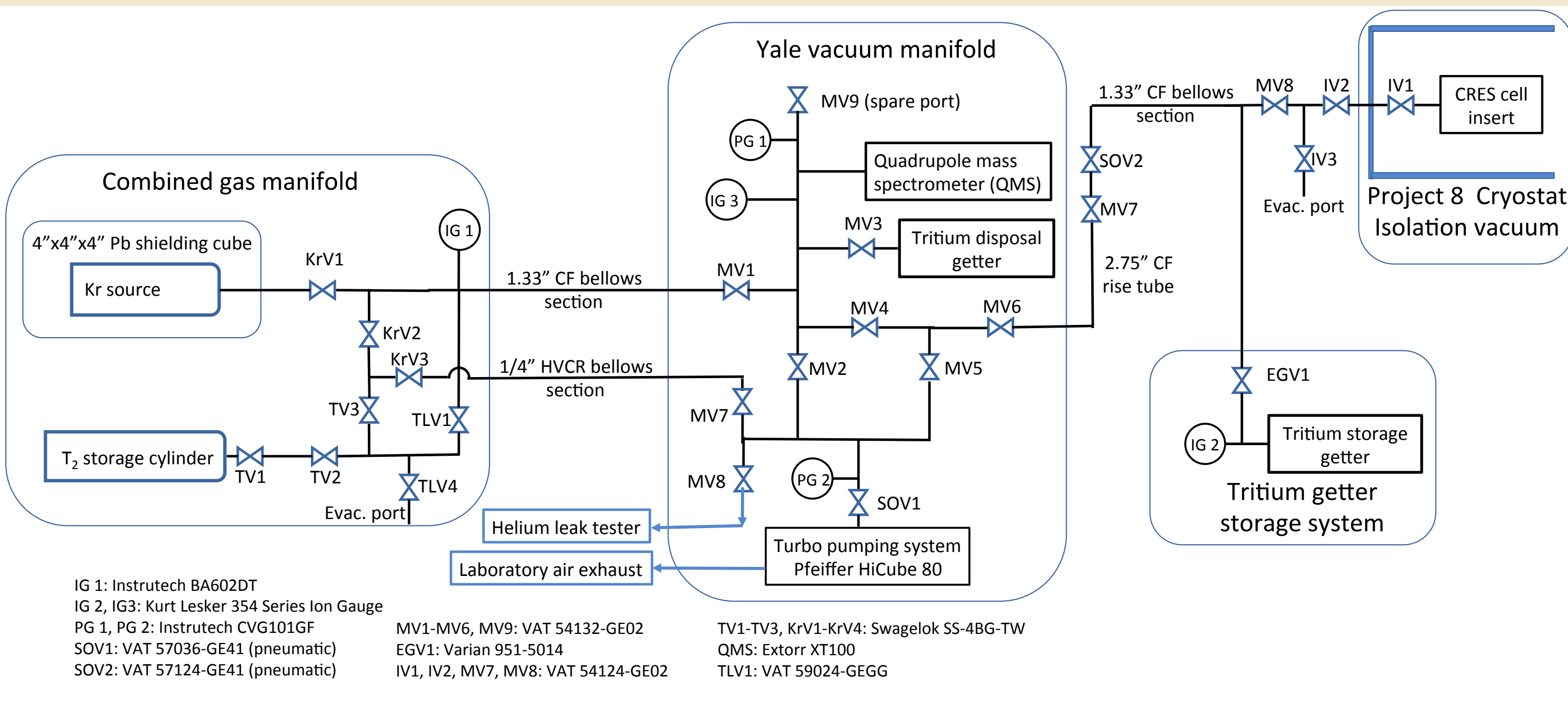
Successfully tested the loading and release of D₂ from a SAES St172 getter sample (together with KIT):

- Continuous pumping of H₂, CO, CO₂, H₂O, CH₄
- Simple pressure regulation
- Expect similar behavior for T₂
- Recapture T₂ at room temperature

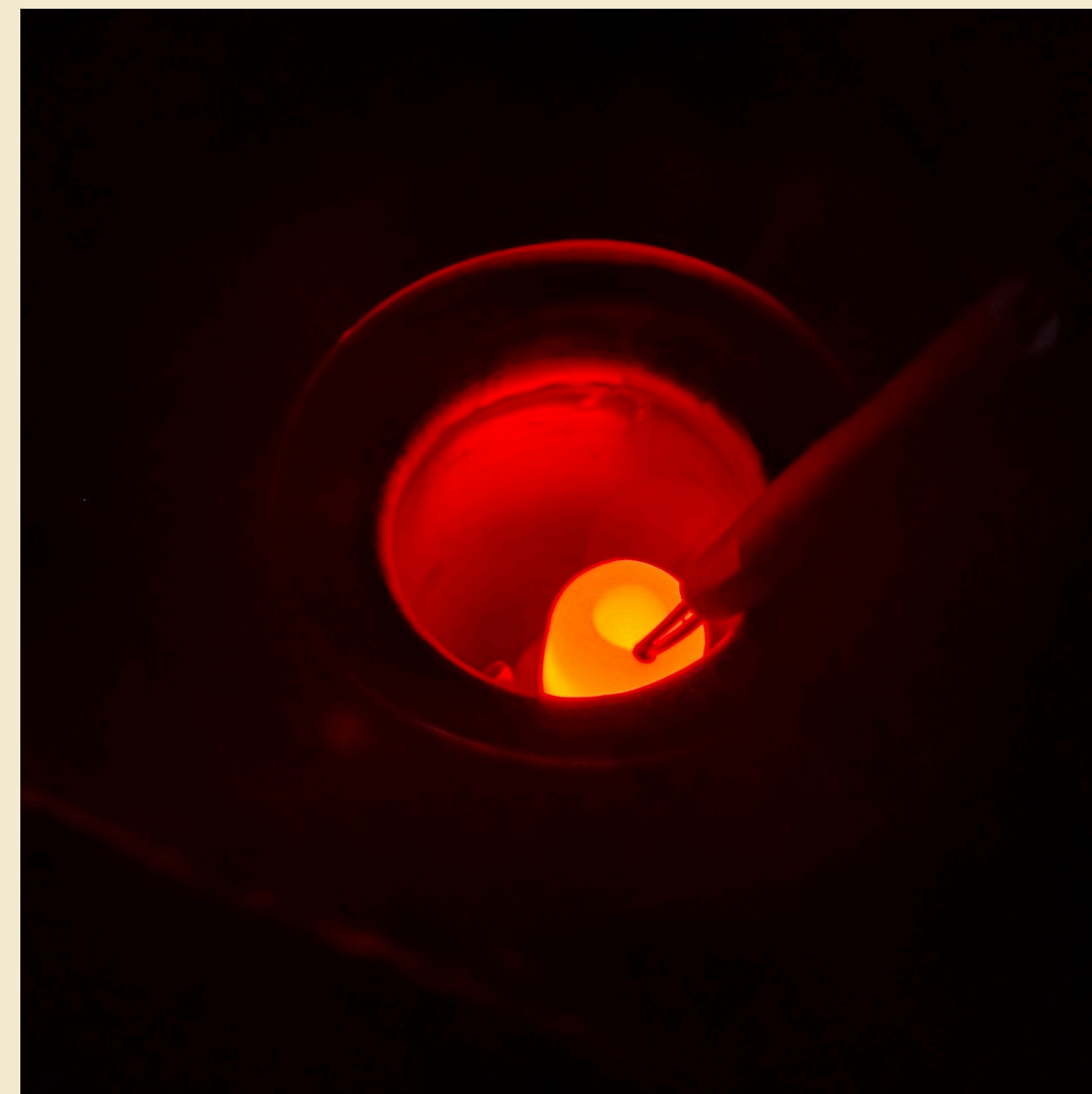
Compact T₂ reservoir built and commissioned @ UW



Getter compartment

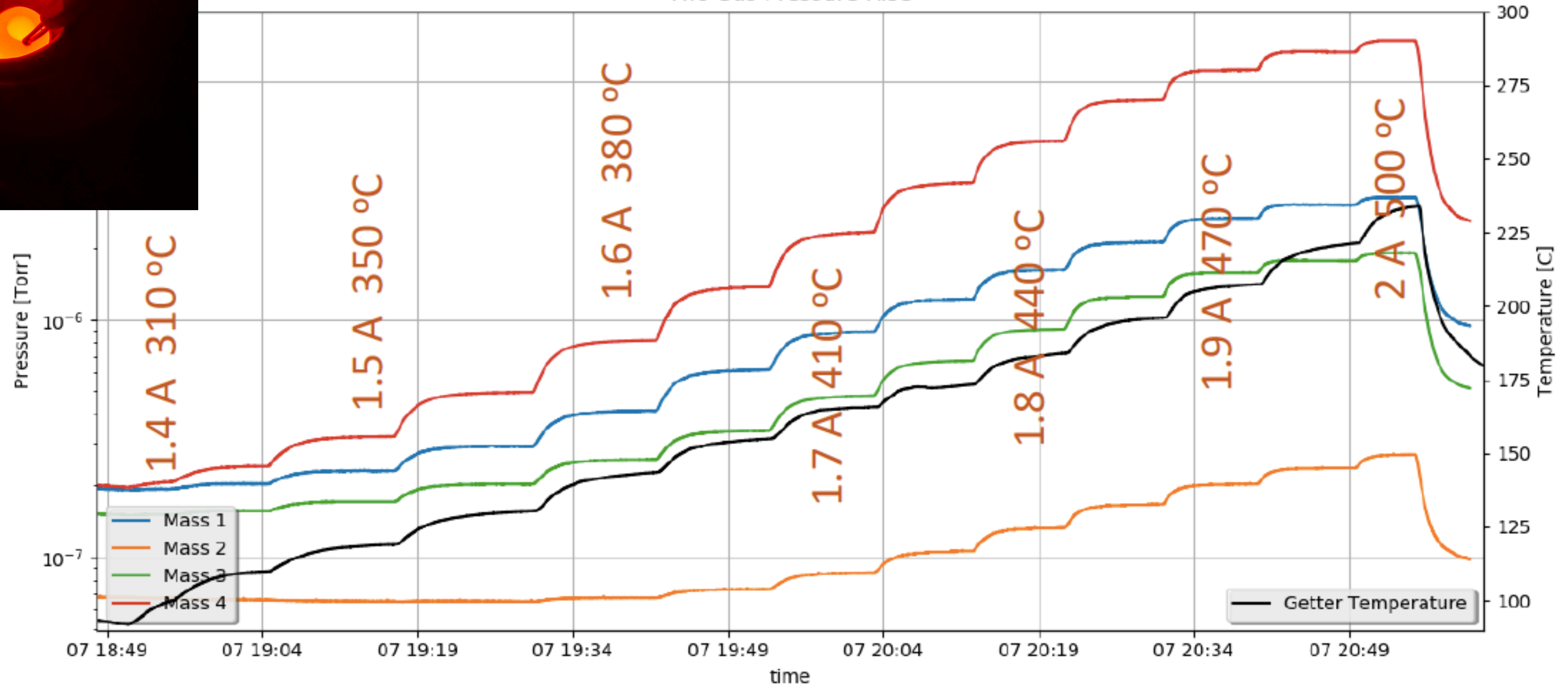


Tritium storage in metal getter



The Gas Pressure Rise

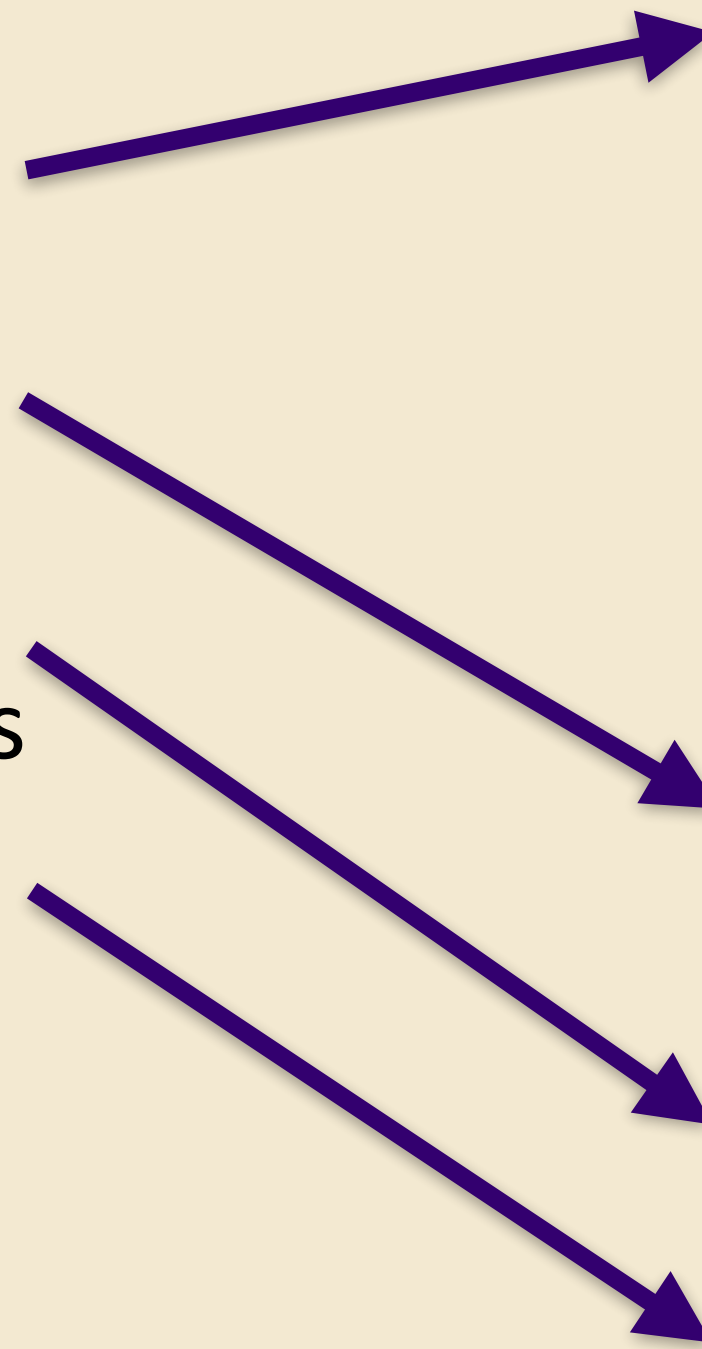
preliminary, A. Ashtari (UW)



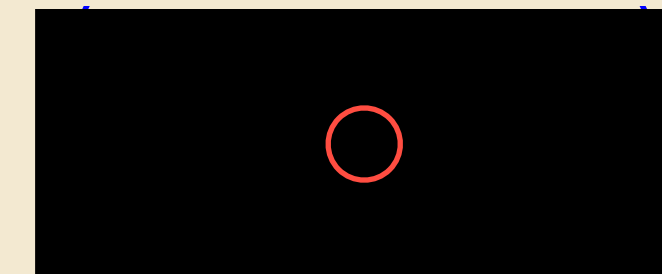
Project 8 Phase II: T₂ compatible CRES cell

Technical challenges for a T₂ compatible CRES cell

1. Transport circularly polarized radiation.
2. Larger volume!
3. Low-loss microwave guide!
4. No ferromagnetic construction material.
5. Hermetic confinement of T₂ gas
6. Low loss microwave transparent windows
7. Cryogenic operation
8. Matched thermal expansion coefficients



Phase I: WR42



0.170" (4.3 mm)

0.420" (10.7 mm)

Oxygen-free high-conductivity copper

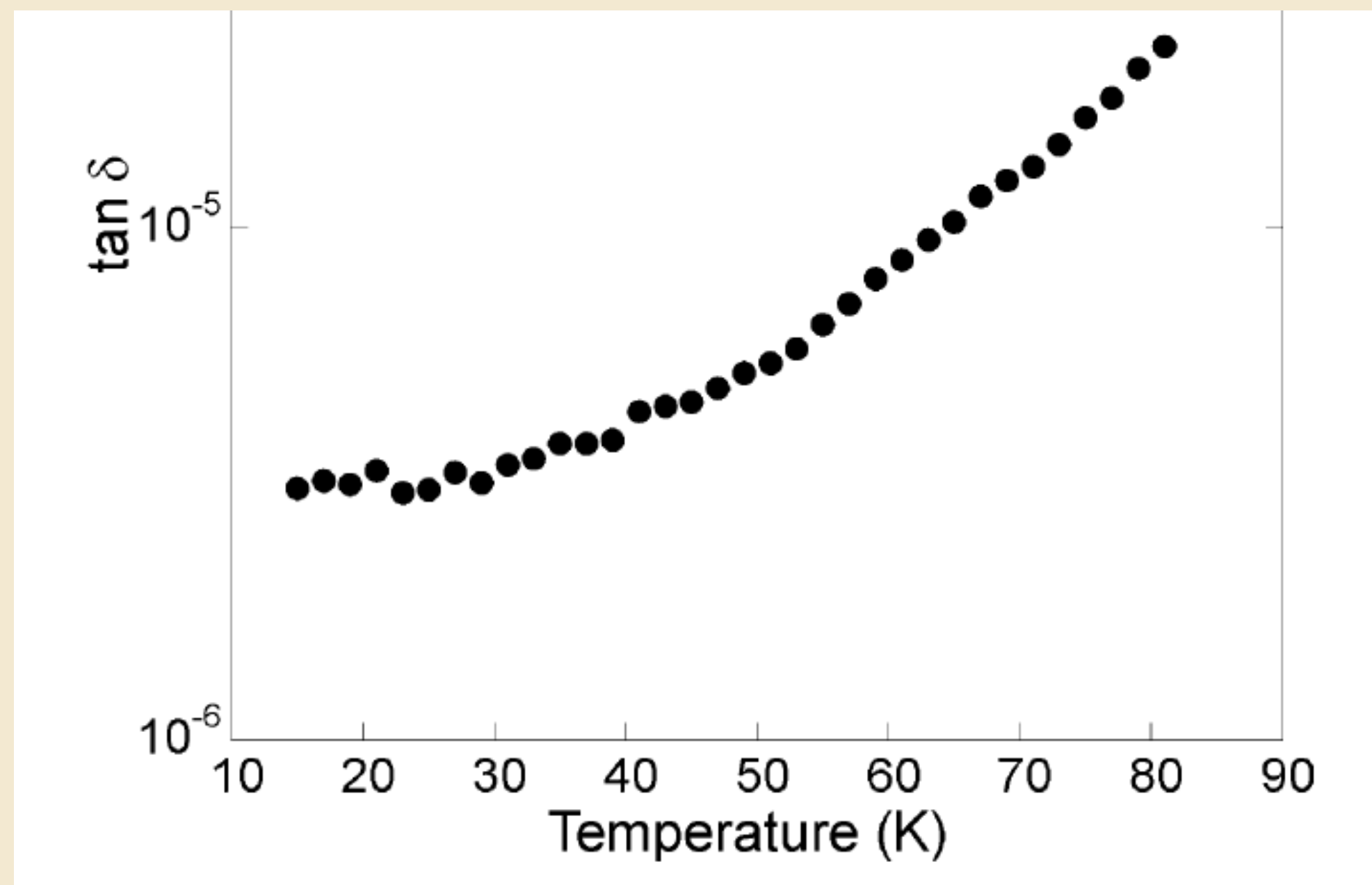
No organic (kapton) window materials!

Calcium difluoride windows, indium sealed

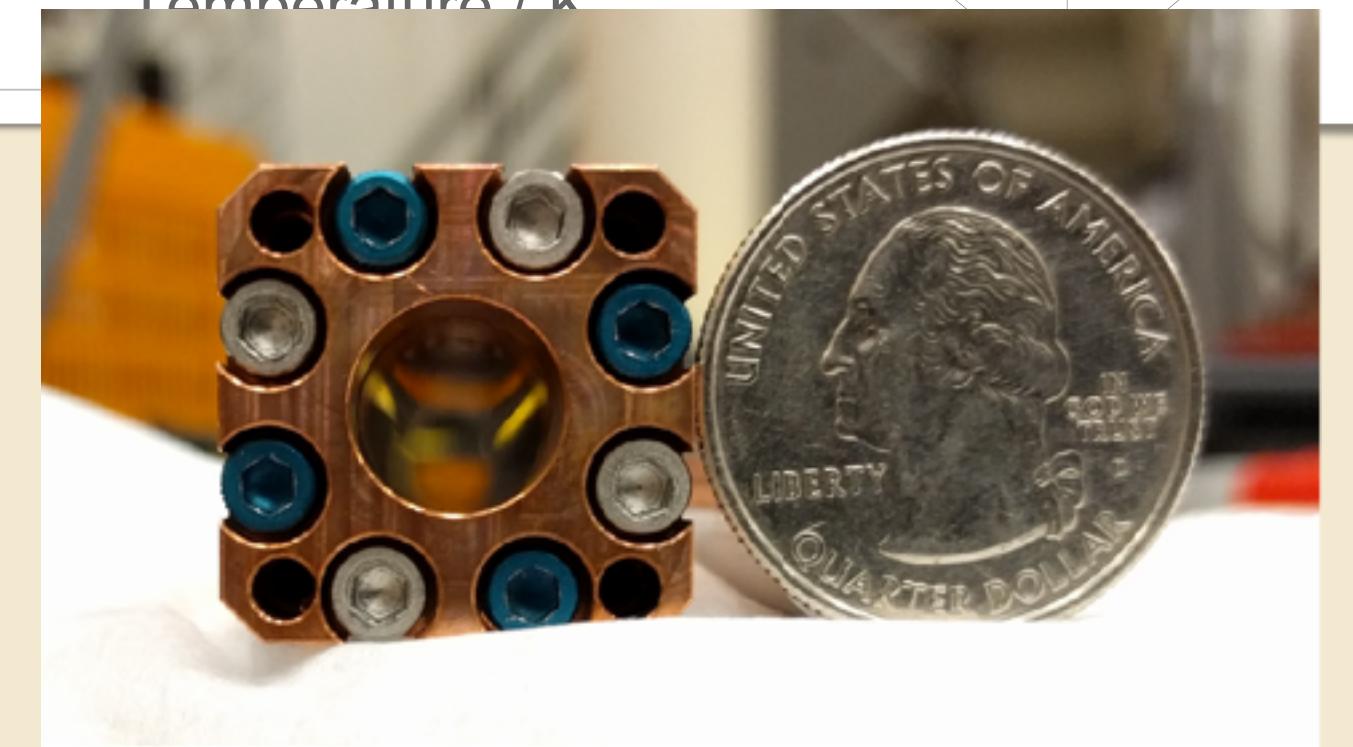
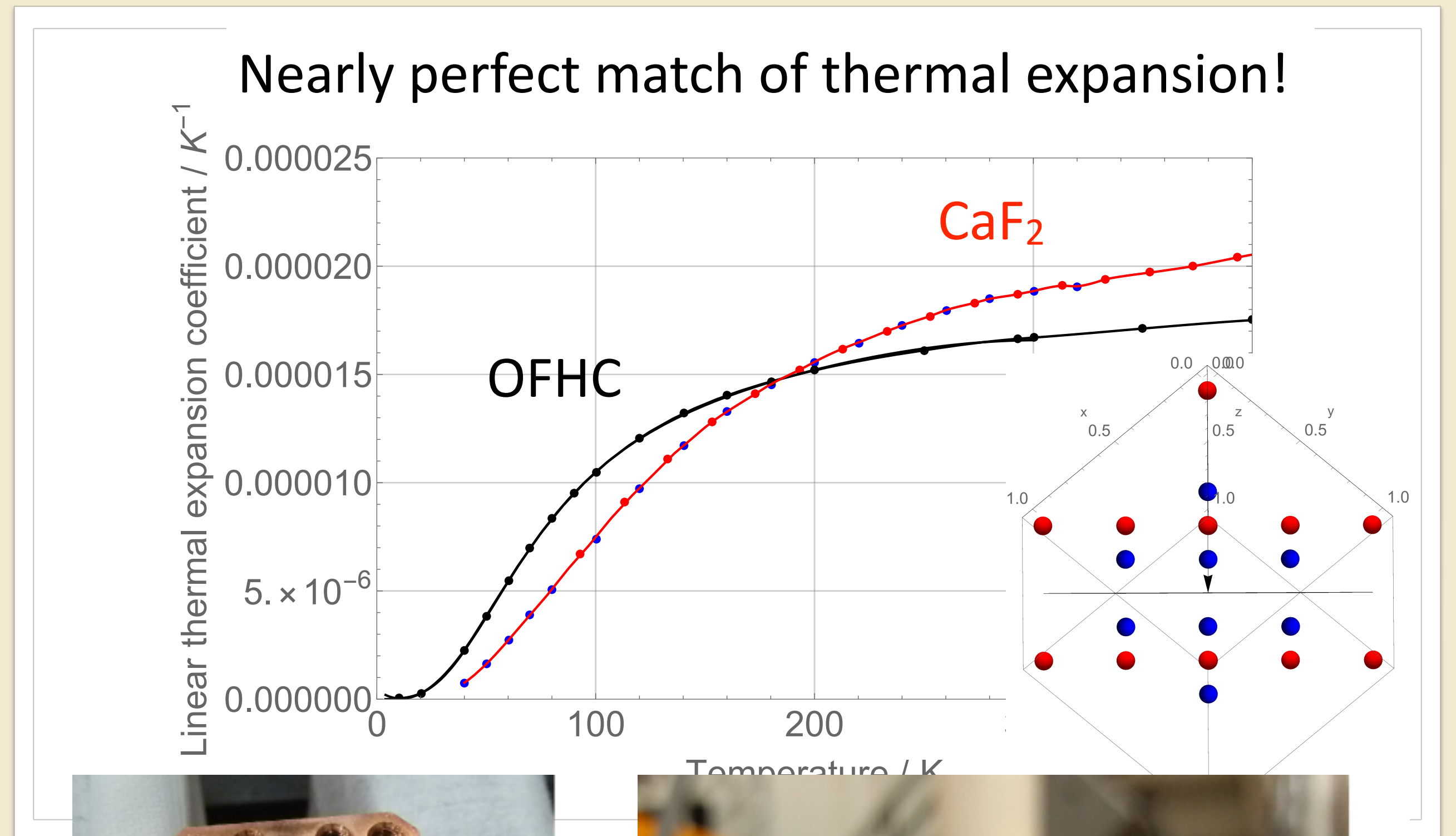
CaF₂ windows are ideal but very brittle!

CaF₂ window material:

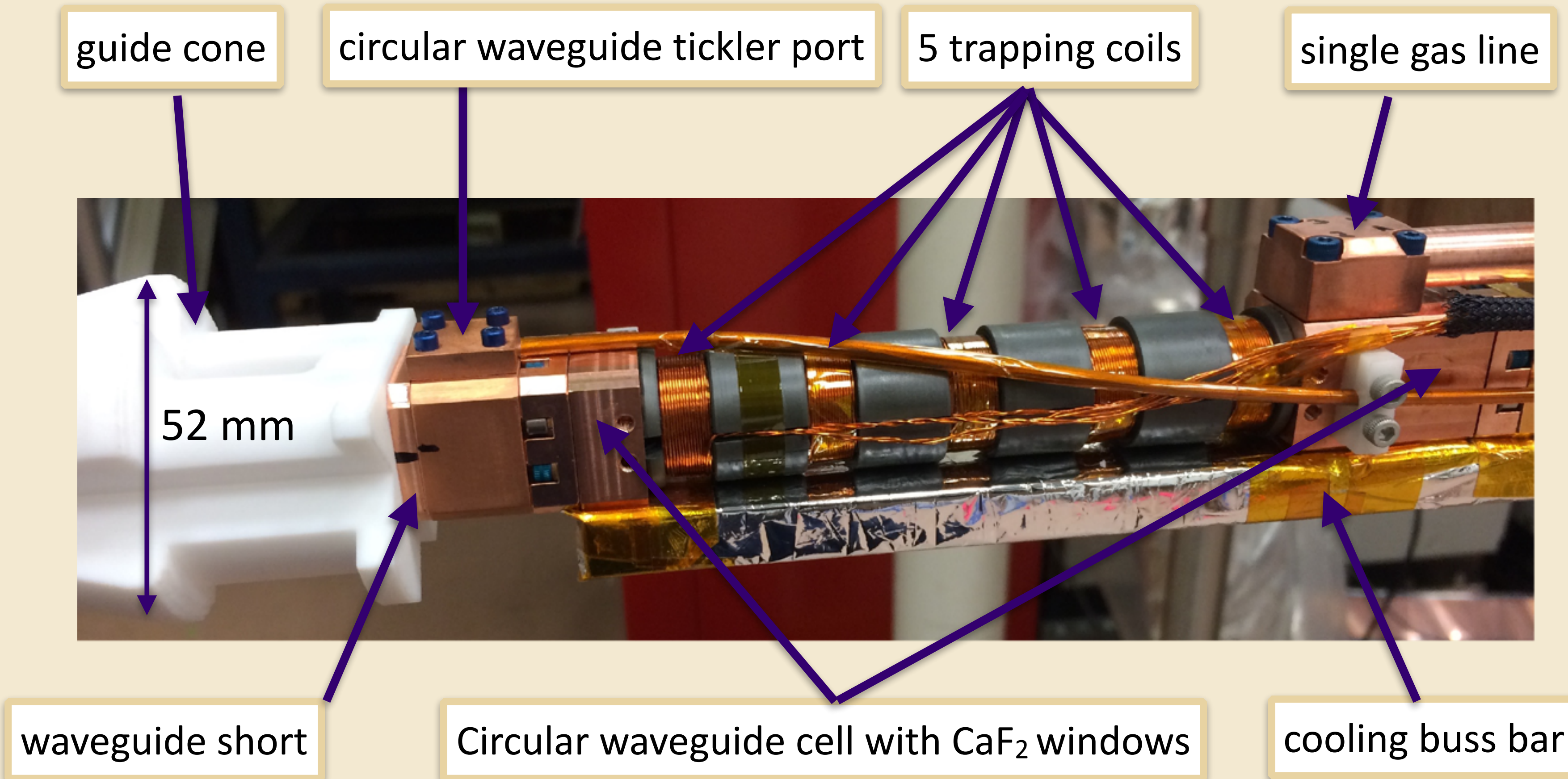
- + Low loss tangent (low absorption loss)
- + Isotropic crystal properties
- + Small dielectric constant (low reflections)
- Very soft and brittle



M. Jacob et al., J. of the Europ. Ceramic Society 23 (2003); 2617



The Phase IIa CRES cell before installation



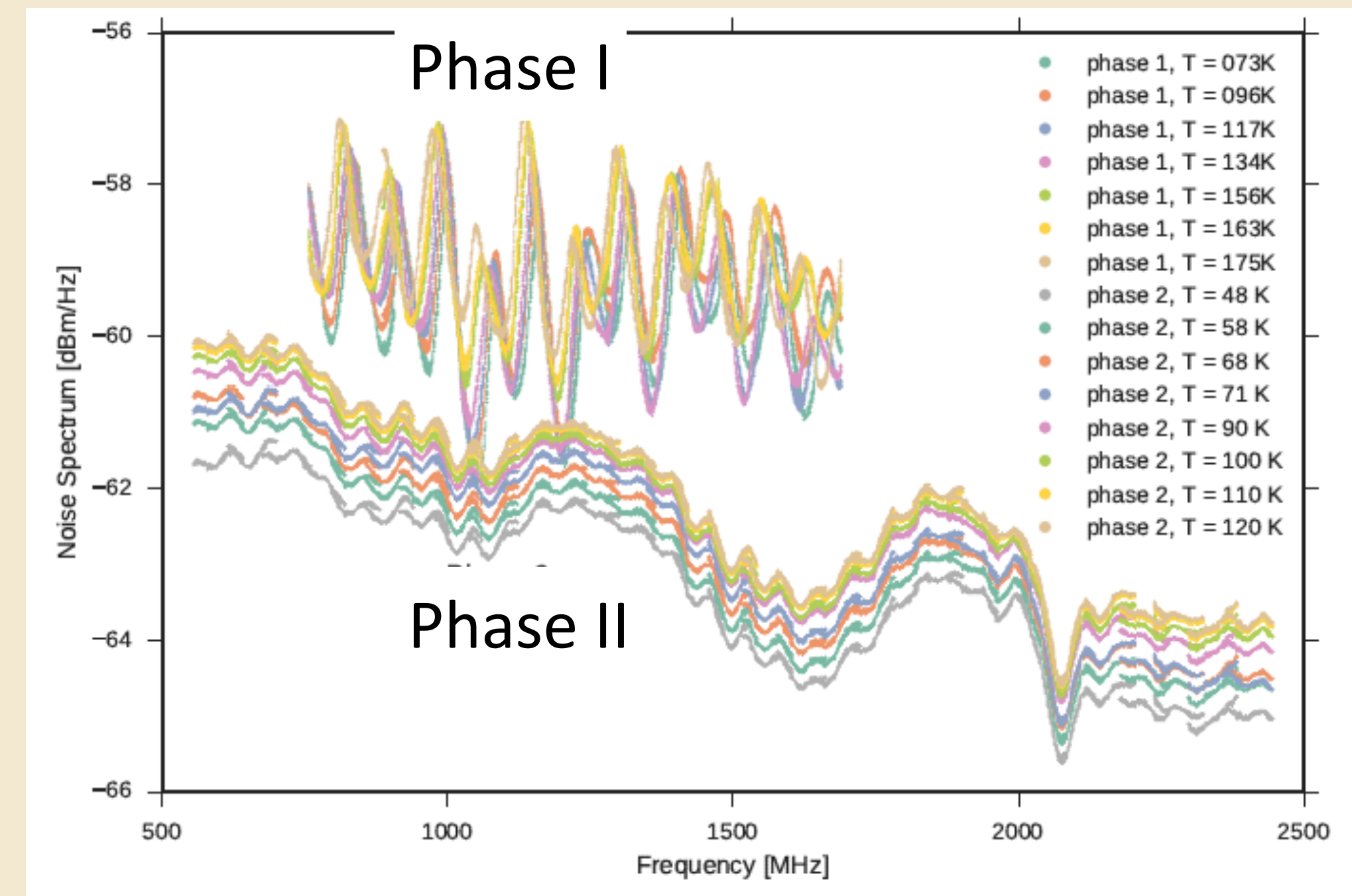
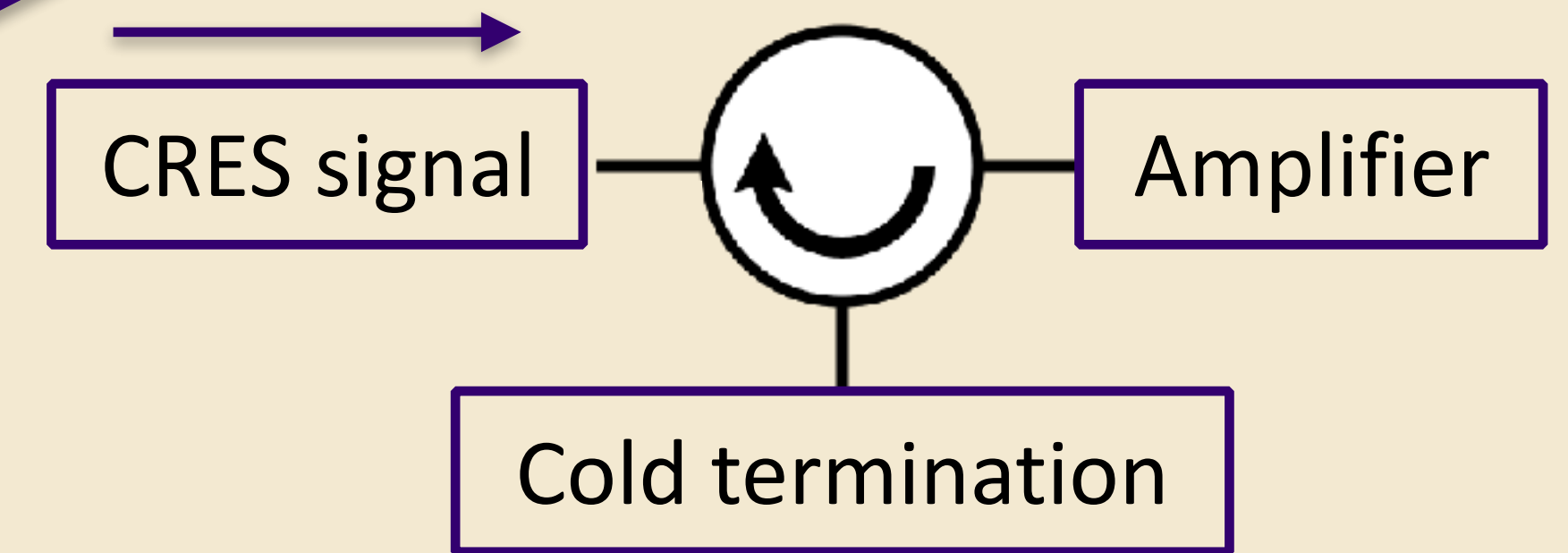
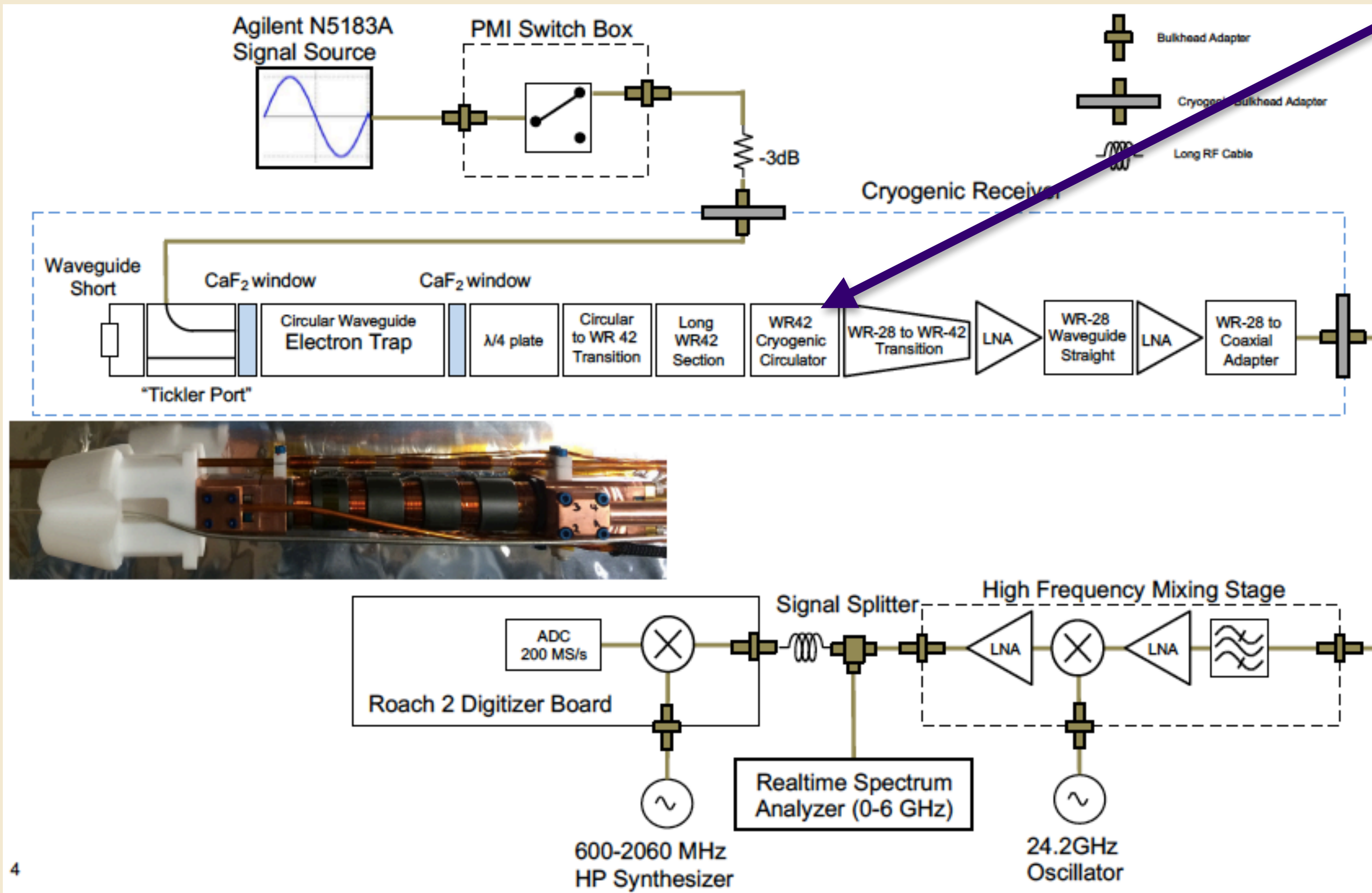
Successful CRES cell operation:

- Commissioning with ^{83m}Kr
- 50-300 K cell temperature

To be used with T₂ gas after late fall 2017!

Upgraded microwave infrastructure

Cryogenic circulator



Typical CRES signals live

17.8 keV electrons, realtime x20

lower energy ⇨

Video removed for web version

ongoing commissioning with $^{83\text{m}}\text{Kr}$

DAQ Phase II upgrade and the path towards Phase III

Multichannel, scalable DAQ system update for Phase II based on ROACH:

- **Reconfigurable Open Architecture Computing Hardware**
- Widely used in radioastronomy
- Based on Xilinx Virtex 5 FPGA
- Project 8 is currently commissioning a 3 channel version

see C. Claessens
(Thursday)



Increase T_2 gas volume for Phase III:
larger magnetic field volume,
free space antenna array



see S. Boeser's talk (Thursday)

Summary

Project 8:

- Phase I: 1st direct observation of cyclotron radiation from a single electron
- Phase I: Successfully measured $^{83\text{m}}\text{Kr}$ spectrum using CRES
- Phase II: Operation of a T₂ compatible CRES cell initially with $^{83\text{m}}\text{Kr}$
Combined T₂ and $^{83\text{m}}\text{Kr}$ gas system built and operation approved
First T₂ CRES measurement later this fall
- Phase III+IV: R&D program concurrent with Phase II

Thank you!

Acknowledgments

Benjamin Monreal, Yu-Hao Sun, Laura Gladstone

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Johannes Gutenberg Universität, Mainz

Thomas Thümmel, Marcel Walter

Karlsruhe Institute of Technology

Kareem Kazkaz

Lawrence Livermore National Laboratory

Nicholas Buzinsky, Joseph Formaggio, Joseph Johnston, Valerian Sibille, Evan Zayas

Massachusetts Institute of Technology

Erin Finn, Mathieu Guigue, Mark Jones, Benjamin LaRoque, Noah Oblath,

Jonathan Tedeschi, Brent VanDevender

Pacific Northwest National Laboratory

Luiz de Viveiros, Timothy Wendler

Pennsylvania State University

Shep Doleman, Jonathan Weintroub, Andre Young

Smithsonian Astrophysical Observatory

Ali Ashtari Esfahani, Raphael Cervantes, Peter Doe, Martin Fertl, Eric Machado,

Walter Pettus, Hamish Robertson, Leslie Rosenberg, Gray Rybka

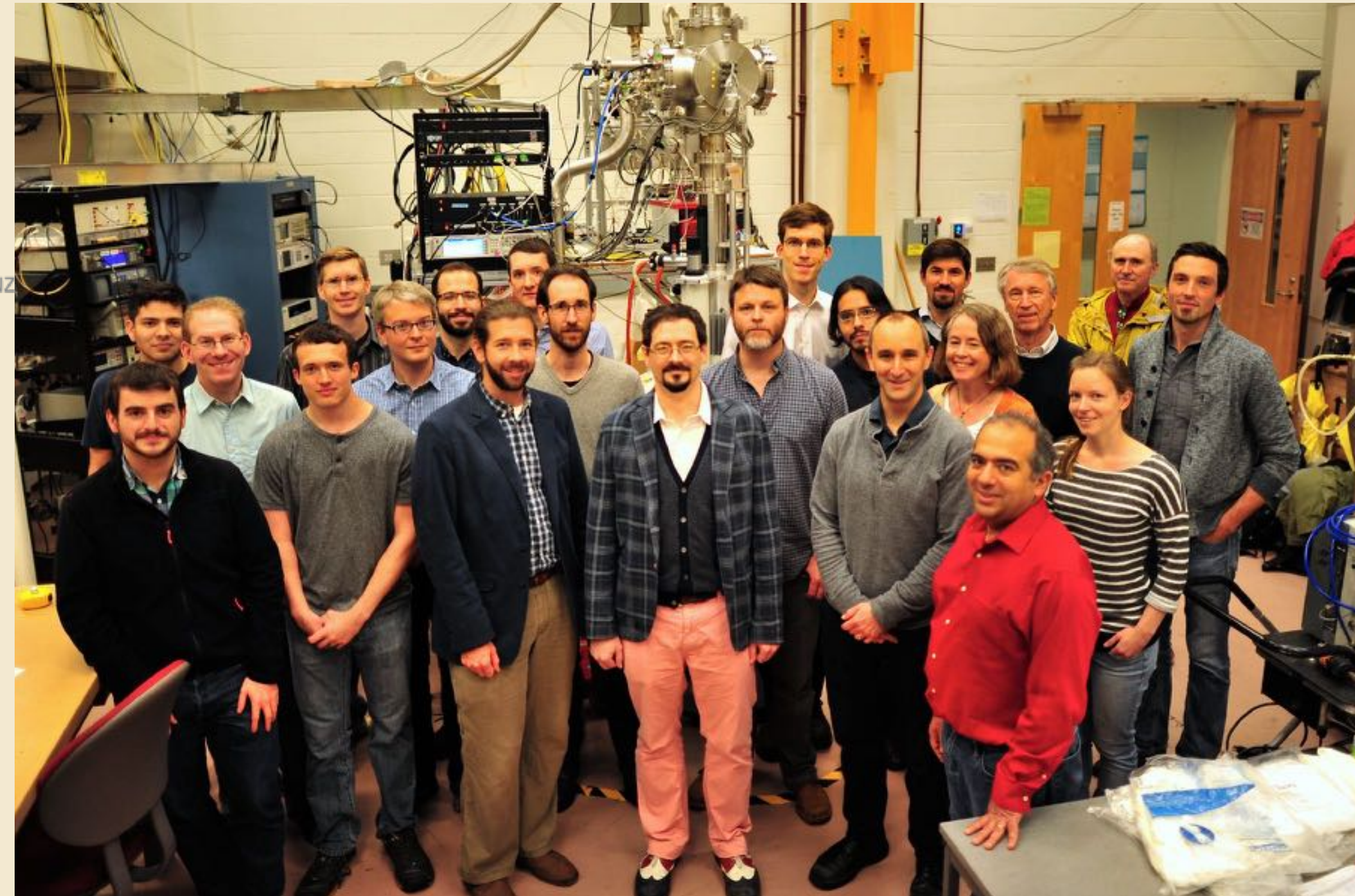
University of Washington

Karsten Heeger, James Nikkel, Luis Saldaña, Penny Slocum

Yale University



JOHANNES GUTENBERG
UNIVERSITÄT MAINZ



PennState



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