

PIs: Jerry P. Draayer, Kristina D. Launey, Tomas Dytrych; Post-Doc Alexis Mercenne Graduate Students: Alison C. Dreyfuss, Robert Baker, Grigor Sargsyan & David S. Kekejian

**U.S. NSF & DOE Sponsored Research** 

Symmetry Adapted No-Core Shell Model

(1 of 26)









Symmetry Adapted No-Core Shell Model

(4 of 26)

International School of Nuclear Physics (Erice, Sicily, 09/16–24/18) 40<sup>th</sup> Course: From Quarks to Nuclei and Stars









# First Results for <sup>6</sup>Li with N<sub>max</sub> = 10



... Proof of Principle ...

> 99% of Physics In < 1% of the "reorganized" space

JISP16\* bare interaction in Nmax = 10 space with  $\hbar\Omega$  = 20 MeV

### ... Team Work ...

Many "helps" along the way e.g., James Vary making his NCSM available to us, Mark Caprio (ND) visiting LSU on a sabbatical, along with quality input from Anna Hayes of LANL, various collaborators from Bulgaria, China, Mexico, and so on. Also to <u>NSF</u> for a PetaApps award, and <u>DOE</u> for an EPSCoR grant, plus SURA for release time and financial assistance with postdoc team!



Symmetry Adapted No-Core Shell Model

Ab Inito No-Core HPC

(9 of 26)



Next demonstrate (via many examples / results) that a collective state representation for  $|\Psi_{space}\rangle$  is a smart and wise choice for basis states! Looking ahead to work with HE/LE NP communities to integrate the symmetry adapted *ab initio* features of strong interaction in  $|\Psi_{spin}\rangle$ ?

Symmetry Adapted No-Core Shell Model



# **Creation of <sup>12</sup>C in hot stars**



# ... The elusive Hoyle state ...



Symmetry Adapted No-Core Shell Model

(11 of 26)



-24/18)

(12 of 26)

Symmetry Adapted

No-Core Shell Model









(15 of 26)

No-Core Shell Model

40<sup>th</sup> Course: From Quarks to Nuclei and Stars

ĹSIJ



40<sup>th</sup> Course: From Quarks to Nuclei and Stars





# Review Article – PPNP 89 (2016) 101-136

Progress in Particle and Nuclear Physics 89 (2016) 101-136



Contents lists available at ScienceDirect

Progress in Particle and Nuclear Physics

journal homepage: www.elsevier.com/locate/ppnp



#### Symmetry-guided large-scale shell-model theory



#### Kristina D. Launev<sup>a,\*</sup>, Tomas Dytrych<sup>a,b</sup>, Jerry P. Draaver<sup>a</sup>

<sup>a</sup> Department of Physics and Astronomy, Louisiana State University, Baton Rouge, LA 70803, USA <sup>b</sup> Nuclear Physics Institute, 250 68 Řež, Czech Republic

#### ARTICLE INFO

#### ABSTRACT

Article history: Available online 15 February 2016

Keywords:

Ab initio shell-model theory Symplectic symmetry Collectivity Clusters Hoyle state Orderly patterns in nuclei from first principles

In this review, we present a symmetry-guided strategy that utilizes exact as well as partial symmetries for enabling a deeper understanding of and advancing ab initio studies for determining the microscopic structure of atomic nuclei. These symmetries expose physically relevant degrees of freedom that, for large-scale calculations with OCD-inspired interactions, allow the model space size to be reduced through a very structured selection of the basis states to physically relevant subspaces. This can guide explorations of simple patterns in nuclei and how they emerge from first principles, as well as extensions of the theory beyond current limitations toward heavier nuclei and larger model spaces. This is illustrated for the ab initio symmetry-adapted no-core shell model (SA-NCSM) and two significant underlying symmetries, the symplectic Sp(3, R) group and its deformation-related SU(3) subgroup. We review the broad scope of nuclei, where these symmetries have been found to play a key role—from the light p-shell systems, such as <sup>6</sup>Li, <sup>8</sup>B, <sup>8</sup>Be, <sup>12</sup>C, and <sup>16</sup>O, and sd-shell nuclei exemplified by <sup>20</sup>Ne, based on first-principle explorations; through the Hoyle state in <sup>12</sup>C and enhanced collectivity in intermediate-mass nuclei, within a no-core shell-model perspective; up to strongly deformed species of the rare-earth and actinide regions, as investigated in earlier studies. A complementary picture, driven by symmetries dual to Sp(3, ℝ), is also discussed. We briefly review symmetry-guided techniques that prove useful in various nuclear-theory models, such as Elliott model, ab initio SA-NCSM, symplectic model, pseudo-SU(3) and pseudo-symplectic models, ab initio hyperspherical harmonics method, ab initio lattice effective field theory, exact pairing-plus-shell model approaches, and cluster models, including the resonating-group method. Important implications of these approaches that have deepened our understanding of emergent phenomena in nuclei, such as enhanced collectivity, giant resonances, pairing, halo, and clustering, are discussed, with a focus on emergent patterns in the framework of the ab initio SA-NCSM with no a priori assumptions.

© 2016 Elsevier B.V. All rights reserved.

## ... Also ...

Kristina Launey – "State of the Art in Nuclear Cluster Physics" (SOTANCP3) Yokohama, Japan May 26-30, 2014

## &

– Kristina Launey – "State of the Art in Nuclear Cluster Physics" (SOTANCP4) Galveston, Texas, USA May 13-18, 2018

LSU

Symmetry Adapted No-Core Shell Model

#### (18 of 26)





Symmetry Adapted No-Core Shell Model (19 of 26) International School of Nuclear Physics (Erice, Sicily, 09/16–24/18) 40<sup>th</sup> Course: From Quarks to Nuclei and Stars

ĹSIJ

# Further sd-shell Results (Robert Baker / GS)

sd-Shell Examples





Experiment ...... -17.7 8 shells ...... -19.3 (no effective charges)

LSU

Symmetry Adapted No-Core Shell Model

(21 of 26)

# Reaction Theory Reaction Theory (Alexis Mercenne / PDoc)

## Nucleosynthesis: Type I X-Ray Burst

TABLE 2 REACTIONS THAT IMPACT THE BURST LIGHT CURVE IN THE MULTI ZONE X-RAY BURST MODEL.

Rank	Reaction	Type <sup>a</sup>	$\operatorname{Sensitivity}^{\mathrm{b}}$	Category	Simulations for XRB are sensitive to certain
$egin{array}{c} 1 \\ 2 \\ 3 \\ 4 \\ 5 \\ 6 \end{array}$	$^{15}O(\alpha,\gamma)^{19}Ne$ $^{56}Ni(\alpha,p)^{59}Cu$ $^{59}Cu(p,\gamma)^{60}Zn$ $^{61}Ga(p,\gamma)^{62}Ge$ $^{22}Mg(\alpha,p)^{25}Al$ $^{14}O(\alpha,p)^{17}F$	D U D D D	16     6.4     5.1     3.7     2.3     5.8	1 1 1 1 1 1	<ul> <li>reaction rates</li> <li><sup>23</sup>Al(p,γ)<sup>24</sup>Si</li> <li>improve rate precision to improve simulations</li> </ul>
$7 \\ 8 \\ 9 \\ 10 \\ 11 \\ 12$		D U D U U U	$4.6 \\ 1.8 \\ 1.4 \\ 1.3 \\ 2.1 \\ 1.8$	$egin{array}{c} 1 \\ 1 \\ 2 \\ 2 \\ 2 \\ 2 \\ 2 \end{array}$	$ \begin{array}{c c} Ab \ initio \ wave \\ functions \\  \Psi\rangle \end{array} \end{array} \begin{array}{c} Widths \\ Phase \\ Shifts \end{array} \begin{array}{c} Cross \ sections \\ Reaction \ rates \end{array} $
13 14 15	${}^{17}F(\alpha,p)^{20}Ne$ ${}^{24}Mg(\alpha,\gamma)^{28}Si$ ${}^{57}Cu(p,\gamma)^{58}Zn$	U U D	$3.5 \\ 1.2 \\ 1.3 \\ 1.1$	2 2 2	Reaction network simulation
17 18 19	$^{17}{ m F}({ m p},\gamma)^{18}{ m Ne}$ $^{40}{ m Sc}({ m p},\gamma)^{41}{ m Ti}$ $^{48}{ m Cr}({ m p},\gamma)^{49}{ m Mn}$	U D D	1.1 1.7 1.1 1.2	2 2 2 2	Abundances

<sup>a</sup> Up (U) or down (D) variation that has the largest impact

<sup>b</sup>  $M_{LC}^{(i)}$  in units of 10<sup>38</sup> ergs/s

Symmetry Adapted No-Core Shell Model International School of Nuclear Physics (Erice, Sicily, 09/16-24/18) 40<sup>th</sup> Course: From Quarks to Nuclei and Stars

\_SU



No-Core Shell Model (23 of 26)

International School of Nuclear Physics (Erice, Sicily, 09/16-24/18) 40<sup>th</sup> Course: From Quarks to Nuclei and Stars





## "Collusion within & among Nucleons"



Symmetry Adapted No-Core Shell Model

(24 of 26)



- Robust stand-alone SA-NCSM code, publicly available (... tunable to available computational resources ...)
- Designed to handle up to 3-body & 4-body interactions (... important for alpha-particle type correlations ...)
- Push forward on the Sp-NCSM Hybrid model N<sup>plus</sup>[N<sub>full</sub>]
   (... importance of generating symmetry mixing...)
- Continue development with applications laptop version (<sup>6</sup>Li, <sup>12</sup>C, <sup>16</sup>O, <sup>20</sup>Ne, <sup>24</sup>Mg ... odd-A too & heavier ...)

## Next Up ... Deformed Versions ... DSA-NCSM & DSp-NCSM

International School of Nuclear Physics (Erice, Sicily, 09/16-24/18)

40<sup>th</sup> Course: From Quarks to Nuclei and Stars

Symmetry Adapted

No-Core Shell Model

(25 of 26)





40<sup>th</sup> Course: From Quarks to Nuclei and Stars

\_SU