

# I Review of relevant parts of Statistical Physics

Statistical physics: study of macroscopic systems with  $N \geq 10^{23}$  particles  
system is characterized by a microstate and a macrostate

microstate: defined by the location and momenta of all  $N$  particles (classical systems) or the  $N$ -body wave function (quantum mechanical systems)

macrostate: characterizes system by a few macroscopic observables like energy ( $E$ ), volume ( $V$ ), pressure ( $P$ ), temperature ( $T$ ), ...

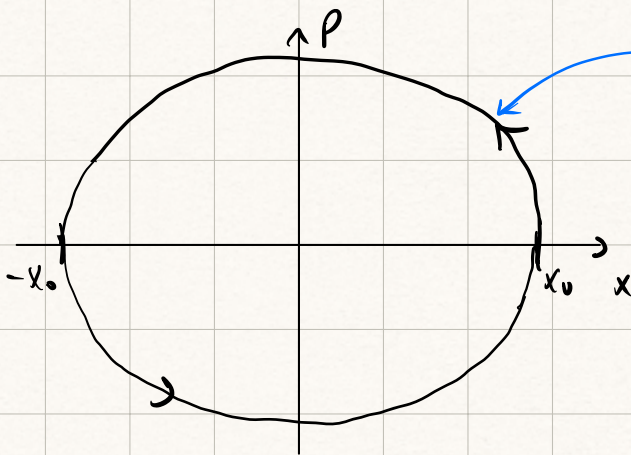
- Microstate is practically impossible to determine.
- Even if we would know this state, how can we solve  $N \sim 10^{23}$  coupled different equations and extract useful information?
- Microstate changes all the time due to interaction processes with environment. These interactions are usually weak enough such that only the microstate changes, but the macrostate remains the same
- basic idea of statistical physics:

Consider all microstates at the same time that are consistent with certain values of macro observables.  $\Rightarrow$  ensembles

ensemble: a large number of virtual copies of a system, where each copy represents a possible microstate the system can be in

phase space (in classical physics) : consider one particle in 1 dim

harmonic oscillator :  $x(t) = x_0 \cos(\omega t)$   
 $p(t) = -m x_0 \omega \sin(\omega t)$



phase space trajectory corresponds to a fixed energy of the system

$$E = \frac{p^2}{2m} + \frac{1}{2} m \omega^2 x^2$$
$$= \frac{1}{2} m x_0^2 \omega^2$$

- state of system at a given time is a point in a 2d phase space
- for  $N$  particles in 3d phase space has  $6N$  dimensions
- MANY different points (microstates) correspond to a macrostate of a given energy
- introduce probability distribution in phase space

$$g(x, p, t) dx dp = g(x_1, x_2, \dots, x_{3N}, p_1, p_2, \dots, p_{3N}, t) \prod_{i=1}^{3N} dx_i dp_i$$

conservation law :  $\frac{\partial g}{\partial t} = -\nabla \cdot v g$  (Liouville equation)

Question: What is  $g(x, p, t)$ ?

## Microcanonical ensemble and partition function

consider isolated macroscopic system  
with  $N$  particles, volume  $V$   
and energy  $E$

$N, V, E$

in reality the total energy is given by a

range  $[E - \Delta E, E]$

macroscopic

microscopic (e.g. fluctuations due to interactions  
with environment)

### Fundamental postulate of statistical physics

All microstates with energy in range  $[E - \Delta E, E]$   
have the same probability (density):

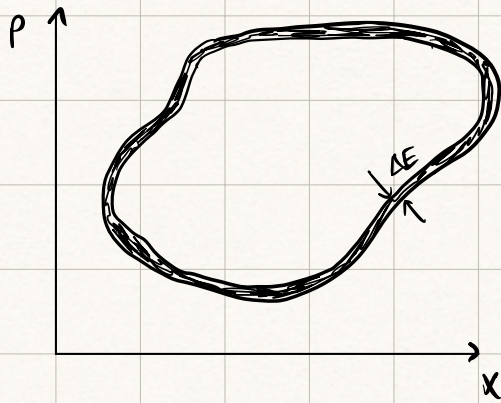
$$p_{mc} = \begin{cases} \frac{1}{Z_{mc}} & \text{for } E - \Delta E \leq H(p, x) \leq E \quad (E - \Delta E \leq H(\omega) \leq E(\omega)) \\ 0 & \text{else} \end{cases}$$

$Z_{mc}$  is a phase space volume in classical physics and  
a discrete number of states in QM.

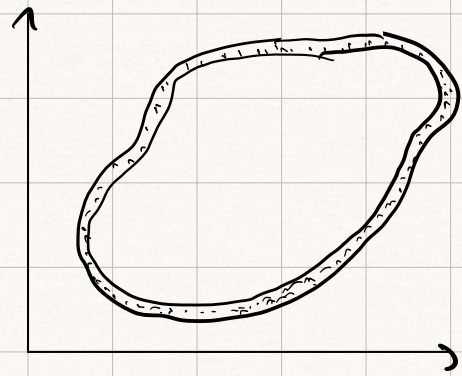
Microcanonical partition function:

$$Z_{mc} = \text{Tr} [\Theta(E - H) - \Theta(H - (E - \Delta E))]$$

classical



Qn



One can show that  $Z_{mc}$  is basically independent of  $\Delta E$   
since the phase space increases VERY rapidly with energy:  $Z_{mc} \sim E^N$   
 $\Rightarrow$  For  $N \approx 10^{23}$  the entire phase space is dominated by  
the upper limit of the energy shell

$\hookrightarrow$  exercise

### density matrix, entropy and expectation values

consider QM system in a microstate  $|\psi_i\rangle$  with probability  $p_i$

$\Rightarrow$  then the density matrix is defined by

$$\rho = \sum_i p_i |\psi_i\rangle\langle\psi_i|$$

$\hookrightarrow$  contains inherent QM uncertainties  
describes statistical uncertainties (ensembles)

$\Rightarrow$  expectation value of operator  $A$ :

$$\langle A \rangle = \sum_i p_i \langle \psi_i | A | \psi_i \rangle$$

$$= \sum_i p_i \sum_n \langle \psi_i | u_n \rangle \langle u_n | A | \psi_i \rangle$$

$$\begin{aligned}
&= \sum_i p_i \sum_n \langle n | A | \psi_i \rangle \langle \psi_i | n \rangle \\
&= \sum_n \langle n | A | n \rangle \\
&= \sum_n \langle n | g A | n \rangle = \text{Tr} (g \cdot A)
\end{aligned}$$

↳ short hand notation for sum over complete set of  $N$ -body states

$$\text{Tr} = \int \frac{d^{3N} x \, d^{3N} p}{h^{3N} N!}$$

semiclassical approx. for  $N$  identical particles

microcanonical density matrix:

$$S_{mc} = \sum_n p_{mc}(E_n) |n\rangle \langle n| = \sum_n \frac{1}{Z_{mc}(E_n)} |n\rangle \langle n|$$

$$\text{Tr} S_{mc} = 1 \quad (\text{normalization of total probability})$$

Definition of entropy (general):

$$S = -k_B \text{Tr} (S \log S) = -k_B \langle \log S \rangle$$

for the microcanonical ensemble:

$$\begin{aligned}
S_{mc} &= -k_B \sum_n p(E_n) \log(p(E_n)) \\
&= -k_B \sum_n Z_{mc}^{-1}(E_n) \log[Z_{mc}^{-1}(E_n)] \\
&= +k_B \log Z_{mc}(E_n)
\end{aligned}$$

↳ Boltzmann entropy

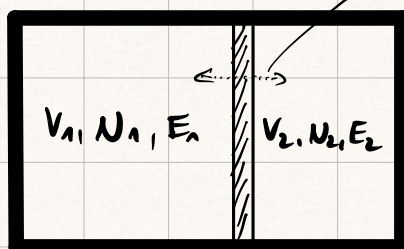
## Temperature

consider a system consisting of two subsystems,  
total system isolated

$$\text{Hamiltonian: } H = H_1 + H_2 + W$$



Small correction to  
total energy of system,  
but allows energy exchange  
between subsystems



energy  
exchange

probability of microstates of system 1+2 with energy  $E = E_1 + E_2$

$$p_{mc}(E) = \begin{cases} \frac{1}{Z_{mc}(E)} & \text{for } E - \Delta E \leq H_1 + H_2 + W \leq E \\ & ((E - \Delta E)/\ln) \leq (H_1 + H_2 + W)/\ln \leq E/\ln \\ 0 & \text{else} \end{cases}$$

probability for subsystem 1 to have energy in range  $[E_1 - \Delta E_1, E_1]$ ,  
and hence subsystem 2 to have energy in range  $[E_2 - \Delta E_2, E_2]$ :

$$p(E_1) = \text{Tr}'_{E_1} \rho_{mc}^{(1+2)}(E) \equiv \text{Tr} \rho_{mc}^{(1+2)}(E) \delta_{E_1, H_1}$$

↓  
restricted trace

↓  
consider only those states  
for which subsystem 1 has  
energy in given range

$$\begin{aligned}
\Rightarrow p(E_1) &= \text{Tr} \delta_{mc}^{(1+2)}(E) \delta_{E_1, H_1} \\
&= \text{Tr} \frac{1}{z_{mc}^{(1+2)}(E)} \delta_{E_1, H_1 + H_2} \delta_{E_1, H_1} \\
&= \text{Tr}_1 \text{Tr}_2 \frac{1}{z_{mc}^{(1+2)}(E)} \delta_{E_1, H_1} \delta_{E-E_1, H_2} \\
&= \frac{z_{mc}^{(1)}(E_1) z_{mc}^{(2)}(E-E_1)}{z_{mc}^{(1+2)}(E)}
\end{aligned}$$

determine the most probable energy  $\bar{E}_1$ :  $\left. \frac{dp(E_1)}{dE_1} \right|_{E_1=\bar{E}_1} = 0$

$$\Rightarrow \left. \frac{\partial z_{mc}^{(1)}(E_1)}{\partial E_1} \right|_{E_1=\bar{E}_1} z_{mc}^{(2)}(E-E_1) + z_{mc}^{(1)}(E_1) \left. \frac{\partial z_{mc}^{(2)}(E-E_1)}{\partial E_1} \right|_{E_1=\bar{E}_1} = 0$$

$$\Rightarrow \underbrace{\left. \frac{\partial \log z_{mc}^{(1)}(E_1)}{\partial E_1} \right|_{E_1=\bar{E}_1}}_{\frac{\partial S_1(E_1)}{\partial E_1} = T_1^{-1}} = \underbrace{\left. \frac{\partial \log z_{mc}^{(2)}(E_2)}{\partial E_2} \right|_{E_2=E-\bar{E}_1}}_{\frac{\partial S_2(E_2)}{\partial E_2} = T_2^{-1}}$$

$$\frac{\partial S_1(E_1)}{\partial E_1} = T_1^{-1}$$

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Definition of temperature:

$$T^{-1} = \frac{\partial S(E)}{\partial E}$$