

Lorentz transformations

- ▶ The laws of nature are invariant under Poincaré transformations
 - = inhomogeneous Lorentz transformations: $x'^\mu = \Lambda^\mu{}_\nu x^\nu + a^\mu$
 - = (homogeneous) Lorentz transformations + translations in space and time
- ▶ Lorentz transformations:
 - = Lorentz boosts + rotations + parity transformations + time reversal
 - proper orhochronous Lorentz transformations
- ▶ Example: Boost along the x axis

$$(\Lambda^\mu{}_\nu) = \begin{pmatrix} \cosh \chi & -\sinh \chi & 0 & 0 \\ -\sinh \chi & \cosh \chi & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$$

- Lorentz transformation of the covariant components:

$$x'^\mu = \Lambda_\mu^\nu x_\nu \quad (= \text{definition of } \Lambda_\mu^\nu)$$

$$= g_{\mu\alpha} x'^\alpha = g_{\mu\alpha} \Lambda^\alpha_\beta x^\beta = g_{\mu\alpha} \Lambda^\alpha_\beta g^{\beta\nu} x_\nu \Rightarrow \boxed{\Lambda_\mu^\nu = g_{\mu\alpha} \Lambda^\alpha_\beta g^{\beta\nu}}$$

i.e., the indices of Λ can be raised and lowered by the metric tensor as well.

- Invariance of the proper time:

$$\Lambda_\mu^\nu x_\nu \Lambda^\mu_\lambda x^\lambda = x'_\mu x'^\mu \stackrel{!}{=} x_\nu x^\nu = x_\nu g^\nu_\lambda x^\lambda \Rightarrow \boxed{\Lambda_\mu^\nu \Lambda^\mu_\lambda = g^\nu_\lambda = \delta_\lambda^\nu}$$

- backtransformation:

$$x^\nu = g^\nu_\lambda x^\lambda = \Lambda_\mu^\nu \Lambda^\mu_\lambda x^\lambda = \Lambda_\mu^\nu x'^\mu \Rightarrow \boxed{x^\nu = x'^\mu \Lambda_\mu^\nu}$$

analogously:

$$\boxed{x_\nu = x'_\mu \Lambda^\mu_\nu}$$

Four-vectors

- ▶ Contra- and covariant four-vectors
 - = objects a with four components which under Lorentz transformations behave like (x^μ) or (x_μ) , respectively:

$$\begin{aligned} a'^\mu &= \Lambda^\mu_{\nu} a^\nu, & a^\nu &= a'^\mu \Lambda_\mu^{\nu}, \\ a'_\mu &= \Lambda_\mu^{\nu} a_\nu, & a_\nu &= a'_\mu \Lambda^\mu_\nu, \end{aligned}$$

- ▶ Scalar product: $a \cdot b \equiv a^\mu b_\mu = a_\mu b^\mu$
 $\Rightarrow a' \cdot b' = a \cdot b$ (because $x'^\mu x'_\mu = x^\mu x_\mu$)
- ▶ Example: **four-momentum** $p = (p^\mu) = \begin{pmatrix} E/c \\ \vec{p} \end{pmatrix}$
with the relativistic energy $E = \sqrt{m^2 c^4 + \vec{p}^2 c^2}$
 $\Rightarrow p^\mu p_\mu = \frac{E^2}{c^2} - \vec{p}^2 = m^2 c^2 \quad \checkmark$

► Four-gradients:

$$\begin{aligned} \frac{\partial}{\partial x'^\mu} &= \frac{\partial x^\nu}{\partial x'^\mu} \frac{\partial}{\partial x^\nu} = \Lambda_\mu^\nu \frac{\partial}{\partial x^\nu} && \text{like } x_\mu &\rightarrow \text{covariant} &\rightarrow \frac{\partial}{\partial x^\mu} \equiv \partial_\mu \\ \frac{\partial}{\partial x'_\mu} &= \frac{\partial x_\nu}{\partial x'_\mu} \frac{\partial}{\partial x_\nu} = \Lambda^\mu_\nu \frac{\partial}{\partial x_\nu} && \text{like } x^\mu &\rightarrow \text{contravariant} &\rightarrow \frac{\partial}{\partial x_\mu} \equiv \partial^\mu \end{aligned}$$

► Relation to the usual three-gradient:

$$\begin{aligned} \nabla^k &= \frac{\partial}{\partial x^k} = -\frac{\partial}{\partial x_k} \\ \Rightarrow (\partial_\mu) &= \begin{pmatrix} \frac{\partial}{\partial x^0} \\ (\frac{\partial}{\partial x^k}) \end{pmatrix} = \begin{pmatrix} \frac{1}{c} \frac{\partial}{\partial t} \\ \vec{\nabla} \end{pmatrix}, \quad (\partial^\mu) = \begin{pmatrix} \frac{\partial}{\partial x_0} \\ (\frac{\partial}{\partial x_k}) \end{pmatrix} = \begin{pmatrix} \frac{1}{c} \frac{\partial}{\partial t} \\ -\vec{\nabla} \end{pmatrix} \end{aligned}$$

► d'Alambert operator: $\square \equiv \partial_\mu \partial^\mu = \frac{1}{c^2} \frac{\partial^2}{\partial t^2} - \vec{\nabla}^2$ Lorentz scalar!

► Tensors of rank 2: $A'^{\mu\nu} = \Lambda^\mu_\alpha \Lambda^\nu_\beta A^{\alpha\beta}$ (like $x^\mu x^\nu$)

Examples: fieldstrength tensor $F^{\mu\nu}$, $g^{\mu\nu}$, Λ^μ_ν

► Tensors of rank n : $A'^{\mu_1 \dots \mu_n} = \Lambda^{\mu_1}_{\nu_1} \dots \Lambda^{\mu_n}_{\nu_n} A^{\nu_1 \dots \nu_n}$

Classification of Lorentz transformations

- ▶ Using $\Lambda_\mu^\nu \Lambda^\mu_\lambda = \delta_\lambda^\nu$ (see above) one can show:
 - ▶ $\det(\Lambda_\mu^\nu) = \pm 1$
 - ▶ $\Lambda_0^0 \geq 1$ or $\Lambda_0^0 \leq -1$
- classification of Lorentz transformations by the sign of $\det(\Lambda_\mu^\nu)$ and Λ_0^0
- ▶ Boosts and rotations: $\det(\Lambda_\mu^\nu) = +1, \quad \Lambda_0^0 \geq 1$
(can be generated continuously from the identity)
- ▶ parity transform.: $x' = \begin{pmatrix} ct' \\ \vec{x}' \end{pmatrix} = \begin{pmatrix} ct \\ -\vec{x}' \end{pmatrix} \Rightarrow (\Lambda^\mu_\nu) = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix}$
 $\Rightarrow \det(\Lambda_\mu^\nu) = -1, \quad \Lambda_0^0 \geq 1$
- ▶ time reversal: $x' = \begin{pmatrix} ct' \\ \vec{x}' \end{pmatrix} = \begin{pmatrix} -ct \\ \vec{x}' \end{pmatrix} \Rightarrow (\Lambda^\mu_\nu) = \begin{pmatrix} -1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$
 $\Rightarrow \det(\Lambda_\mu^\nu) = -1, \quad \Lambda_0^0 \leq -1$

3.3 The Klein-Gordon equation



- ▶ Analogous procedure to the “derivation” of the Schrödinger equation:

- ▶ relativistic energy-momentum relation: $E^2 = \vec{p}^2 c^2 + m^2 c^4$

- ▶ replace by operators: $E \rightarrow i\hbar \frac{\partial}{\partial t}$, $\vec{p} \rightarrow \frac{\hbar}{i} \vec{\nabla}$

$$\Rightarrow -\hbar^2 \frac{\partial^2}{\partial t^2} \Phi(\vec{r}, t) = \left(-\hbar^2 c^2 \vec{\nabla}^2 + m^2 c^4 \right) \Phi(\vec{r}, t)$$

“Klein-Gordon equation” (Schrödinger, Fock, Klein, Gordon 1926)

- ▶ partial diff. equation of 2nd order in time and spatial derivatives
(Schrödinger equation: 1st order in time, 2nd order in space)

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- ▶ partial diff. equation of 2nd order in time and spatial derivatives
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- ▶ Lorentz invariant form:

$$\Leftrightarrow \left[\frac{1}{c^2} \frac{\partial^2}{\partial t^2} - \vec{\nabla}^2 + \left(\frac{mc}{\hbar} \right)^2 \right] \Phi(\vec{r}, t) = 0$$

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$$\Leftrightarrow \left[\square + \left(\frac{mc}{\hbar} \right)^2 \right] \Phi(x) = 0$$

$$\frac{\hbar}{mc} = \frac{\hbar c}{mc^2} \equiv \lambda_C \quad \text{“Compton wavelength”}$$

► plane-wave ansatz for the solution: $\Phi(t, \vec{x}) = \mathcal{N} e^{-\frac{i}{\hbar}(Et - \vec{p} \cdot \vec{x})}$

$$\Rightarrow \left[\frac{1}{c^2} \frac{\partial^2}{\partial t^2} - \vec{\nabla}^2 + \left(\frac{mc}{\hbar} \right)^2 \right] \Phi(t, \vec{x}) = \left[-\frac{E^2}{\hbar^2 c^2} + \frac{\vec{p}^2}{\hbar^2} + \frac{m^2 c^2}{\hbar^2} \right] \Phi(t, \vec{x}) \stackrel{!}{=} 0$$

$$\Rightarrow E^2 = m^2 c^4 + \vec{p}^2 c^2 \quad \text{relativistic energy-momentum relation} \quad \checkmark$$

► using four-vectors:

$$x = \begin{pmatrix} ct \\ \vec{x} \end{pmatrix}, \quad p = \begin{pmatrix} \frac{E}{c} \\ \vec{p} \end{pmatrix} \quad \Rightarrow \quad p \cdot x = Et - \vec{p} \cdot \vec{x}$$

$$\Rightarrow \text{The solution is Lorentz invariant: } \Phi(t, \vec{x}) = \mathcal{N} e^{-\frac{i}{\hbar} p \cdot x} \equiv \Phi(x)$$

insert into the Klein-Gordon equation:

$$\left[\square + \left(\frac{mc}{\hbar} \right)^2 \right] \Phi(x) = \left[\partial_\mu \partial^\mu + \left(\frac{mc}{\hbar} \right)^2 \right] \Phi(x) = \left[-\frac{p^2}{\hbar^2} + \frac{m^2 c^2}{\hbar^2} \right] \Phi(x) \stackrel{!}{=} 0$$

$$\Rightarrow p^2 = \frac{E^2}{c^2} - \vec{p}^2 \stackrel{!}{=} m^2 c^2 \quad \Leftrightarrow \quad E^2 = m^2 c^4 + \vec{p}^2 c^2 \quad \checkmark$$

Current conservation

► Klein-Gordon equation: $\left[\square + \left(\frac{mc}{\hbar} \right)^2 \right] \Phi(x) = 0$

complex-conjugate equation: $\left[\square + \left(\frac{mc}{\hbar} \right)^2 \right] \Phi^*(x) = 0$

$$\begin{aligned} \Rightarrow 0 &= \Phi^*(x) \square \Phi(x) - \Phi(x) \square \Phi^*(x) \\ &= \Phi^*(x) \partial_\mu \partial^\mu \Phi(x) - \Phi(x) \partial_\mu \partial^\mu \Phi^*(x) \\ &= \partial_\mu [\Phi^*(x) \partial^\mu \Phi(x) - \Phi(x) \partial^\mu \Phi^*(x)] \quad \Rightarrow \quad \boxed{\partial_\mu j^\mu(x) = 0} \end{aligned}$$

with the **conserved 4-current**

$$\begin{aligned} j^\mu(x) &= \alpha (\Phi^*(x) \partial^\mu \Phi(x) - \Phi(x) \partial^\mu \Phi^*(x)) \\ &\equiv \alpha \Phi^*(x) \left(\partial^\mu - \overleftarrow{\partial}^\mu \right) \Phi(x), \quad \alpha: \text{arbitrary constant} \end{aligned}$$

► time- and space-like components: $(j^\mu) = \begin{pmatrix} j^0 \\ \vec{j} \end{pmatrix} \equiv \begin{pmatrix} c\rho \\ \vec{j} \end{pmatrix}$

$$\Rightarrow 0 = \partial_\mu j^\mu = \partial_0 j^0 + \partial_k j^k = \frac{1}{c} \frac{\partial}{\partial t} c\rho + \vec{\nabla} \cdot \vec{j}$$

$$\Rightarrow \boxed{\frac{\partial \rho}{\partial t} + \vec{\nabla} \cdot \vec{j} = 0} \quad \text{continuity equation}$$

interpretation in the nonrelativistic case: ρ = probability density
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- ▶ determination of the constant: $\vec{j} = \vec{j}_{\text{Schrödinger}}^!$

$$\vec{j} \equiv (j^k) = \alpha (\Phi^*(\partial^k)\Phi - \Phi(\partial^k)\Phi^*)$$

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$$\vec{j} \equiv (j^k) = \alpha (\Phi^*(-\vec{\nabla})\Phi - \Phi(-\vec{\nabla})\Phi^*) \stackrel{!}{=} \frac{\hbar}{2mi} (\Phi^* \vec{\nabla} \Phi - \Phi \vec{\nabla} \Phi^*)$$

$$\Rightarrow \alpha = \frac{i\hbar}{2m}$$