

So a_p creates a particle of momentum p
 But it's spread over whole universe. Not something you can
 actually do.

You can make something local in both x and p space

Consider $f(x)$ ~~with dual~~ with dual $\tilde{f}(p)$.

Single-particle state $|f\rangle = \sum_{\vec{p} = \frac{2\pi}{L}\vec{n}} \tilde{f}(p) a_p^\dagger |0\rangle$ sensible.

Normalization: $\sum_p \tilde{f}^* \tilde{f}(p) = 1$. Each $\tilde{f}(p) \rightarrow 0$ as $L \rightarrow \infty$
~~for some fixed $f(x)$. sad.~~

Means function $f(x)$ must get smaller as I make L bigger???

I would like to change normalization convention

so things will be written in terms of

$$\sum_{\vec{p} = \frac{2\pi}{L}\vec{n}} L^{-3} \Rightarrow \int \frac{d^3p}{(2\pi)^3} \quad (\text{sometimes written } \int d^3p)$$

and such that $\tilde{f}(p) \rightarrow$ something finite in the $L \rightarrow \infty$ limit

Only way: change normalization ~~of~~ I use for a_p^\dagger .

I wanna do that anyways, since $[a_p, a_q^\dagger] = \delta_{pq}$ Kronecker δ
Not limit of Dirac delta.

So I replace:

old normalization

$$a_{old,p} = \frac{1}{\sqrt{2\omega_p L^3}} \int d^3x e^{-i\vec{p}\cdot\vec{x}} (\omega_p \phi(x) + i\pi \dot{\phi}(x))$$

state: $|f\rangle = \sum_{\vec{p}=\frac{2\pi\vec{n}}{L}} \tilde{f}_{old}(\vec{p}) a_{old,p}^\dagger |0\rangle$

normalization: $\sum_{\vec{p}} \tilde{f}_{old}^* \tilde{f}_{old}(\vec{p}) = 1$

$$[a_p, a_q^\dagger] = \delta_{p,q}$$

Good: $\|a^\dagger|0\rangle\| = 1$ well normalized

Bad: to build wave packets, need $f(x) \propto \frac{1}{L^{3/2}}$ funny.

new normalization

$$a_{new,p} = \frac{1}{\sqrt{2\omega_p}} \int d^3x e^{-i\vec{p}\cdot\vec{x}} (\omega_p \phi(x) + i\pi \dot{\phi}(x))$$

$$= L^{3/2} a_{old}$$

$|f\rangle = \sum_{\vec{p}=\frac{2\pi\vec{n}}{L}} L^{-3} \tilde{f}_{new}(\vec{p}) a_{new,p}^\dagger |0\rangle$

$$\tilde{f}_{new}(\vec{p}) = L^{3/2} \tilde{f}_{old}, \text{ so}$$

$$\sum_{\vec{p}} L^{-3} \tilde{f}_{new}^* \tilde{f}_{new}(\vec{p}) = 1 = \int \frac{d^3p}{(2\pi)^3} f_{new}^* f_{new}(\vec{p})$$

$$[a_p, a_q^\dagger] = L^3 \delta_{p,q}$$

$$\approx (2\pi)^3 \delta^3(\vec{p}-\vec{q})$$

Good: $L \rightarrow \infty$ limit now transparent.

Bad: $\|a^\dagger|0\rangle\| = L^3$ diverges

Must use wave packets to get physically reasonable states. (which is OK!)

Actually I will shift normalization even a bit more!

LS P3

$$a_{\text{new } \vec{p}} = L^{3/2} \sqrt{2\omega_{\vec{p}}} a_{\text{old}} \quad \text{extra factor of } \sqrt{2\omega_{\vec{p}}}$$

Then:

$$\psi(x) = \int \frac{d^3p}{(2\pi)^3 2\omega_{\vec{p}}} \left[a_{\vec{p}} e^{i\vec{p}\cdot\vec{x}} + a_{\vec{p}}^\dagger e^{-i\vec{p}\cdot\vec{x}} \right]$$

$$[a_{\vec{p}}, a_{\vec{p}'}^\dagger] = 2\omega_{\vec{p}} (2\pi)^3 \delta^3(\vec{p} - \vec{p}')$$

so $\langle p | p' \rangle = 2\omega_{\vec{p}} (2\pi)^3 \delta^3(\vec{p} - \vec{p}')$ if $|p'\rangle \equiv a_{\vec{p}'}^\dagger |0\rangle$

$$|f\rangle = \int \frac{d^3p}{(2\pi)^3 2\omega_{\vec{p}}} f_{\vec{p}} |p\rangle \quad \text{where } \int \frac{d^3p}{(2\pi)^3 2\omega_{\vec{p}}} f_{\vec{p}}^* f_{\vec{p}} = 1$$

correctly normalizes a wave packet.

why?

$$\int \frac{d^4p}{(2\pi)^4} (2\pi) \delta(p^2 - m^2) \Theta(p^0) \quad \text{is Lorentz invariant}$$

obviously Lorentz inv
 Lorentz inv inv. when next to Θ , explain

$$= \int \frac{d^3p}{(2\pi)^3} \underbrace{\int \frac{dp^0}{2\pi} 2\pi \delta(p^0^2 - \omega_{\vec{p}}^2) \Theta(p^0)}_{\frac{1}{2\omega_{\vec{p}}}} \Theta(p^0)$$

$\frac{1}{2\omega_{\vec{p}}}$ why?

So this normalization has smooth $\hbar \rightarrow \infty$ limit & is naturally Covariant.

Time evolution

Heisenberg picture $[H, a_p] = -\omega_p a_p$ easy to show

$$a_p(t) = e^{iHt} a_p e^{-iHt} = e^{-i\omega_p t} a_p$$

call $\omega_p = p^0$ or vice versa
nice, as $p^2 = m^2$ then,
usual relativ. relation.

then
$$\psi(x,t) = \int \frac{d^3p}{(2\pi)^3 2p^0} (e^{-ip_\mu x^\mu} a_p + e^{ip_\mu x^\mu} a_p^\dagger)$$

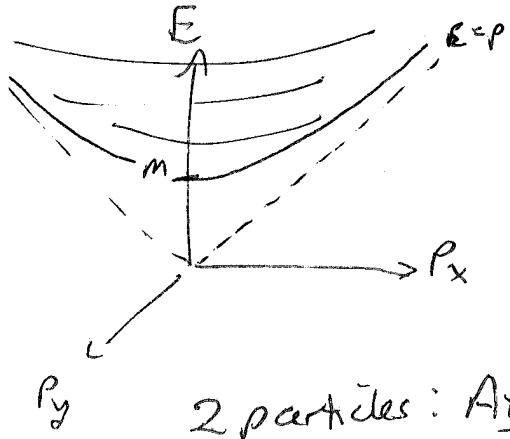
Oh look, using Heisenberg picture, $\psi(x^\mu)$ has covariant expression

Heisenberg is way to go in QFT.

This thing is a thing of single particles & multi-particle states.

Is there really an absolute difference? Yes!

Single particle: $E = \sqrt{p^2 + m^2}$. At rest, $E = m$

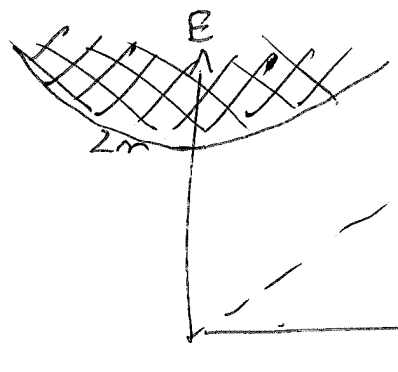


hyperbolic surface of revolution,
"the mass shell" in E-p space.

2 particles: At Rest $E = \sqrt{\vec{p}_1^2 + m^2} + \sqrt{\vec{p}_2^2 + m^2}$ with $\vec{p}_1 = -\vec{p}_2$ (cm)

(CM frame)
But $|\vec{p}|$ can be anything. $\vec{p}^2 \in [0, \infty]$ any value.

$E \geq 2m$ but can take any value $> 2m$.



Fills the interior of a "mass shell"
physically different

We see this in 2-point correlation function

Define

$$G_{\phi\phi}^{\rightarrow}(x) = \langle 0 | \phi(x) \phi(0) | 0 \rangle \quad 4\text{-vector notation } x^\mu.$$

$$\tilde{G}_{\phi\phi}^{\rightarrow}(q) \equiv \int d^4x e^{iq \cdot x} G_{\phi\phi}^{\rightarrow}(x) \quad \text{its momentum-space version}$$

Just compute it

$$G^{\rightarrow}(x) = \langle 0 | \int \frac{d^3p d^3k}{(2\pi)^3 2p^0 (2\pi)^3 2k^0} (a_p e^{-ip \cdot x} + a_p^\dagger e^{+ip \cdot x}) (a_k e^{i k \cdot 0} + a_k^\dagger e^{i k \cdot 0}) | 0 \rangle$$

$$= \int \frac{d^3p}{(2\pi)^3 2p^0} e^{-ip \cdot x} \cdot G^{\rightarrow}(q) = 2\pi \delta(q^2 - m^2) \Theta(q^0)$$

mass-shell requirement

$\phi(0)$ makes a particle (a_k^\dagger)

It propagates in mass-shell respecting fashion

$\phi(x)$ catches it & sees how it propagated to x .

~~Note, for $x^2 < 0$, $G^{\rightarrow}(x) = 0$ identically~~

Can I do this integral? yes!

for $x^2 < 0$: use Lorentz, Luke. Go to frame $x^0 = 0$ $|\vec{x}| = -\sqrt{x^2} \equiv r$ put it on z -axis.

$$G^{\rightarrow}(r) = \int e^{i\vec{p} \cdot \vec{r}} \frac{d^3p}{(2\pi)^3 2\sqrt{p^2 + m^2}} = \frac{1}{16\pi^3} \int \frac{p^2 dp}{\sqrt{p^2 + m^2}} \int_0^{2\pi} d\phi \int_0^\pi d\theta \sin\theta e^{ipr \cos\theta}$$

$$= \frac{1}{4\pi^2 r} \int_0^\infty \frac{p \sin(pr)}{\sqrt{p^2 + m^2}} dp \quad \left(\int_{-1}^1 d\cos\theta = \frac{2}{pr} \sin pr \right)$$

$$= \frac{m}{4\pi^2 r} K_1(mr) \quad \text{Modified Bessel Function} \rightarrow \begin{cases} \frac{1}{4\pi^2 r^2} & mr \ll 1 \\ \frac{m^{1/2}}{2^{5/2} \pi^{3/2} \Gamma^{3/2}} e^{-mr} & mr \gg 1 \end{cases}$$

$$K_1(x) = \frac{\sqrt{\pi}}{2} e^{\frac{1}{2}ix} H_{1+i}^{(1)}(ix) \quad \text{Hankel Contour}$$

Creepy part: $G^2(r) \rightarrow \frac{1}{4\pi^2 r^2}$ at small r .

$\langle \varphi^2 \rangle$ is like square of fluct. in fields.

$$\text{But: } \langle \varphi^2 \rangle = \lim_{r \rightarrow 0} \langle \varphi(r) \varphi(0) \rangle$$

$$= \lim_{r \rightarrow 0} \frac{1}{4\pi^2 r^2} \text{ eek!}$$

At scale r , fluct. in field have size $\frac{1}{2\pi r}$

The $\frac{1}{2\pi}$ is nice - effects of fluct. will often have $(\frac{1}{2\pi})^2$

The $\frac{1}{r}$ means field picture really weird.

Don't Ask "what is fluct. in φ at a point?"

Do ask "fluct at some scale r ?"

say, smear field $\varphi_r \equiv \int_{(2\pi)^{3/2} r^3} d^3x e^{-x^2/2r^2} \varphi(x)$ (so normalized to 1)

$$\langle \varphi_r^2 \rangle = \int \frac{d^3x d^3y}{(2\pi)^3 r^6} e^{-\frac{(x-y)^2}{2r^2}} \frac{1}{4\pi^2 (x-y)^2}$$

$$= \frac{1}{8\pi^2 r^2} \text{ yup. Gets bigger the smaller you make } r.$$

Oh, and $G^2(t) =$ analytic continuation of $G^2(r)$

$$= \frac{m}{8\pi^2 t} H_1(t) \text{ Hankel function}$$