

$$\mathcal{L} = \frac{1}{2} (\partial_\mu \phi)^2 - \frac{m^2}{2} \phi^2 - \frac{\lambda}{24} \phi^4$$

What do I know about this theory?

- 1 - Is it Harmonic oscillators?
- 2 - Are there particles?
- 3 - Does it exist at all?

Totally not clear!

- 1 - NO
- 2 - Yes, but see 3
- 3 - Yes for finite UV cutoff  
but limit  $a \rightarrow 0$  ( $\Lambda \rightarrow \infty$ ) is subtle - theory "loses"  $\lambda \phi^4$  term

How can I learn anything about such a theory?

What if I could somehow compute  $\langle \phi(x) \phi(0) \rangle = G^2(x)$ .

What would it teach me? (Free thg - knew about particles)

Or more generally, for some <sup>local</sup> operator  $\mathcal{O}$  in QFT, what do we learn from

$$G_{\mathcal{O}\mathcal{O}}^2(x) = \langle 0 | \mathcal{O}^\dagger(x) \mathcal{O}(0) | 0 \rangle \quad ?$$

Well, assume there are states (maybe contin or discrete...)

organize as e states of  $P^\mu$  ( $P^\mu = \int T^{\mu 0} d^3x$ )  $|n\rangle : P^\mu |n\rangle = p_n^\mu |n\rangle$

Assume  $P_n^0 > 0$  except for vacuum ("Mass Gap")

$$\mathcal{O}(0) |0\rangle = \underbrace{\sum_n |n\rangle \langle n | \mathcal{O}(0) |0\rangle}_{C_{\mathcal{O},n} \text{ some } \mathbb{C} \text{ number}} = \sum_n C_{\mathcal{O},n} |n\rangle$$

$$\begin{aligned} G_{\mathcal{O}\mathcal{O}}^2(x) &= \langle 0 | e^{i\hat{p}^\mu x_\mu} \mathcal{O}^\dagger e^{-i\hat{p}^\mu x_\mu} \mathcal{O} | 0 \rangle \\ &= \sum_{n,m} C_{\mathcal{O},m}^* C_{\mathcal{O},n} \langle m | e^{-i\hat{p}^\mu x_\mu} |n\rangle e^{-i\hat{p}_n^\mu x_\mu} \langle m | n \rangle = \sum_n e^{-i\hat{p}_n^\mu x_\mu} C_{\mathcal{O},n}^* C_{\mathcal{O},n} \end{aligned}$$

If  $\langle \phi | \phi \rangle = C_0 \neq 0$ , replace  $\mathcal{O} \rightarrow \mathcal{O} - C_0$   
 (vacuum  $\rightarrow$  vacuum is boring)

Then  $P_n^0 > 0$

Note:  $1 = \int_{M^2}^{\infty} \frac{ds}{2\pi} \int \frac{d^4 p}{(2\pi)^4} (2\pi) \delta(p^2 - s) \Theta(p^0) (2\pi)^4 \delta^4(p - p_n)$   
 trivially

$$G_{\phi\phi}^{\vec{0}}(x) = \int_{M^2}^{\infty} \frac{ds}{2\pi} \int \frac{d^4 p}{(2\pi)^4} e^{-ip \cdot x} \Theta(p^0) 2\pi \delta(p^2 - s) \times \left[ \sum_n (2\pi)^4 \delta^4(p - p_n) C_n^* C_n \right]$$

this is unchanged by boosts: value for  $p_n^\mu =$  value for  $\Lambda_r^\mu{}_\nu p_n^\nu$   
 depends only on  $P_n^\mu$  through invariant  $P_n^2$ .

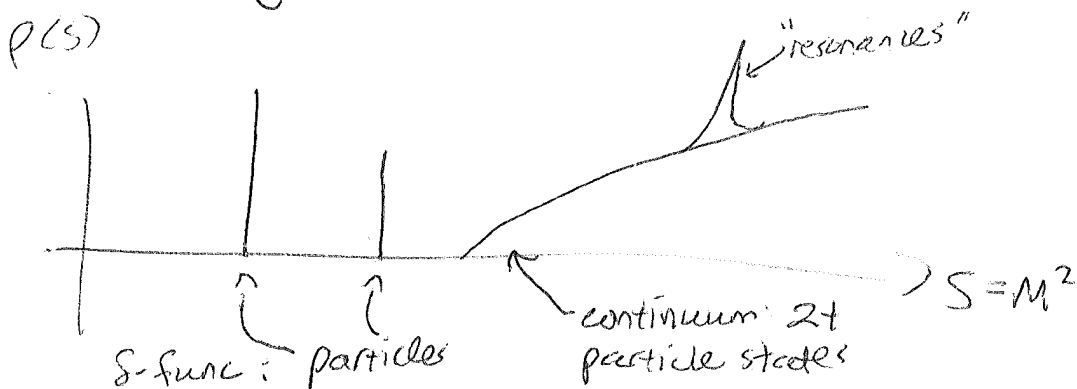
Name it  $\rho(P_n^2) = \rho(s)$

$$G_{\phi\phi}^{\vec{0}}(x) = \int_{M^2}^{\infty} \frac{ds}{2\pi} \rho(s) \Delta(x; s) \rightarrow \int \frac{d^4 p}{(2\pi)^4} (2\pi) \delta(p^2 - s) \Theta(p^0) e^{-ip \cdot x}$$

Behavior we saw for a massive free particle.

Integral over  $M^2$  values, of

$\rho(s)$   $\times$  what a massive particle does.



What else is in correlation functions?

L6P3

$$G^{\leftarrow}(x) = \langle 0 | \varphi(\omega) \varphi(x) | 0 \rangle = \text{same as } G^{\rightarrow} \text{ but } \textcircled{A} (-p^0)$$

$$D(x) = G^{\rightarrow}(x) - G^{\leftarrow}(x) = \langle 0 | [\varphi(x), \varphi(\omega)] | 0 \rangle$$

Cares about causality & teaches about propagation (how?)

First: for  $x^2 < 0$ ,  $[\varphi(x), \varphi(\omega)] = 0$  by causality. Also explicitly

~~E~~

For instance:  $\mathcal{L} \rightarrow \mathcal{L}_{\text{old}} - \varphi(x) J(x)$

$J(x)$  "external" (I-choose)  $x^\mu$ -dependent function. "Handle" to "wiggle" field.

With this  $\mathcal{L}$ , what is  $\langle 0, t=-\infty | \varphi(y) | 0, t=-\infty \rangle$ ?

Insert time evolution  $\langle 0 | U^\dagger(y^0, -\infty) \varphi(\vec{y}) U(y^0, -\infty) | 0 \rangle$

where  $U$  obeys  $HU(t-t_0) = i \mathcal{H} U(t-t_0)$  is time-evolution operator.

$$\mathcal{L} = \mathcal{L}_0 + \mathcal{L}_1 \rightarrow -\varphi(x) J(x) \quad \text{expand to 1-order in } \mathcal{H}_1$$

$$\mathcal{H} = \mathcal{H}_0 + \mathcal{H}_1 \rightarrow +\varphi(x) J(x)$$

$$U(t_1, -\infty) = 1 - i \int_{-\infty}^{t_1} \varphi(x) J(x) dt' \quad \text{in int. picture (Time-Dependent Pert. Thy)}$$

$$\begin{aligned} \langle 0 | \varphi(y) | 0 \rangle &= \langle 0 | \left( 1 + i \int_{x^0 < t}^{x^0 < t} \varphi(x) J(x) \right) \varphi(y) \left( 1 - i \int_{z^0 < t}^{z^0 < t} \varphi(z) J(z) \right) | 0 \rangle \\ &= -i \int_{x^0 < t}^{x^0 < t} \langle 0 | [\varphi(y), \varphi(x)] | 0 \rangle J(x) \\ &= -i \int_{x^0 < t}^{x^0 < t} J(x) \sigma(y-x) \end{aligned}$$

$\sigma(y-x)$  tells how cause at  $x$  affects field at  $y$ .

Technically the thing we wanted there was

$$G_R(y-x) = \langle 0 | [\phi(y), \phi(x)] | 0 \rangle \Theta(y^0 - x^0)$$

$$\langle \phi(y) | 0 \rangle_J = \int d^4x \quad -iJ(x) G_R(y-x) \quad \text{Retarded Function}$$

usually

~~In spacetime~~

usually we absorb the  $i$  factor into definition of  $G_R$ :

$$G_R(y-x) = -i \langle 0 | [\phi(y), \phi(x)] | 0 \rangle \Theta(y^0 - x^0)$$

Note: I can define  $G_R(p)$  at positive-imaginary-part-

frequency:  $G_R(p) = \int d^4y e^{iP_\mu y^\mu} \langle 0 | [\phi(y), \phi(0)] | 0 \rangle \Theta(y^0)$

for  $p^0 = p_r^0 + i p_i^0$ , that's  $\int d^4y e^{-p_r^0 y^0} e^{i y^\mu k_\mu}$   
 ↑  
 decaying factor. Only evaluate at  $y^0 > 0$ !

$G_R(p)$  is defined & free of singularities in upper-half  $p^0$  plane.

Let's compute it in our free theory!

We saw that  $\langle 0 | \phi(y) \phi(0) | 0 \rangle = \int \frac{d^4q}{(2\pi)^4} e^{-iP_\mu y^\mu} \Theta(q^0) 2\pi \delta(q^2 - m^2)$   
 $\langle 0 | \phi(0) \phi(y) | 0 \rangle = \int \frac{d^4q}{(2\pi)^4} e^{iP_\mu y^\mu} \Theta(-q^0) 2\pi \delta(q^2 - m^2)$

$$G_R(p) = \int d^4y \Theta(y^0) e^{iP_\mu y^\mu} \left[ \int \frac{d^4q}{(2\pi)^4} -i 2\pi \delta(q^2 - m^2) (\Theta(q^0) - \Theta(-q^0)) \right] e^{-i q_\mu y^\mu}$$

the  $\int dy^3 e^{-i(\vec{p}-\vec{q})\cdot\vec{y}} \rightarrow (2\pi)^3 \delta^3(\vec{p}-\vec{q})$  | L6P5

performs the  $\int \frac{d^3 q}{(2\pi)^3}$  part, leaving

$$G_R(p) = \int dy^0 \Theta(y^0) \int \frac{dq^0}{2\pi} (-2\pi i) e^{i(p^0 - q^0)y^0} (\Theta(q^0) - \Theta(-q^0)) \delta(q_0^2 - \omega_p^2)$$

How do I do that?

1) do  $\int dq^0$  using the  $\delta$ -function.  $q^0 = \omega_p$  with +  
 $q^0 = -\omega_p$  with -

2) again take  $p^0 \rightarrow p^0 + i\frac{\epsilon}{2}$  we will see why  
 so I can do  $y$ -int.

$y$ -int: 
$$\frac{-i}{2\omega_p} \left( e^{i(p^0 - \omega_p)y^0} - e^{i(p^0 + \omega_p)y^0} \right)$$

Recall  $\int_0^\infty dy e^{-ny} = \frac{1}{n}$ .  $\int dy e^{-(\epsilon - i\omega)y} = \frac{1}{\epsilon - i\omega}$ . so

$$\begin{aligned} \left(\frac{-i}{2\omega_p}\right) \int_0^\infty dy^0 \left( e^{i(p^0 - \omega_p)y^0} - e^{-(p^0 + \omega_p)y^0} \right) e^{-\epsilon/2 y^0} &= \left(\frac{-i}{2\omega_p}\right) \left( \frac{1}{\frac{\epsilon}{2} - i(p^0 - \omega_p)} - \frac{1}{\frac{\epsilon}{2} - i(p^0 + \omega_p)} \right) \\ &= \frac{1}{2\omega_p} \frac{2\omega_p}{(p^0 - \omega_p + i\frac{\epsilon}{2})(p^0 + \omega_p + i\frac{\epsilon}{2})} = \frac{1}{p_0^2 - \omega_p^2 + i\epsilon p_0} = \frac{1}{p^2 - m^2 + i\epsilon p_0} \end{aligned}$$

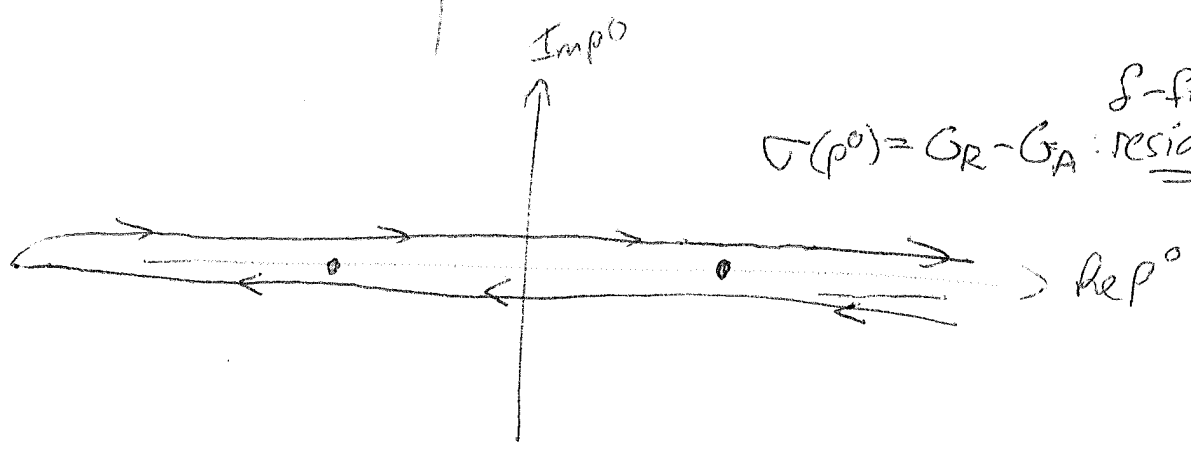
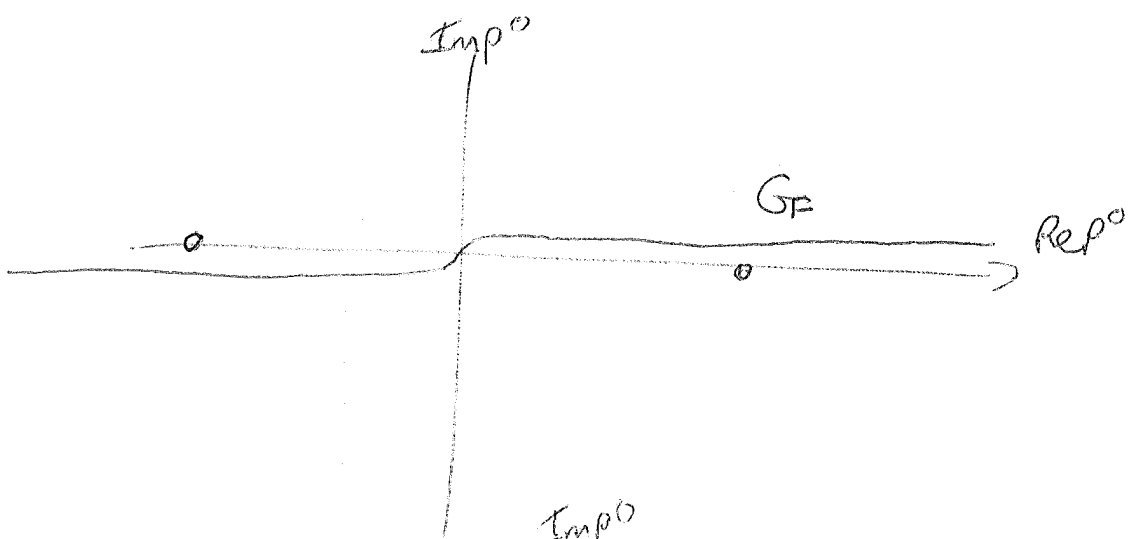
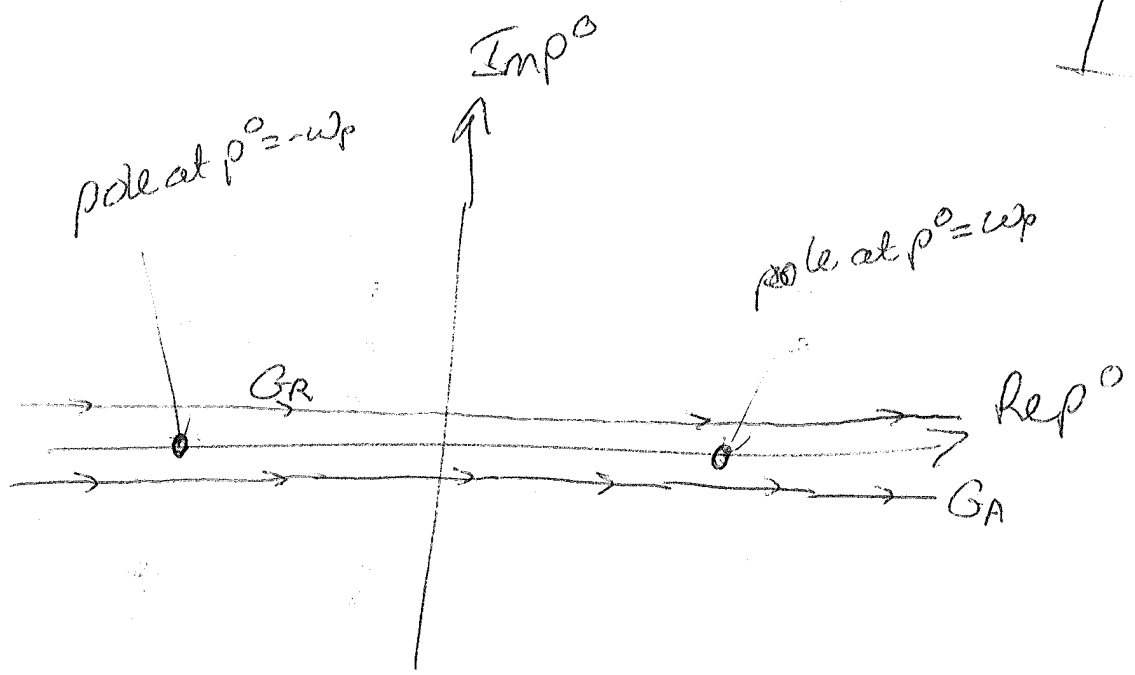
Some problems call for

$G_A(x) = \int \frac{d^4 p}{(2\pi)^4} \frac{e^{ipx}}{p^2 - m^2 - i\epsilon p_0} \rightarrow G_A(p) = \frac{1}{p^2 - m^2 - i\epsilon p_0}$  sign on  $\epsilon$  tells side of  $p^0 = \text{Real contour to use}$

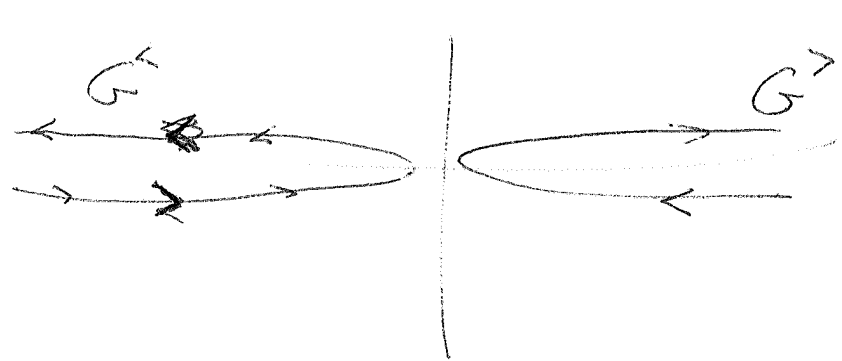
Some problems need

$G_F(x) = -i \langle 0 | ( \varphi(x) \varphi(0) \Theta(x^0) + \varphi(0) \varphi(x) \Theta(-x^0) ) | 0 \rangle \stackrel{?}{=} \langle 0 | T(\varphi(x) \varphi(0)) | 0 \rangle$

$G_F(p) = \frac{1}{p^2 - m^2 + i\epsilon}$  Book uses  $\Delta_F$



$\int$ -func. times  
 $\Gamma(p^0) = G_R - G_A$ : residues of poles



feels singularities but only on  $\frac{1}{2}$  the axis.